# Resummed lattice QCD equation of state at finite baryon density: Strangeness neutrality and beyond 

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#### Abstract

We calculate a resummed equation of state with lattice QCD simulations at imaginary chemical potentials. This work presents a generalization of the scheme introduced in [Phys. Rev. Lett. 126, 232001 (2021) to the case of nonzero $\mu_{S}$, focusing on the line of strangeness neutrality. We present results up to $\mu_{B} / T \leq 3.5$ on the strangeness neutral line $\langle S\rangle=0$ in the temperature range $130 \mathrm{MeV} \leq T \leq 280 \mathrm{MeV}$. We also extrapolate the finite baryon density equation of state to small nonzero values of the strangeness-tobaryon ratio $R=\langle S\rangle /\langle B\rangle$. We perform a continuum extrapolation using lattice simulations of the 4stoutimproved staggered action with $8,10,12$ and 16 time slices.


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## I. INTRODUCTION

The properties of hot and dense strongly interacting matter are the subject of many theoretical and experimental studies. At zero baryon density, first principle lattice QCD simulations showed that the transition between deconfined and confined matter is a crossover [1]. It is conjectured that the crossover line in the temperature $(T)$-baryochemical

[^0]potential $\left(\mu_{B}\right)$ plane eventually turns into a line of first order transition at a critical endpoint [2-7].

At finite baryon density, lattice simulations face the infamous complex action or sign problem [8,9]. Most results on finite density QCD therefore come from indirect methods, such as reweighting from zero chemical potential [10-16], Taylor expansion around zero chemical potential [17-28], or analytic continuation from imaginary chemical potentials, where the sign problem is absent [29-44]. More direct simulation approaches at finite density include certain reweighting techniques $[16,45,46]$, the complex Langevin equation [47-49] and methods based on Lefschetz thimbles [50-52]. Results with these more direct methods for full QCD are for the moment nonexistent (for

Lefschetz thimbles) or sparse (for complex Langevin and reweighting).

Another approach to locate the critical endpoint is the analysis of data from heavy ion collision experiments. An important tool for heavy ion phenomenology is the use of relativistic hydrodynamic simulations [53-55]. These in turn need the equation of state as a theoretical input in the whole range of temperatures and densities covered by the experiment.

The equation of state at zero density (or baryochemical potential) is known in the continuum limit from the crossover region [56-58] up to very high temperatures [59] where it can be matched with results from resummed perturbation theory [60-62].

The most straightforward way to extend the equation of state to finite chemical potentials is the use of a Taylor expansion in the chemical potential. For the pressure this reads:

$$
\begin{equation*}
\frac{p}{T^{4}}\left(T, \hat{\mu}_{B}, \hat{\mu}_{S}, \hat{\mu}_{Q}\right)=\sum_{i j k} \frac{1}{i!j!k!} \chi_{i j k}^{B S Q}(T) \hat{\mu}_{B}^{i} \hat{\mu}_{\mu^{j}}^{j} \hat{\mu}_{Q}^{k}, \tag{1}
\end{equation*}
$$

where the generalized susceptibilities are defined as

$$
\begin{equation*}
\chi_{i j k}^{B S Q}=\frac{\partial^{i+j+k}(\hat{p})}{\partial^{2} \hat{\mu}_{B} \partial^{2} \hat{\mu}_{S} \partial^{2} \hat{\mu}_{Q}}, \tag{2}
\end{equation*}
$$

and the dimensionless baryochemical potential is $\hat{\mu}_{B}=\frac{\mu_{B}}{T}$, and similarly for the strangeness chemical potential $\mu_{S}$ and the electric charge chemical potential $\mu_{Q}$. The dimensionless pressure is $\hat{p}=\frac{p}{T^{4}}$. The expansion coefficients $\chi_{i j k}^{B S Q}$ can be calculated either by direct simulations at zero chemical potential, or by simulations at imaginary chemical potentials followed by a fit. The Taylor coefficients are known up to fourth order in the continuum limit [22-24]. At high temperatures, these coefficients match those calculated from resummed perturbation theory [63-65]. At finite lattice spacings, also the sixth and eighth derivatives have been calculated, albeit with modest precision [28,42].

Naive truncations of the Taylor expansion show undesirable properties above $\mu_{B} / T \gtrsim(2-2.5)$ : namely, unphysical oscillations in the equation of state. This is due to the sign structure of higher order coefficients starting from $\chi_{6}^{B}$, which is in turn caused by the $\hat{\mu}_{B}$ dependence of the crossover transition [4,66-69]. It appears that extrapolating through the crossover transition-as one is forced to do with Taylor expansions at a fixed temperature-is hard.

One can go beyond fixed order polynomials, by considering resummations of the Taylor series. The classic example of this is Padé resummation, which has been used also for finite density QCD [32,44,70,71]. While the convergence properties of Padé approximants are in general superior to ordinary Taylor expansions, it is hard to tell
which expansion will work better at the low orders of the expansion available.

A way to extend the chemical potential range accessible by indirect methods already at low orders of the approximation is the use of a physically motivated extrapolation scheme. Recently, in Ref. [69] the authors of this paper introduced exactly such a scheme. It was motivated by an important finding in recent studies of the crossover region at imaginary chemical potentials: the existence of an approximate scaling variable $T\left(1+\kappa \hat{\mu}_{B}^{2}\right)$, where $T$ is the temperature, and $\mu_{B}$ the baryochemical potential. As a function of this variable, many observables-such as the chiral condensate or the baryon number-to-chemical potential ratio $\chi_{1}^{B} / \hat{\mu}_{B}$-collapse to a single curve when measured at different fixed values of $\operatorname{Im} \mu_{B} / T$ as a function of temperature $[43,69]$. Up to $\mu_{B} / T=1.5$ the existence of an approximate scaling variable has also been confirmed directly at a real baryochemical potential, by reweighting from the sign quenched ensemble [46].

