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1 The anisotropy of magnetic susceptibility of uniaxial
2 superparamagnetic particles: Consequences for its
3 interpretation in magnetite and maghemite bearing
4 rocks

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5 **Abstract.** A simple model that provides a quantitative description of the
6 magnetic susceptibility of superparamagnetic to stable single-domain uni-
7 axial magnetic particles can be built in the framework of the theory of stochas-
8 tic resonance. This model expands that of *Mullins and Tile* [1973] for super-
9 paramagnetic grains by considering the dependence of superparamagnetic
10 susceptibility on the particle orientation and thus describes the anisotropy
11 of magnetic susceptibility (AMS) of ensembles of superparamagnetic as well
12 as single-domain particles. The theory predicts that, on the contrary of sta-
13 ble single-domain, the maximum anisotropy of superparamagnetic particles
14 is parallel to their easy axis and shows that the AMS of ensembles of uni-
15 axial particle is strongly dependent on the distribution of particle grain-size,
16 coercivity, measurement temperature and frequency. It also explains why the
17 inverse AMS pattern expected for stable single-domain particles is rarely ob-
18 served in natural samples. We use examples of well-characterized obsidian
19 specimens to show that, as predicted by the theory, in the presence of sig-
20 nificant superparamagnetic contributions the maximum susceptibility axis
21 of AMS is directed along the preferential direction of particles easy axis.

1. Introduction

22 Fine-grained magnetic particles are very common in nature and their anisotropy of
23 magnetic susceptibility (AMS) has been commonly used in a variety of environmental
24 and tectonic studies [e.g., *Rochette et al.*, 1992]. In these magnetic particles of nanomet-
25 ric size, the transition from stable single-domain to superparamagnetic state is marked,
26 among other effects, by a severalfold increase of magnetic susceptibility. This transition
27 occurs in a relatively narrow interval of temperature and volumes when the particle relax-
28 ation time becomes comparable to the measurement time or, if measurements are made in
29 alternating field, to about the half-period. Their presence can be quantified with suscep-
30 tibility measurements at different temperatures or frequencies, which are often employed
31 in environmental studies on sediments and soils. However, despite the interest in AMS
32 and in superparamagnetic grain, the AMS of superparamagnetic grain is not well studied.

33 Neglecting the effect of temperature, the orientation of magnetic moment in uniaxial
34 single-domain particles is determined by the local minima of the particle self-energy and
35 an induced magnetization, hence their susceptibility, results from the shift of such min-
36 ima in an applied field [*Stoner and Wohlfarth*, 1948]. When the probability of energy
37 barrier hopping caused by thermal fluctuations becomes significant, the susceptibility is
38 increased by a superparamagnetic term that adds to the stable single-domain suscepti-
39 bility. The superparamagnetic susceptibility of an ensemble of non-interacting particles
40 can be described as that of a paramagnetic gas only if the particles blocking energy is
41 negligible compared to thermal energy. A more complete model for an ensemble of par-
42 ticles with easy-axis parallel to the magnetizing field, was proposed by *Mullins and Tile*

43 [1973], based on *Néel* [1949] theory. This model explains phenomena occurring during the
44 superparamagnetic-stable single-domain transition such as the quadrature susceptibility
45 (i.e., the susceptibility due to the component of magnetization 90° out of phase from
46 the driving field) and frequency dependence. In the rock- and paleo-magnetic literature,
47 the latter was discussed in detail by *Worm* [1998] while *Shcherbakov and Fabian* [2005]
48 and *Egli* [2009] investigated inverse methods to compute magnetic grain-size distributions
49 using the frequency-dependent susceptibility measured at different temperatures.