One might suspect that the existence of such an approximate scaling variable is ultimately related to criticality in the two-flavor chiral limit of QCD [72,73]. If the universal contribution to the equation of state is large, one expects a single scaling variable-which is a combination of the quark mass, the temperature and the chemical potential-to govern most thermodynamic observables. If in the low chemical potential region the strength of the transition is ultimately governed by the quark mass, one expects curves sensitive to criticality to not change shape significantly at small chemical potentials. This is also consistent with the observation that the width of the crossover transition is, to a good approximation, constant at small $\mu_{B}[35,43]$.

Based on these observations, one can define for an observable $F$ of interest (of sigmoid shape in the temperature, such as $\chi_{1}^{B} / \hat{\mu}_{B}$ ) the following ansatz:

$$
\begin{equation*}
F\left(T, \hat{\mu}_{B}\right)=F\left(T^{\prime}, 0\right), \tag{3}
\end{equation*}
$$

where the chemical potential dependence is absorbed in the rescaled temperature $T^{\prime}=T\left(1+\kappa_{2}^{F}(T) \hat{\mu}_{B}^{2}+\kappa_{4}^{F}(T) \hat{\mu}_{B}^{4}+\cdots\right)$. The superscripts on the $\kappa_{n}^{F}$ denote that the expansion coefficients are different for different observables. The approximate constant strength of the crossover transition at small chemical potentials is manifest in the approximate temperature independence of the coefficient $\kappa_{2}(T)$ in the crossover region. A strengthening or weakening of the transition at larger chemical potentials could in turn manifest itself in a nontrivial temperature dependence of the higher order expansion coefficients $\kappa_{n}$ in the crossover region.

While in our previous study in Ref. [69] we showed the QCD equation of state up to $\mu_{B} / T=3.5$ for the simplest case of $\mu_{S}=\mu_{Q}=0$, for the phenomenologically more relevant strangeness neutral line $\langle S\rangle=0$, the equation of state from first principle calculations is still only available
as a fixed order Taylor expansion at fixed $T$, and it therefore covers a rather limited range in the chemical potentials, due to the aforementioned unphysical oscillations.

Here, we extend our previous investigations using our novel resummation scheme, conducted at zero strangeness chemical potential $\mu_{S}=0$ in Ref. [69] to the strangeness neutral line, and present results up to $\mu_{B} / T=3.5$ in this setting. This equation of state can be used in hydrodynamic simulations where local strangeness neutrality is enforced.

We also introduce an additional component to the resummation scheme-the Stefan-Boltzmann correction discussed in the next section. This will improve the convergence of our ansatz at high temperatures, where the aforementioned approximate scaling does not hold.

Finally, we go beyond strangeness neutrality by calculating, at finite real $\mu_{B}$, the expansion coefficients needed to calculate the equation of state at small values of the strangeness-to-baryon number ratio:

$$
\begin{equation*}
R \equiv \frac{\langle S\rangle}{\langle B\rangle}=\frac{\chi_{1}^{S}}{\chi_{1}^{B}} . \tag{4}
\end{equation*}
$$

This allows one to relax the condition of local strangeness neutrality in hydrodynamic simulations. This is necessary, as in a heavy ion collision, only global strangeness neutrality is guaranteed, and local charge fluctuations can be large.

The structure of the paper is as follows. In Sec. II we briefly describe the strangeness neutrality condition. In Sec. III. we formally define our extrapolation ansatz on the strangeness neutral line, including the Stefan-Boltzmann correction. In Sec. IV we describe our method to determine the extrapolation coefficients. In Sec. V we use the coefficients determined in Sec. IV to calculate thermodynamics at finite real chemical potentials both on the strangeness neutral line and in its vicinity.

## II. STRANGENESS NEUTRALITY

When enforcing strangeness neutrality, two sets of conditions are widely used in the literature. In the first, one sets $\mu_{Q}=0$ and defines a curve in the $\mu_{B}-\mu_{S}$ plane by the condition $\chi_{1}^{S}=0$. By differentiation, this implies:

$$
\begin{equation*}
\frac{\mathrm{d} \mu_{S}}{\mathrm{~d} \mu_{B}}=-\frac{\chi_{11}^{B S}}{\chi_{2}^{S}} \tag{5}
\end{equation*}
$$

On this line, total derivatives with respect to the baryochemical potential read

$$
\begin{equation*}
\frac{\mathrm{d}}{\mathrm{~d} \hat{\mu}_{B}}=\frac{\partial}{\partial \hat{\mu}_{B}}+\frac{\mathrm{d} \hat{\mu}_{S}}{\mathrm{~d} \hat{\mu}_{B}} \frac{\partial}{\partial \hat{\mu}_{S}}=\frac{\partial}{\partial \hat{\mu}_{B}}-\frac{\chi_{11}^{B S}}{\chi_{2}^{S}} \frac{\partial}{\partial \hat{\mu}_{S}} . \tag{6}
\end{equation*}
$$

We denote the total derivatives of the dimensionless pressure with respect to the baryochemical potential along the line $\mu_{Q}=0$ and $\chi_{1}^{S}=0$ as:

$$
\begin{equation*}
\left.c_{n}^{B}\left(T, \hat{\mu}_{B}\right) \equiv \frac{\mathrm{d}^{n} \hat{p}\left(T, \hat{\mu}_{B}\right)}{\mathrm{d} \hat{\mu}_{B}^{n}}\right|_{\substack{\mu_{0}=0 \\ x_{1}^{S}=0}} . \tag{7}
\end{equation*}
$$

The $c_{n}^{B}\left(T, \hat{\mu}_{B}\right)$ essentially are the expansion coefficients of a Taylor expansion carried out along the line with $\mu_{Q}=0$ and $\chi_{1}^{S}=0$, where the leading one

$$
\begin{equation*}
c_{1}^{B}\left(T, \hat{\mu}_{B}\right)=\chi_{1}^{B}-\frac{\chi_{11}^{B S}}{\chi_{2}^{S}} \chi_{1}^{S}=\chi_{1}^{B} \tag{8}
\end{equation*}
$$

reduces to the net baryon density, while the second one reads:

$$
\begin{equation*}
c_{2}^{B}\left(T, \hat{\mu}_{B}\right)=\chi_{2}^{B}-\frac{\left(\chi_{11}^{B S}\right)^{2}}{\chi_{2}^{S}} . \tag{9}
\end{equation*}
$$

Since a vanishing electric charge chemical potential $\mu_{Q}=0$ is equivalent to zero isospin, this condition is equivalent to $\chi_{1}^{Q}=0.5 \chi_{1}^{B}$. While this is a much better approximation for the conditions of a heavy ion collision than simply taking $\mu_{S}=\mu_{Q}=0$, there is a further correction which can be taken into account: the slight isospin imbalance of the colliding nuclei (typically lead or gold). This amounts to the two constraints $\chi_{1}^{S}=0$ and $0.4 \chi_{1}^{B}=\chi_{1}^{Q}$, which define a curve in the $\mu_{B}-\mu_{S}-\mu_{Q}$ space. Along this curve the total derivatives are

$$
\begin{equation*}
\frac{\mathrm{d}}{\mathrm{~d} \hat{\mu}_{B}}=\frac{\partial}{\partial \hat{\mu}_{B}}+\frac{\mathrm{d} \hat{\mu}_{S}}{\mathrm{~d} \hat{\mu}_{B}} \frac{\partial}{\partial \hat{\mu}_{S}}+\frac{\mathrm{d} \hat{\mu}_{Q}}{\mathrm{~d} \hat{\mu}_{B}} \frac{\partial}{\partial \hat{\mu}_{Q}} \tag{10}
\end{equation*}
$$

where $\frac{\mathrm{d} \hat{\mu}_{S}}{\mathrm{~d} \hat{\mu}_{B}}$ and $\frac{\mathrm{d} \hat{\mu}_{Q}}{\mathrm{~d} \hat{\mu}_{B}}$ are given by the solution of the constraints:
$\chi_{11}^{B S}+\chi_{11}^{S Q} \frac{\mathrm{~d} \hat{\mu}_{Q}}{\mathrm{~d} \hat{\mu}_{B}}+\chi_{2}^{S} \frac{\mathrm{~d} \hat{\mu}_{S}}{\mathrm{~d} \hat{\mu}_{B}}=0$,
$\chi_{11}^{B Q}+\chi_{11}^{S Q} \frac{\mathrm{~d} \hat{\mu}_{S}}{\mathrm{~d} \hat{\mu}_{B}}+\chi_{2}^{Q} \frac{\mathrm{~d} \hat{\mu}_{Q}}{\mathrm{~d} \hat{\mu}_{B}}=0.4\left(\chi_{2}^{B}+\chi_{11}^{B S} \frac{\mathrm{~d} \hat{\mu}_{S}}{\mathrm{~d} \hat{\mu}_{B}}+\chi_{11}^{B Q} \frac{\mathrm{~d} \hat{\mu}_{Q}}{\mathrm{~d} \hat{\mu}_{B}}\right)$.

In this case, total derivatives will be denoted:

$$
\begin{equation*}
\left.d_{n}^{B}\left(T, \hat{\mu}_{B}\right) \equiv \frac{\mathrm{d}^{n} \hat{p}\left(T, \hat{\mu}_{B}\right)}{\mathrm{d} \hat{\mu}_{B}^{n}}\right|_{\substack{x_{1}^{Q}=0.4 X_{1}^{B} \\ x_{1}^{S}=0}} \tag{12}
\end{equation*}
$$

For simplicity, through most of this manuscript, we will use the first set of conditions with $\mu_{Q}=0$. In Sec. IV B, we will consider the difference between the two schemes in the leading order of the Taylor expansion-i.e., we will calculate $c_{2}^{B}(T, 0)$ and $d_{2}^{B}(T, 0)$ and their temperature derivatives, at which point we will also discuss the rationale for our choice of the first setting with $\mu_{Q}=0$.

## III. THE EXTRAPOLATION SCHEME

Before describing our extrapolation ansatz, we note that the ansatz given by Eq. (3), introduced in Ref. [69] for $\mu_{S}=0$, would also work at strangeness neutrality. The existence of the approximate scaling variable on the strangeness neutral line is shown in Fig. 1 for the quantity $c_{1}^{B} / \mu_{B}$, where on the left panel we show the data points of our simulations for a $48^{3} \times 12$ lattice, while on the right panel we show the same data points as a function of a rescaled temperature $T\left(1+\kappa \hat{\mu}^{2}\right)$. Notice that the collapse plot with a constant $\kappa$ does not work quite as well at high temperatures. Indeed, one does not expect an approximate scaling variable outside the crossover range. Our scheme can still incorporate this behavior by the temperature dependence of the $\kappa_{n}$ coefficients. In fact, with the ansatz given by Eq. (3) the coefficient $\kappa_{2}$ grows at high temperatures (see later).