50 Although the *Mullins and Tile* [1973] model is still the main reference within the rock-
51 and paleo-magnetic scientific community, a vast amount of work on AC susceptibility is
52 available in the physics literature. The theory of stochastic resonance has been applied
53 to the AC susceptibility to describe interwell hopping both in the case of uniaxial and
54 triaxial particles [e.g., *Coffey et al.*, 2001; *Raikher et al.*, 2003; *Kalmykov et al.*, 2005, and
55 references therein]. The effect of intrawell contribution was introduced by *Svedlindh et*
56 *al.* [1997] and a semi-analytical expressions for the in-phase and quadrature susceptibility
57 that include the effect of surface anisotropy and (weak) dipolar interactions in the limit
58 of small field was developed by *Vernay et al.* [2014]. Many of these models attempt
59 to solve the most general problem based on the theory of *Brown* [1963], considering
60 simultaneously both interwell and intrawell fluctuations over a wide range of controlling
61 parameters. This generally involves solving the Fokker-Plank equation with a periodically
62 varying potential and leads to complicated calculations that can be evaluated only using
63 a numerical approach. Moreover most calculations contemplate only the case of particles
64 with anisotropy axis parallel to the field direction.

65 This paper presents a model describing the superparamagnetic susceptibility (χ_{SP}) of
66 uniaxial particles from the point of view of the theory of stochastic resonance [e.g., *McNara*
67 *and Wiesenfeld*, 1989; *Gammaitoni et al.*, 1998]. The proposed model is simplified
68 by restricting to the case of low-field susceptibility measured at AC frequencies satisfying
69 the adiabatic assumption. Within these limitations, which comprise virtually all kind of
70 rock-magnetic measurements, it is possible to consider a straightforward, bi-state model
71 that captures an accurate representation of uniaxial magnetic particles and yield simple
72 analytical expressions. It is shown that the χ_{SP} derived from this model is equivalent
73 to that of *Mullins and Tile* [1973] for particles with easy axis parallel to the field, hence
74 it is supported by the experimental evidence available in the literature. The proposed
75 model, however, expands the previous one introducing the dependence of χ_{SP} on particle
76 orientation and combining the interwell (superparamagnetic) and the intrawell (ferrimag-
77 netic) susceptibility. We focus on this aspect in order to quantify the AMS contribution
78 of superparamagnetic and stable single-domain grains showing that superparamagnetic
79 susceptibility is very likely to dominate the AMS pattern in many natural rock samples.
80 Experimental measurements from obsidians are shown to support the theory and the
81 consequence on AMS measurements in rock-magnetism are discussed.

2. Theory

2.1. Stochastic Resonance of Bi-state Magnetic Particles

82 In ferromagnetic (*s.l.*) material the magnetic susceptibility χ is defined as $\chi = \frac{\partial M}{\partial H}$ at
83 $H = 0$ [e.g., *Bertotti*, 1998]. Let's consider the magnetic susceptibility χ_{SP} due to the
84 barrier hopping caused by thermal fluctuation in a uniaxial particle of volume v , whose
85 geometry is depicted in Fig. 1a, subject to an alternating field with intensity H and

86 angular frequency ω . In zero field, the minima of the particle potential energy E are
 87 symmetrical and separated by the potential barrier $E_b = K_u v$, where K_u is the anisotropy
 88 constant. Thermally-induced hopping between the potential wells occurs but in this con-
 89 dition the symmetry of the system enforces the average effect to vanish. In the presence
 90 of a periodic field H , the double-well potential E is tilted back and forth, thereby raising
 91 and lowering successively the potential barriers of the right and the left well, respectively,
 92 in an antisymmetric manner (Fig. 1b). The periodic forcing due to the alternating field
 93 is too weak to let the magnetic moment move periodically from one potential well into
 94 the other one, however it introduces an asymmetry in the system and lets the stochastic
 95 interwell hopping come into play. Statistical effects of the thermal switching becomes par-
 96 ticularly relevant when the average waiting time between two thermally-induced interwell
 97 transitions is comparable with the half-period of the alternating field, causing an increase
 98 of the interwell hopping frequency. This phenomenon is called stochastic resonance.