One shortcoming of this scheme is that its region of applicability is restricted by the Stefan-Boltzmann limit of the right hand side of Eq. (3). When the quantity $F\left(T, \hat{\mu}_{B}\right)$ becomes larger than its infinite temperature limit at $\mu_{B}=0$, the ansatz in Eq. (3) must break down. It is easy to address this shortcoming, however, using the scheme only for observables $F$ that have an infinite temperature limit that is independent of $\hat{\mu}_{B}$. Given an observable that does not possess this property, one can easily construct another observable, simply dividing it by its own Stefan-Boltzmann limit:

$$
\begin{equation*}
F\left(T, \hat{\mu}_{B}\right) \rightarrow \frac{F\left(T, \hat{\mu}_{B}\right)}{\bar{F}\left(\hat{\mu}_{B}\right)} \tag{13}
\end{equation*}
$$

where the Stefan-Boltzmann limits [74] are denoted by

$$
\begin{equation*}
\bar{F}\left(\hat{\mu}_{B}\right)=\lim _{T \rightarrow \infty} F\left(T, \hat{\mu}_{B}\right) \tag{14}
\end{equation*}
$$

By construction, the ratio on the right hand side of Eq. (14) has an infinite temperature limit equal to one, at all values
of $\hat{\mu}_{B}$. By using the ansatz from Eq. (3) on this StefanBoltzmann corrected observable, we arrive at our new scheme, given by

$$
\begin{equation*}
\frac{F\left(T, \hat{\mu}_{B}\right)}{\bar{F}\left(\hat{\mu}_{B}\right)}=\frac{F\left(T_{F}^{\prime}, 0\right)}{\bar{F}(0)} \tag{15}
\end{equation*}
$$

where the temperature on the right-hand side is expanded as

$$
\begin{equation*}
T_{F}^{\prime}=T\left(1+\lambda_{2}^{F}(T) \hat{\mu}_{B}^{2}+\lambda_{4}^{F}(T) \hat{\mu}_{B}^{4}+\cdots\right) \tag{16}
\end{equation*}
$$

As shown in Fig. 2, this Stefan-Boltzmann correction does not spoil the collapse plot in the approximate scaling variable, meaning that the fast convergence of the scheme in the crossover region is maintained, with a $\lambda_{2}(T)$ coefficient that is approximately constant in the crossover range. The limitation at high temperature is however removed. Furthermore, as can be seen on the left panel of Fig. 2, the coefficients $\lambda_{n}$ must go to zero at high temperatures, as the data points for the different imaginary chemical potentials almost overlap. This is in contrast to the scheme of Eq. (3), where $\kappa_{2}$ grows at high temperatures.

For any finite order in the expansion in the $\lambda_{n}$, Eq. (15) generates an infinite number of terms in the Taylor expansion of the quantity $F$, thus the ansatz achieves a particular resummation of the Taylor expansion. As discussed in the introduction, this resummation is expected to converge fast in the crossover region, where the strength of the transition stays approximately constant.

In this work, we will consider three different observables $F$. First, we consider the normalized net baryon density $F=c_{1}^{B} / \hat{\mu}_{B}$. By noticing that $\lim _{\hat{\mu}_{B} \rightarrow 0} \frac{c_{1}^{B}\left(T, \hat{\mu}_{B}\right)}{\hat{\mu}_{B}}=c_{2}^{B}(T, 0)$, Eq. (15) becomes:

$$
\begin{equation*}
\frac{c_{1}^{B}\left(T, \hat{\mu}_{B}\right)}{\overline{c_{1}^{B}}\left(\hat{\mu}_{B}\right)}=\frac{c_{2}^{B}\left(T_{B B}^{\prime}, 0\right)}{\overline{c_{2}^{B}}(0)} \tag{17}
\end{equation*}
$$

where the infinite temperature limits of $c_{1}^{B}$ and $c_{2}^{B}$ are denoted $\overline{c_{1}^{B}}$ and $\overline{c_{2}^{B}}$ respectively, and


FIG. 1. Left: the scaled total derivative $c_{1}^{B} / \hat{\mu}_{B}$ on the strangeness neutral line from our imaginary chemical potential simulations. The data points at $\mu_{B}=0$ show the second derivative $\frac{\mathrm{d}^{2} \hat{p}}{\mathrm{~d} \hat{\mu}_{B}^{2}}$. Right: same observables, with the temperature rescaled by a factor $1+\kappa \hat{\mu}_{B}^{2}$.


FIG. 2. Left: the scaled total derivative $c_{1}^{B} / \hat{\mu}_{B}$ on the strangeness neutral line from our imaginary chemical potential simulations, divided by its chemical potential dependent Stefan-Boltzmann limit. The data points at $\mu_{B}=0$ show the second derivative $c_{2}^{B}$ divided by its Stefan-Boltzmann limit. Right: same observables, with the temperature rescaled by a factor $1+\lambda \hat{\mu}_{B}^{2}$.

$$
\begin{equation*}
T_{B B}^{\prime} \equiv T\left(1+\lambda_{2}^{B B}(T) \hat{\mu}_{B}^{2}+\lambda_{4}^{B B}(T) \hat{\mu}_{B}^{4}+\cdots\right) \tag{18}
\end{equation*}
$$

The Stefan-Boltzmann limits are easily obtained:

$$
\begin{align*}
\overline{c_{1}^{B}}\left(\hat{\mu}_{B}\right) & =\hat{\mu}_{B} \overline{c_{2}^{B}}(0)+\hat{\mu}_{B}^{3} \overline{c_{4}^{B}}(0), \\
\overline{c_{2}^{B}}(0) & =\frac{2}{9}, \quad \overline{c_{4}^{B}}(0)=\frac{4}{27 \pi^{2}} . \tag{19}
\end{align*}
$$

The $c_{2}^{B}(T, 0)$ function at $\mu_{B}=0$, together with the temper-ature-dependent coefficients in Eq. (18), are sufficient to extrapolate the strangeness-neutral equation of state to finite baryon density as we will show in Sec. V.