99 The theory presented in this paper assumes a small driving AC field H (ideally $H \rightarrow 0$
 100 for the initial susceptibility) and a field frequencies $\omega \ll f_0$ where f_0 is the atomic attempt
 101 frequency, with $f_0 \approx 1$ GHz when computed from Néel's relaxation times [*Moskowitz*
 102 *et al.*, 1997]. These assumptions are fulfilled by rock-magnetic measurements at room-
 103 temperature and low-temperature. The discrete two-states model implies that the dis-
 104 tribution of the moment orientation is sharply peaked at the minima of the potential
 105 energy, which is a reasonable assumption for $\frac{K_u v}{k_B T} \geq 5$, hence for magnetic particles with a
 106 spherical equivalent diameter larger than a few nanometers [e.g., *García-Palacios*, 2000].
 107 In extremely small particles, however, quantum fluctuations become relevant and set a
 108 more stringent limit to the validity of models based on classical mechanics. Although this

109 limit is not precisely defined, it has been suggested [*Jones and Srivastava, 1989*] that
 110 a number of atoms $< 10^3$, which roughly corresponds to about 5 nm diameter, are the
 111 smallest particles that can be studied with classic models.

112 Within the above limits, this theory provides a useful model to calculate the average
 113 magnetization caused by thermally-induced interwell hopping of uniaxial particles subject
 114 to an alternating magnetic field, hence their AC superparamagnetic susceptibility.

2.2. Superparamagnetic Susceptibility

In the bi-state system considered above, the magnetic moment can be found in the
 states (potential minima) \pm with a probability (n_{\pm}) given by the master equation:

$$\frac{dn_+(t)}{dt} = -n_+(t) W_+ + n_-(t) W_-, \quad (1)$$

115 which is equivalent to that commonly used for deriving Néel relaxation time except that
 116 here the transition rate $W_{\pm}(t)$ out of the \pm state, is periodically modulated. The solution
 117 to this first-order differential equation (1) was given by *McNamara and Wiesenfeld* [1989]

$$\begin{aligned} n_+(t) &= g^{-1}(t) \left(n_+(t_0) g(t_0) + \int_{t_0}^t W_-(t') g(t') dt' \right) \\ g(t) &= \exp \left(\int^t (W_+(t') + W_-(t')) dt' \right) \end{aligned} \quad (2)$$

who proposed to use a periodically modulated escape rate W_{\pm} of the type

$$W_{\pm}(t) = f(\mu \pm \eta_0 \cos(\omega t)) \quad (3)$$

118 where μ is a dimensionless ratio between potential barrier and thermal noise of the un-
 119 perturbed system, and η_0 is the amplitude of the periodical modulation.

In a uniaxial magnetic particle the escape rate function $f(t)$ is proportional to an
 exponential function [e.g., *Néel, 1949*], the energy barrier of the unperturbed particle is

$\mu = -K_u v/k_B T$ and periodical fluctuation $\eta_0 = -E_H/k_B T$ is given by the ratio between the Zeeman energy and the thermal noise. Following *McNamara and Wiesenfeld* [1989], eq. (3) can be expanded in a Taylor series for small $\eta_0 \cos(\omega t)$ and after substituting μ and η_0 we obtain,

$$W_{\pm}(t) = C e^{-\frac{K_u v}{k_B T}} \left(1 \mp \frac{E_h}{k_B T} \cos(\omega t) + \frac{1}{2} \left(\frac{E_h}{k_B T} \right)^2 \cos^2(\omega t) \mp \frac{1}{6} \left(\frac{E_h}{k_B T} \right)^3 \cos^3(\omega t) + \dots \right), \quad (4)$$

hence

$$W_+(t) + W_-(t) = 2C e^{-\frac{K_u v}{k_B T}} \left(1 + \frac{1}{2} \left(\frac{E_h}{k_B T} \right)^2 \cos^2(\omega t) + \dots \right), \quad (5)$$

where C is a proportionality factor taken such that $2C$ corresponds to the Néel pre-exponential factor f_0 , hence $2C e^{-\frac{K_u v}{k_B T}} = 1/\tau$ is the inverse of Néel's relaxation time.

The integral (1) can now be performed analytically to the first order in $\eta_0 = -E_H/k_B T$ [e.g., *McNamara and Wiesenfeld*, 1989; *Gammaitoni et al.*, 1998],

$$n_+(t|x_0, t_0) = \frac{1}{2} \left(e^{-\frac{1}{\tau}(t-t_0)} (\delta_{x_0} - 1 - \kappa(t_0)) + 1 + \kappa(t) \right) \quad (6)$$

where $\kappa(t) = 1/\tau \frac{E_h}{k_B T} \cos(\omega t - \Phi)/\sqrt{1/\tau^2 + \omega^2}$ and $\Phi = \arctan(\omega \tau)$. According to *McNamara and Wiesenfeld* [1989] the quantity $n_+(t|x_0, t_0)$ represent the probability that the magnetic moment in the state $+$ at time t given the initial state, and the Kronecker delta δ_{x_0} is 1 when the system initially in state $+$. The mean value $\langle n_+(t) \rangle$ is obtained by averaging over a sufficiently long time (ideally $t_0 \rightarrow -\infty$) so that the memory of the initial conditions gets lost obtaining,

$$\langle n_+(t) \rangle = \frac{E_h}{k_B T \sqrt{1 + \omega^2 \tau^2}}. \quad (7)$$

The average superparamagnetic magnetization of a particle can then be expressed as

$$M = \langle n_+ \rangle M_s \cos(\phi - \theta). \quad (8)$$

where $\phi - \theta$ is the angle between the direction of the time-dependent field H and the magnetic moment M_s . For uniaxial particles in the hypothesis of small field one can find [e.g., *Lanci*, 2010]

$$M = \langle n_+ \rangle M_s \cos \left(\phi - \frac{\mu_0 M_s H \sin(\phi)}{\mu_0 M_s H \cos(\phi) + 2K_u} \right). \quad (9)$$

Moreover, in small field H , and consequently small angle θ , the Zeeman energy can be reduced to the first term of its Taylor series expansion around $\theta = 0$ leading to $E_H = \mu_0 M_s v H (\cos(\phi) + \sin(\phi)\theta)$. Substituting in (7) one obtains the following expression for $\langle n_+ \rangle$

$$\langle n_+ \rangle = \frac{\mu_0 H M_s v \cos(\phi)}{k_B T \sqrt{1 + \omega^2 \tau^2} - \mu_0 H M_s v \sin(\phi)}. \quad (10)$$

The variation of $\langle n_+ \rangle$ as a function of the temperature and grain orientation ϕ is shown in Fig. 2. Intuitively, the rapid initial increase of $\langle n_+ \rangle$ is due to magnetic moment unblocking, while the subsequent $\propto 1/T$ decrease can be explained by the increasing number of random interwell jumps, which cause a stronger randomization of the system.

The superparamagnetic susceptibility χ_{SP} of a grain with orientation ϕ can be calculated from the equations (9) and (10)

$$\chi_{SP}(\phi) = \frac{\partial}{\partial H} \left[\frac{\mu_0 H M_s v \cos(\phi)}{k_B T \sqrt{1 + \omega^2 \tau^2} - \mu_0 H M_s v \sin(\phi)} M_s \cos \left(\phi - \frac{\mu_0 M_s H \sin(\phi)}{\mu_0 M_s H \cos(\phi) + 2K_u} \right) \right]. \quad (11)$$

For $H \rightarrow 0$ one obtains

$$\chi_{SP}(\phi) = \frac{\mu_0 M_s^2 v \cos^2(\phi)}{k_B T \sqrt{1 + \omega^2 \tau^2}}. \quad (12)$$

The in-phase χ'_{SP} e quadrature χ''_{SP} components of χ_{SP} can be obtained straightaway using the phase angle $\Phi = \arctan(\omega \tau)$

$$\chi'_{SP}(\phi) = \frac{\mu_0 M_s^2 v \cos^2(\phi)}{k_B T (1 + \omega^2 \tau^2)} \quad (13)$$

$$\chi''_{SP}(\phi) = \frac{\mu_0 M_s^2 v \cos^2(\phi) \tau \omega}{k_B T (1 + \omega^2 \tau^2)}. \quad (14)$$

Equations 13 and 14 generalize *Mullins and Tile* [1973] introducing the dependence on particle orientation ϕ . $\chi_{SP}(\phi)$ shows a dependence on $\cos^2(\phi)$ indicating that the susceptibility of grains with easy-axis orthogonal to the field direction is null and that the largest contribution to superparamagnetic susceptibility is given by grains with easy-axis parallel to the field direction.