Second, we will consider the normalized strangeness chemical potential that is needed to realize the $\chi_{1}^{S} \equiv 0$ condition in a grand canonical ensemble: $F=\mathbf{M}\left(T, \hat{\mu}_{B}\right) \equiv$ $\frac{\hat{\mu}_{S}}{\hat{\mu}_{B}}\left(T, \hat{\mu}_{B}\right)$. Since

$$
\begin{equation*}
\lim _{\hat{\mu}_{B} \rightarrow 0} \mathrm{M}\left(T, \hat{\mu}_{B}\right)=-\frac{\chi_{11}^{B S}(T, 0)}{\chi_{2}^{S}(T, 0)} \equiv \mathrm{M}(T, 0) \tag{20}
\end{equation*}
$$

Eq. (15) becomes:

$$
\begin{equation*}
\frac{\mathbf{M}\left(T, \hat{\mu}_{B}\right)}{\overline{\mathbf{M}}\left(\hat{\mu}_{B}\right)}=\frac{\mathbf{M}\left(T_{B S}^{\prime}, 0\right)}{\overline{\mathbf{M}}(0)} \tag{21}
\end{equation*}
$$

with the Stefan Boltzmann limit $\overline{\mathbf{M}}\left(\hat{\mu}_{B}\right)=\lim _{T \rightarrow \infty} \mathbf{M}\left(T, \hat{\mu}_{B}\right)$ and

$$
\begin{equation*}
T_{B S}^{\prime}=T\left(1+\lambda_{2}^{B S}(T) \hat{\mu}_{B}^{2}+\lambda_{4}^{B S}(T) \hat{\mu}_{B}^{4}+\cdots\right) \tag{22}
\end{equation*}
$$

Finally, we will consider $F=\chi_{2}^{S}$, and denote its StefanBoltzmann limit by $\overline{\chi_{2}^{S}}$. For this observable, Eq. (15) reads:

$$
\begin{equation*}
\frac{\chi_{2}^{S}\left(T, \hat{\mu}_{B}\right)}{\overline{\chi_{2}^{S}}\left(\hat{\mu}_{B}\right)}=\frac{\chi_{2}^{S}\left(T_{S S}^{\prime}, 0\right)}{\overline{\chi_{2}^{S}}(0)} \tag{23}
\end{equation*}
$$

where

$$
\begin{equation*}
T_{S S}^{\prime}=T\left(1+\lambda_{2}^{S S}(T) \hat{\mu}_{B}^{2}+\lambda_{4}^{S S}(T) \hat{\mu}_{B}^{4}+\cdots\right) \tag{24}
\end{equation*}
$$

Note that, at strangeness neutrality, the Stefan-Boltzmann limits of M and $\chi_{2}^{S}$ are independent of $\hat{\mu}_{B}$,

$$
\begin{equation*}
\overline{\mathbf{M}}\left(\hat{\mu}_{B}\right)=\overline{\mathbf{M}}(0)=\frac{1}{3}, \quad \overline{\chi_{2}^{S}}\left(\hat{\mu}_{B}\right)=1 \tag{25}
\end{equation*}
$$



FIG. 3. The strangeness to baryon chemical potential ratio (upper panel) and the strangeness susceptibility (lower panel) at simulated imaginary baryochemical potentials on our $48^{3} \times 12$ ensembles.


FIG. 4. The $\kappa_{2} \equiv \kappa_{2}^{B B}$ coefficients obtained in the scheme of Eq. (3) without the Stefan-Boltzmann correction (in blue), compared to the coefficients $\lambda_{2} \equiv \lambda_{2}^{B B}$ obtained in the scheme of Eq. (15) with the Stefan-Boltzmann correction (in red) on our $48^{3} \times 12$ lattices.
thus $\kappa_{n}^{B S}=\lambda_{n}^{B S}$ and $\kappa_{n}^{S S}=\lambda_{n}^{S S}$. This concludes the definition of our scheme for all observables of interest. We show the lattice data for M and $\chi_{2}^{S}$ on our $48^{3} \times 12$ ensembles in Fig. 3.

Finally, in order to illustrate the advantageous properties of the Stefan-Boltzmann correction mentioned above, we compare the $\kappa_{2}^{B B}$ and $\lambda_{2}^{B B}$ coefficients obtained without and with the Stefan-Boltzmann correction respectively, in Fig. 4 for our $48^{3} \times 12$ lattices. One can clearly see that the diverging behavior of the $\kappa_{2}$ coefficient at large temperatures is not present in the $\lambda_{2}$ coefficient. Thanks to this, our scheme can be applied at large temperatures as well. Moreover, the milder temperature dependence of $\lambda_{2}$, and its smaller magnitude, extend that range in $\hat{\mu}_{B}$ where our scheme can be applied.

## IV. DETERMINATION OF THE EXPANSION COEFFICIENTS

## A. Lattice setup

We perform simulations with $N_{f}=2+1+1$ dynamical quark flavors, with physical light, strange and charm quark masses. We use staggered fermions with fat links constructed with 4 steps of stout smearing [75] with the smearing parameter $\rho=0.125$ and a tree-level Symanzikimproved gauge action. This discretization of the QCD action was first used in Ref. [22], where information about the line of constant physics can be found. For the scale setting we use either the pion decay constant $f_{\pi}=$ 130.41 MeV [76], or the Wilson flow based $w_{0}=$ 0.1725 fm scale introduced in Ref. [77]. Taking into account the difference between the two scale settings is part of our systematic error analysis. We use lattices of temporal extent $N_{\tau}=8,10,12$ ad 16 to perform a continuum limit. The spatial volume is given by the aspect
ratio of $L T=4$. We performed simulations for imaginary baryochemical potentials given by $\operatorname{Im} \hat{\mu}_{B} \frac{8}{\pi}=0,3,4,5,6$, 6.5. In addition, for the $N_{\tau}=12$ lattices we also have data at $\operatorname{Im} \hat{\mu}_{B} \frac{8}{\pi}=5.5$. Strangeness neutrality was enforced on our imaginary chemical potential ensembles via the procedure discussed in Ref. [43].