The in-phase χ'_{SP} e quadrature χ''_{SP} superparamagnetic susceptibility can be reduced to the isotropic case of *Mullins and Tile* [1973] by averaging them over ϕ uniformly distributed on a sphere obtaining

$$\chi'_{SP} = \frac{\mu_0 M_s^2 v}{3 k_B T} \frac{1}{1 + \omega^2 \tau^2} \quad (15)$$

$$\chi''_{SP} = \frac{\mu_0 M_s^2 v}{3 k_B T} \frac{\omega \tau}{1 + \omega^2 \tau^2}. \quad (16)$$

where the two factors are separated to highlight the low-field approximation of the Curie law term and the stochastic term.

The derivation of eq. (13) and eq. (14) has been criticized by one of the reviewer (A. Newell), although he admits that the result is correct. For this reason we forced ourself to adhere pedantically the original theory developed by *McNamara and Wiesenfeld* [1989] and revised by *Gammaitoni et al.* [1998] in such a way that their derivation can be easily followed by the readers.

144 One further criticism concern the concept of stochastic resonance, in particular neglect-
 145 ing that the peak shown in Fig. 2 represent the effect of stochastic resonance. Here we
 146 answer quoting *Gammaitoni et al.* [1998] who, referring to equivalent of $\langle n_+ \rangle$ (their x)
 147 write: “. . . we note that the amplitude x first increases with increasing noise level, reaches
 148 a maximum, and then decreases again. This is the celebrated stochastic resonance effect.”

2.3. Stable Single-domain and Superparamagnetic Susceptibility

In our two-state model, with the distribution of the moment orientation is sharply peaked at the potential energy minima, the intrawell contribution to magnetic susceptibility consists of the ferromagnetic (*s.l.*) susceptibility χ_F due to the shift of the self-energy minima in the applied field [e.g., *O'Reilly*, 1984; *Lanci*, 2010]. In single uniaxial particles with orientation ϕ (Fig. 1a), the initial ferromagnetic susceptibility χ_F is described by [e.g., *Lanci*, 2010]

$$\chi_F(\phi) = \frac{\mu_0 M_s^2 \sin^2(\phi)}{2 K_u}. \quad (17)$$

Coupling together the superparamagnetic in-phase χ'_{SP} and the stable single-domain susceptibility χ_F , the interwell jumps and intrawell contribution in the physics literature [e.g., *Svedlindh et al.*, 1997], the (in-phase) magnetic susceptibility per unit of volume, as generally measured by K-bridge, for an ensemble of grains with orientation ϕ can be expressed as the sum of equations (12) and (17) i.e.:

$$\chi'(\phi) = \frac{\mu_0 M_s^2 v \cos^2(\phi)}{k_B T (1 + \omega^2 \tau^2)} + \frac{\mu_0 M_s^2 \sin^2(\phi)}{2 K_u}. \quad (18)$$

In the isotropic case of an ensemble of single-domain uniaxial grains with uniformly distributed orientation on a sphere one has

$$\chi' = \frac{\mu_0 M_s^2 v}{3 k_B T} \frac{1}{1 + \omega^2 \tau^2} + \frac{\mu_0 M_s^2}{3 K_u} \quad (19)$$

149 which is equivalent to the formulation of *Shcherbakov and Fabian* [2005] and the so-called
 150 Néel model of *Egli* [2009].