## B. $c_{2}^{B}, d_{2}^{B}$ and their $T$ derivative at $\mu_{B}=0$

In order to determine thermodynamics at finite real chemical potentials, the right hand side of Eq. (17), i.e., $c_{2}^{B}(T, 0)$ must be known. For some quantities, like the entropy, its temperature derivative is also needed. We describe the determination of this quantity and its derivative here.

Given the second order quark number susceptibilities at vanishing chemical potentials, one can express $c_{2}^{B}(T, 0)$ as in Eq. (9):

$$
\begin{equation*}
c_{2}^{B}(T, 0)=\chi_{2}^{B}-\frac{\left(\chi_{11}^{B S}\right)^{2}}{\chi_{2}^{S}} \tag{26}
\end{equation*}
$$

which we continuum extrapolate using our $N_{\tau}=10,12$ and 16 lattices (for $c_{2}^{B}(T)$ and $\left.\mathrm{M}(T)\right)$ or $N_{\tau}=12,16$ and 20 [for $\chi_{2}^{S}(T)$ ], using a tree level improvement for the observables, as detailed in Refs. [20,56]. For the calculation of the temperature derivative we use the same procedure as in Ref. [69]. A procedure which we also describe here, to make the manuscript self-contained.

The derivatives are obtained by fitting the data with suitable ansätze, which are then differentiated. We first divide the temperature range into two parts: the crossover region and the high temperature region. In the first, we interpolate $c_{2}^{B}$ with cubic basis splines with nodes points in the range $T \in[130 \mathrm{MeV}, 300 \mathrm{MeV}]$. We perform a joint fit of the splines in temperature and in $1 / N_{\tau}^{2}$. The ansatz reads:

$$
\begin{equation*}
c_{2}^{B}(T, 0)=\sum_{i=1}^{n} \alpha_{i} b_{i}(T)+\frac{1}{N_{\tau}^{2}} \sum_{i=1}^{n} \beta_{i} b_{i}(T) . \tag{27}
\end{equation*}
$$

Here, the basis splines satisfy $b_{i}\left(T_{j}\right)=\delta_{i j}$ where the $T_{j}$ with $1 \leq j \leq n$ are the node points for the given basis function. To estimate the systematics of the interpolation, we combine results from four sets of the node points $T_{j}$. The systematic error estimation also includes the difference between the $f_{\pi^{-}}$and $w_{0}$-based scale settings. We only kept fits with a $Q$ value higher than $5 \%$, which we then combined with uniform weights. The $c_{2}^{B}(T)$ and $\mathrm{M}(T)$ fits were constrained at temperatures below 130 MeV to match predictions from the hadron resonance gas (HRG) model.

In the large temperature range we performed a polynomial fit in $1 / T$. The known Stefan-Boltzmann limit was not set as the constant term of the polynomial. Hence, the fitted value of the constant is not equal to the known infinite temperature limit, and our fit only allows for interpolation


FIG. 5. $c_{2}^{B}\left(T, \hat{\mu}_{B}=0\right), d_{2}^{B}\left(T, \hat{\mu}_{B}=0\right)$ and their logarithmic temperature derivatives in the continuum limit, extrapolated from our $N_{\tau}=10,12$ and 16 lattices.
in the range where we have lattice data. The region where the two ansätze overlap give consistent results in the temperature range between 200 MeV and 280 MeV for both $c_{2}^{B}$ and its $T$ derivative. In the final result, we simply concatenate the two results at $T=250 \mathrm{MeV}$. Final results for $c_{2}^{B}(T, 0)$ and its logarithmic temperature derivative are shown in Fig. 5.

We also performed the same analysis for $d_{2}^{B}(T, 0)$, corresponding to $\chi^{Q}=0.4 \chi^{B}$. The results for this quantity and its temperature derivative are also shown in Fig. 5. At large temperatures, there is a small but statistically significant difference between $d_{2}^{B}$ and $c_{2}^{B}$. The difference of these Taylor coefficients leads to a small difference between the leading order chemical potential dependence in these two cases for high temperatures. The next corrections, corresponding to the $\lambda_{n}^{B B}$ coefficients of our resummation scheme, would probably also slightly differ in the two cases, but our lattice results are not yet precise enough to detect this difference. Therefore, we continue with the $\mu_{Q}=0$ setting for simplicity.

## C. Analysis of the coefficients $\lambda_{2}$ and $\lambda_{4}$

The analysis proceeds in the same way for all three observables we study. Denoting by $A$ either one of the observables $c_{1}^{B}, \mathrm{M}=\frac{\mu_{S}}{\mu_{B}}$ and $\chi_{2}^{S}$ and calling $B$ one of $c_{2}^{B}$, $-\frac{\chi_{1 /}^{B S}}{\chi_{2}^{S}}$ or $\chi_{2}^{S}$ respectively, and denoting the corresponding Stefan-Boltzmann corrected observables by $\tilde{A}=A / \bar{A}$ and $\tilde{B}=B / \bar{B}$ respectively, our extrapolation ansatz is defined as

$$
\begin{equation*}
\tilde{A}\left(T, \hat{\mu}_{B}\right)=\tilde{B}\left(T^{\prime}, 0\right) \tag{28}
\end{equation*}
$$

In the first step of the analysis, a spline interpolation is performed for $\tilde{A}$ at finite imaginary $\mu_{B}$ and for $\tilde{B}$ at $\mu_{B}=0$. In the second step, these splines are used for each


FIG. 6. The determination of the rescaled temperatures $T_{B B}^{\prime}$ by spline fits to the data at zero and imaginary baryochemical potential. Solid lines represent the spline fits to the lattice data, while transparent lines represent the shifts $\Delta T_{B B}=T_{B B}^{\prime}-T$ extracted from the splines.
imaginary $\hat{\mu}_{B}$, and several values of $\tilde{A}\left(T, \hat{\mu}_{B}\right)$ to determine a $T^{\prime}$ that solves Eq. (28). This procedure is illustrated in Fig. 6. After finding the values of $T^{\prime}$, we construct the quantity

$$
\begin{equation*}
\Pi\left(T, \hat{\mu}_{B}, N_{\tau}\right)=\frac{T^{\prime}\left(T, \hat{\mu}_{B}, N\right)-T}{T \hat{\mu}_{B}} \tag{29}
\end{equation*}
$$

Since $\lim _{\hat{\mu}_{B} \rightarrow 0} \Pi=\lambda_{2}^{i j}$, with $i j$ being either one of $B B, B S$ or $S S$, we can add a data point at $\mu_{B}=0$ by utilizing the formulas connecting the ordinary Taylor coefficients defined in Eq. (2) to the $\lambda_{n}$ coefficients defined in Sec. III. For reference, these formulas are listed in the Appendix. In the third step, we fit the quantity $\Pi\left(T, \hat{\mu}_{B}, N_{\tau}\right)$ with an ansatz that is linear in $1 / N_{\tau}^{2}$ and either linear or parabolic in $\mu_{B}^{2}$-i.e.:

$$
\begin{align*}
\Pi\left(T, \hat{\mu}_{B}, N_{\tau}\right)= & \lambda_{2}^{A}+\lambda_{4}^{A} \hat{\mu}_{B}^{2}+\lambda_{6}^{A} \hat{\mu}_{B}^{4} \\
& +\frac{1}{N_{\tau}^{2}}\left(\alpha^{A}+\beta^{A} \hat{\mu}_{B}^{2}+\gamma^{A} \hat{\mu}_{B}^{4}\right) \tag{30}
\end{align*}
$$

where we either fix $\lambda_{6}^{A}=\gamma^{A}=0$ or leave both as free parameters in the fit.

Systematic errors are estimated by combining several fits with uniform weights, as long as their $Q$ value is above $1 \%$. The different choices in the analysis procedure include:
(i) 3 different sets of spline node points at $\mu_{B}=0$,
(ii) 2 different sets of spline node points at finite imaginary $\mu_{B}$,
(iii) $w_{0}$ or $f_{\pi}$ based scale setting,
(iv) 2 different chemical potential ranges in the global fit: $\hat{\mu}_{B} \leq 5.5$ or $\hat{\mu}_{B} \leq 6.5$,
(v) 2 functions for the chemical potential dependence of the global fit: linear or, parabola
(vi) including the coarsest lattice, $N_{\tau}=8$, or not, in the continuum extrapolation.
This amounts to a total of $3 \times 2^{5}=96$ fits entering the systematic error estimation.

In order to calculate certain thermodynamic quantities at finite chemical potential, such as the entropy, the



FIG. 7. Expansion coefficients $\lambda_{n}^{B B}$ (top panel), $\lambda_{n}^{B S}$ (middle panel) and $\lambda_{n}^{S S}$ (bottom panel) in the continuum. Also shown are the polynomial fits used in the next section, as well as predictions from the hadron resonance gas model, which were used to constrain the fits at low temperatures.
temperature derivative of the $\lambda_{n}$ coefficients is also needed. To estimate these derivatives we perform an uncorrelated fit of the obtained expansion coefficients with a fourth order polynomial ansatz for the $\lambda_{2}^{i j}$ and a second order polynomial ansatz for the $\lambda_{4}^{i j}$. For the lower end of our temperature range, the fits are constrained to the values predicted by the hadron resonance gas model. Two points from the HRG model, at $T=125,130 \mathrm{MeV}$, are included in the fit with fictitious errors of the approximate size of the error in our lattice data, so not to dominate the fit. This approximately amounts to a $10 \%$ and a $100 \%$ errors in the $\lambda_{2}^{i j}$ and $\lambda_{4}^{i j}$, respectively. The expansion coefficients and the fits used to estimate their temperature derivative are shown in Fig. 7. All of the $\lambda_{2}$ coefficients are approximately constant in the crossover range, as is expected from the existence of the approximate scaling variable, discussed in the introduction. With the exception of $\lambda_{2}^{S S}$, the $\lambda_{2}$ coefficients all go to zero within error bars at higher end of our temperature range, as was anticipated in the discussion of Fig. 2. The $\lambda_{2}^{S S}$ is still nonzero, as the strangeness susceptibility $\chi_{2}^{S}$ tends to its Stefan-Boltzmann more slowly, due to the larger strange quark mass.

## V. THERMODYNAMICS AT REAL $\mu_{B}$

## A. Thermodynamics on the $\chi_{1}^{S}=0$ line

As we already stated, since $\mu_{Q}=0$ and $\chi_{1}^{S}=0$, on the strangeness neutral line we have $c_{1}^{B}=\chi_{1}^{B}$. Therefore, the ansatz in Eq. (17) gives us the baryon number at finite chemical potential directly. In a similar manner, the ratio $\mu_{S} / \mu_{B}$ and the strangeness susceptibility $\chi_{2}^{S}$ are calculated directly from Eq. (21) and Eq. (23), respectively. To obtain the pressure, the entropy density and the energy density, we start from $c_{1}^{B}$. For the pressure, one has to calculate the following integral:

$$
\begin{equation*}
\frac{p\left(T, \hat{\mu}_{B}\right)}{T^{4}}=\frac{p(T, 0)}{T^{4}}+\int_{0}^{\hat{\mu}_{B}} c_{1}^{B}\left(T, \hat{\mu}_{B}^{\prime}\right) \mathrm{d} \hat{\mu}_{B}^{\prime}, \tag{31}
\end{equation*}
$$