151 Eq. (18) shows clearly that the dependence of χ on $\cos^2(\phi)$ of the superparamagnetic
 152 state (first term) is orthogonal to the $\sin^2(\phi)$ dependence of χ in the stable single-domain
 153 state (second term). In an anisotropic assemblages the prevalence of either χ_{SP} or χ_F
 154 will result in a different direction of the AMS maximum axis and of the AMS ellipsoid
 155 shape, going from the inverse pattern of a stable single-domain to normal pattern pre-
 156 dicted for superparamagnetic grains. This is shown in Fig. 3 by plotting $\chi(\phi)$ for different
 157 grains with increasing $K_u v/k_b T$ ratios. In stable single-domain grains ($K_u v/k_b T > 18$ at
 158 the 100 Hz frequency) $\chi(\phi)$ is largest at $\phi = \pi/2$. On the other hand, $\phi = 0$ increases
 159 and soon became dominant upon rising $K_u v/k_b T$. The transition from inverse to normal
 160 AMS occurs over a narrow range of $K_u v/k_b T$ values corresponding to the onset of su-
 161 perparamagnetic effect. Due to their much higher susceptibility, even small amounts of
 162 superparamagnetic grains are likely to dominate the total susceptibility signal, becoming
 163 the main AMS carriers in samples where grain sizes are not strictly confined to the stable
 164 single-domain range.

3. Comparison with Experimental Data

165 Natural obsidian samples taken from different localities (Lipari Is., Palmarola Is. and
 166 Sardinia) and flows, have been used to test the normal AMS pattern of superparamagnetic
 167 magnetite particles predicted by the theory. Volcanic glasses are a well-suited testing ma-
 168 terial, since they contain very fine-grained iron oxides. Furthermore, it is possible to select
 169 samples with negligible contributions from non-SD particles. Obsidian samples are often
 170 very anisotropic, due to the alignment of ferrimagnetic inclusions along the flow direction

171 [*Canón-Tapia and Castro, 2004*]. Because of the dominant magnetite mineralogy, and the
172 abovementioned properties, obsidians can be used to test if the inverse AMS pattern of
173 the stable single-domain is dominated by the normal AMS pattern of superparamagnetic
174 particles.

175 Obsidian samples have been selected on the basis of mineralogy and grain size consider-
176 ations derived from standard rock-magnetic measurements. The acquisition of isothermal
177 remanent magnetization (IRM) at room ($\sim 300\text{K}$) and liquid nitrogen (77K) temperature
178 was used to retrieve the contribution of superparamagnetic particles and investigate the
179 magnetic mineralogy. The IRM was acquired with a pulse magnetizer and measured mea-
180 sured with a 2G DC-SQUID cryogenic magnetometer. Comparison of measurements at
181 77K and 300K (Fig. 4) shows that all selected obsidian samples have a large superpara-
182 magnetic contribution with a ratio $\text{IRM}_{77\text{K}}$ to $\text{IRM}_{300\text{K}}$ of ~ 2 . The IRM acquisition
183 for both low- and room-temperature curves is compatible with a predominant magnetite
184 mineralogy, while the fraction of remanent magnetization acquired at field higher than
185 300 mT could be tentatively explained with strong magnetostriction or by partially oxi-
186 dized magnetite grains. Samples SB2 and Palmarola shows higher saturation field at 77K
187 that could count for the larger magnetocrystalline anisotropy of the monoclinic phase
188 below the Verwey transition temperature [*Abe et al., 1976*] or strong magnetostriction in
189 the smaller grains.

190 IRM results are supported by hysteresis loops (Fig. 5), which were measured with
191 Princeton Instrument vibrating sample magnetometer equipped with a cryostat for low
192 temperature measurements at 80K . Low-temperature loops have thicker hysteresis loops
193 and higher remanences compared to room temperature, as expected from theoretical mod-

194 els [*Lanci and Kent, 2003*], confirming presence of a large superparamagnetic fraction. The
195 increased coercivity of samples SB2 and Palmarola, seen with IRM_{77K} acquisition curves,
196 is also visible in the hysteresis loop measured at 80K, which is not saturated in the 0.7 T
197 maximum measurement field. However, the hysteresis loops do not shows the constricted
198 shape characteristic of a mixture of minerals with distinct (bi-modal) coercivity spectra,
199 such as magnetite and hematite, suggesting a monodispersed coercivity spectrum end
200 corroborating the hypothesis of monoclinic phase or strong magnetostriction of the SP
201 grains.

202 The absence of a significant fraction of magnetization carried by multi-domain grains
203 was verified by letting the samples cross the Verwey transition [*Verwey, 1939*]. The
204 switch between cubic and monocline lattice remove the remanence carried by magneto-
205 crystalline anisotropy, hence carried by multi-domain grains as well as equidimensional
206 single domain particles [e.g., *Muxworthy and McClelland, 1999*]. This was performed by
207 cooling the specimens at 77K applying a saturating field of 2 T and let them warm up to
208 300K and, the opposite, saturating the samples at 300K and measuring them after cooling
209 at 77K. The presence of the Verwey transition was observed in other obsidians samples
210 from the same flows that had a significant contribution of multi-domain grains and were,
211 therefore, rejected for the purpose of this study. In the selected samples instead, both
212 up-temperature and down-temperature measurements gave very similar magnetization
213 slightly lower than the room temperature measurements. Results are shown in Fig. 6 and
214 compared with the remanences at 300K and 77K, summarizing the negligible contribution
215 of multi domain and large contribution of superparamagnetic grains that characterize these
216 samples.

AMS measurements were performed using a KLY-3 Kappa Bridge and the 15 positions protocol, while the anisotropy of isothermal remanent magnetization (AIRM) was measured, on the same specimens, with a JR6 spinner magnetometer using a 12 positions protocol. The AIRM remanence was acquired applying a magnetic field of 20 mT to the samples, which were AF demagnetized before the next IRM along a different direction. The relatively low field was used because experimental studies have demonstrated the equivalence of anisotropy of thermal remanence with the low-field AIRM [Stephenson *et al.*, 1986], which became a standard procedure in rock magnetism. However, limited to the Lipari obsidians, we have tested the correspondence of AIRM acquired at 20 mT and 100 mT fields, which have virtually identical directions.

The directions of AMS and AIRM eigenvectors and the Flinn [2001] anisotropy parameters are plotted in Fig. 7. There are no practical differences between the direction of the principal axes of AMS and AIRM directions, indicating that all samples have a normal AMS pattern with the maximum susceptibility aligned with the preferential direction of the particle's easy axis indicated by the AIRM. The larger differences in the direction of the maximum anisotropy axes (about 20°) are observed in the SB2 and Palmarola specimens. The Flinn diagram shows similar degrees on anisotropy and similar shapes for AIRM and AMS. The AMS is better clustered and slightly less anisotropic than AIRM. This is a common experimental result [e.g. Stephenson *et al.*, 1986] that can be explained by the fact that AMS combines the inverse contribution of the stable single-domain grains with the predominant normal AMS of superparamagnetic grains.

4. Conclusions

238 We have described a simple model of magnetic susceptibility for uniaxial superparamag-
239 netic and stable single-domain particles based on the theory of stochastic resonance. This
240 model emphasizes the dependence of the susceptibility on the particle's orientation and
241 in particular it shows that stable single-domain and superparamagnetic particles possess
242 orthogonal maximum susceptibility axes. This means that in an ensemble of mixed stable
243 single-domain and superparamagnetic particles with a preferential orientation, the AMS
244 pattern can drastically change as function of grain size distribution, anisotropy constant
245 or even measurement frequency and temperature, ranging from an oblate inverse pattern
246 with the minimum eigenvalue along the field direction, which is characteristic of the sta-
247 ble single-domain state [e.g., *Rochette et al.*, 1992], to a prolate pattern with maximum
248 eigenvalue along the field direction predicted for superparamagnetic.

249 Because of this complex behavior a quantitative interpretation of the AMS pattern in
250 uniaxial magnetite/maghemite bearing rock seems rather complicated. In ensembles of
251 identical particles, there is sharp temperature dependence of the AMS pattern that is
252 related to the switch from stable single-domain to superparamagnetic, however in natural
253 samples with a wider distribution of $K_u v / k_b T$ ratios the transition can be more gradual.
254 In principle, this could be computed from (18) if the grain-size and coercivity distribu-
255 tions were accurately known, but this is unlikely in natural samples. Even if a complete
256 inversion of the AMS pattern does not occur because, for instance, the contribution of
257 superparamagnetic grains is not large enough, the strong dependence of AMS from the
258 $K_u v / k_b T$ ratio will introduce a bias in the AMS eigenvalues complicating their inter-

259 pretation. It is suggested that AMS measurements at different frequencies could help
260 recognizing the effect of superparamagnetic grains on AMS pattern.

261 Theoretical predictions are confirmed by results from obsidians samples, which have a
262 large superparamagnetic and negligible multi-domain grains population, and shows that
263 AMS axes are consistent with the AIRM axes, hence maximum anisotropy axes are align
264 to the easy axes. Other similar examples can be found in the literature *Canón-Tapia*
265 *and Castro* [2004]; *Canón-Tapia and Caródenas* [2012] for instance, have reported cases
266 of obsidians where the magnetic mineralogy was identified as a mixture of single-domain
267 magnetite with a substantial contribution of the superparamagnetic fraction and none of
268 them shows a inverse AMS pattern.

269 Our theory give an alternative explanation to the common case of coinciding AMS
270 and AIRM axes, which are usually interpreted as due to the presence of multi-domain
271 grains dominating the AMS [e.g., *Tarling and Hrouda*, 1993] and justify why the inverse
272 AMS is very rarely, if ever, observed in natural samples. In fact, inverse AMS is actually
273 restricted to the true stable single-domain state having a narrow range of grain sizes in
274 magnetite and maghemite. In natural samples stable single-domain particles are most
275 often combined with superparamagnetic and/or multi-domain particles, which are likely
276 to dominate the inverse AMS pattern either because of the much higher susceptibility of
277 the former or because larger volumes of the latter.

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Figure 1. (a) Geometrical description of the elements for uniaxial particles. (b) Sketch of the double-well potential $E = K_u v \sin^2 \phi$. In absence of periodic field H , the minima are located at a distance of π radiant and separated by a potential barrier with height $E_b = K_u v$. In the presence of periodic field H , the double-well potential is tilted back and forth raising and lowering the potential barriers of the right and the left well, respectively. In the figure the effect of the magnetic field on the potential E is exaggerated for clarity.

Figure 2. Amplitude of $\langle n_+(t) \rangle$ as function of the temperature for different orientation orientations ϕ (in radians) of the easy-axis. Peak-shaped function results from the effect of stochastic resonance. The stochastic resonance effect is maximum for grain with easy-axis along the field direction and null for grain with easy-axis orthogonal to the field direction, hence no superparamagnetic susceptibility is expected for the latter.

Figure 3. (a) Susceptibility $\chi(\phi)$ (in logarithmic scale) as function of the easy axis orientation ϕ . Lines of different colors represent grains with increasing $K_u v/k_b T$ ratios ranging approximately from 15 to 25, from superparamagnetic to stable single-domain. Stable single-domain grains dominated grains are characterized by maximum $\chi(\phi)$ at $\phi = \pi/2$, hence showing the characteristic inverse AMS pattern. On the contrary, at smaller $K_u v/k_b T$ ratio, the the susceptibility became much larger at $\phi = 0$ and exhibit the normal AMS pattern expected when superparamagnetism is dominant. (b) Susceptibility χ averaged over uniformly distributed ϕ as a function of the $K_u v/k_b T$ ratio. Black circles correspond to the same set of instances shown in panel (a). Other parameters used in the plot are $M_s = 480000$ A/m, and frequency $2\pi\omega = 100$ Hz.

Figure 4. IRM acquisition of obsidian samples at 300K (closed symbols) and 77K (open symbols). Palmarola and SB2 specimens show an increased coercivity at low temperature suggesting a higher degree of oxidation in superparamagnetic grains.

Figure 5. Hysteresis loops of obsidian samples. Thin blue line represent measurements at 80K and red thicker line represent room temperature measurements.

Figure 6. Low temperature (77K), room temperature (300K), up-temperature and down-temperature Verwey transition of obsidian samples. Differences between different measurements estimates the superparamagnetic, stable single-domain and multi domain contribution as described in the text.

Figure 7. Pattern of principal axes of AMS and AIRM in the obsidian samples (a) Flinn diagram [Flinn, 2001] indicating a generally try-axial shape of the anisotropy ellipsoids with similar values for AMS and AIRM. (b) Equal-area plot (lower hemisphere) of the directions of the principal anisotropy axes.