where

$$
\begin{equation*}
c_{1}^{B}\left(T, \hat{\mu}_{B}\right)=c_{2}^{B}\left(T^{\prime}, 0\right) \frac{\overline{c_{1}^{B}}\left(\hat{\mu}_{B}\right)}{\overline{c_{2}^{B}}(0)} \tag{32}
\end{equation*}
$$

and the pressure at zero chemical potential is taken from Ref. [57]. The entropy density is defined as

$$
\begin{equation*}
s=\left.\frac{\partial p}{\partial T}\right|_{\mu_{B}, \mu_{S}} \tag{33}
\end{equation*}
$$

which can be rewritten in terms of dimensionless quantities as:

$$
\begin{align*}
\hat{s} & =4 \hat{p}+\left.T \frac{\partial \hat{p}}{\partial T}\right|_{\mu_{B}} \\
& =4 \hat{p}+\left.T \frac{\partial \hat{p}}{\partial T}\right|_{\hat{\mu}_{B}}-\hat{\mu}_{B} \chi_{1}^{B} \tag{34}
\end{align*}
$$

where $\hat{s} \equiv \frac{s}{T^{3}}$ and we took into account the difference between derivatives at fixed $\mu_{B}$ versus at fixed $\hat{\mu}_{B}$. By noticing that on the strangeness neutral line

$$
\begin{align*}
\frac{\mathrm{d} \hat{p}\left(T, \hat{\mu}_{B}, \hat{\mu}_{S}\left(T, \hat{\mu}_{B}\right)\right)}{\mathrm{d} T} & =\chi_{1}^{S} \frac{\partial \hat{\mu}_{S}}{\partial T}+\frac{\partial \hat{p}}{\partial T} \\
& =\frac{\partial \hat{p}\left(T, \hat{\mu}_{B}, \hat{\mu}_{S}\left(T, \hat{\mu}_{B}\right)\right)}{\partial T} \tag{35}
\end{align*}
$$

we can write the logarithmic temperature derivative of the pressure as:

$$
\begin{align*}
& \left.T \frac{\partial \hat{p}\left(T, \hat{\mu}_{B}\right)}{\partial T}\right|_{\hat{\mu}} \\
& =T \frac{\partial \hat{p}(T, 0)}{\partial T}+\left.\frac{1}{2} \int_{0}^{\hat{\mu}_{B}^{2}} T \frac{\mathrm{~d} c_{2}^{B}\left(T^{\prime}, 0\right)}{\mathrm{d} T^{\prime}}\right|_{T^{\prime}=T\left(1+\lambda_{2}^{B B} y+\lambda_{4}^{B B} y^{2}\right)} \\
& \quad \times\left[1+\lambda_{2}^{B B} y+\lambda_{4}^{B B} y^{2}+T\left(\frac{\mathrm{~d} \lambda_{2}^{B B}}{d T} y+\frac{\mathrm{d} \lambda_{4}^{B B}}{d T} y^{2}\right)\right] d y \tag{36}
\end{align*}
$$

where $\frac{\mathrm{d} c_{2}^{B}(T)}{\mathrm{d} T}$ is calculated at $\mu_{B}=0$ as discussed previously and shown in Fig. 5.

Given the pressure and the entropy, the dimensionless energy density is given by:

$$
\begin{equation*}
\hat{\epsilon}=\hat{s}-\hat{p}+\hat{\mu}_{B} \chi_{1}^{B} \tag{37}
\end{equation*}
$$

where $\hat{\epsilon}=\frac{\epsilon}{T^{4}}$.
The continuum estimates of the dimensionless baryon number, pressure, entropy density, energy density, $\mu_{S} / \mu_{B}$ ratio and strangeness susceptibility-as computed from the expansion coefficients up to order $\lambda_{4}^{i j}$ —are shown in the various panels of Fig. 8. Notice that even with the inclusion of the $\lambda_{4}^{i j}$ coefficients, the statistical errors of our results stay well under control in the chemical potential range we study. We also compare our results to predictions from the hadron resonance gas model, which at low enough temperatures shows an excellent agreement with our lattice data. As the chemical potential increases, the crossover temperature decreases, and the agreement between out lattice results and the hadron resonance gas gets pushed to lower temperatures, as expected. Note that, similarly to our previous results for the $\mu_{Q}=\mu_{S}=0$ case in Ref. [69], none of the observables display the pathological oscillations of truncated Taylor expansions. We explicitly show in Fig. 9 a comparison of the results for the baryon density in the present work to those at $\mu_{Q}=\mu_{S}=0$ in Ref. [69]. The two sets of results obviously tend to different Stefan-Boltzmann
limits, as a result of the different value of the strange chemical potential (see Fig. 8, bottom left panel). Besides the covered temperature range being larger in this work, we also note that the size of the error is slightly improved. A similar, although less pronounced effect, will appear in the pressure, entropy and energy density, as they are dominated by their value at $\mu_{B}=0$, which is the same in both cases.

For the extrapolation beyond the strangeness neutral line, the ratio of the baryon-strangeness correlator to the strangeness fluctuations $\chi_{11}^{B S} / \chi_{2}^{S}$ will also be needed. In order to obtain this ratio, we simply note that:

$$
\begin{align*}
-\frac{\chi_{11}^{B S}}{\chi_{2}^{S}} & =\frac{\mathrm{d} \hat{\mu}_{S}}{\mathrm{~d} \hat{\mu}_{B}} \\
& =\frac{\mathrm{d}}{\mathrm{~d} \hat{\mu}_{B}}\left[\hat{\mu}_{B} f_{B S}\left(T_{B S}^{\prime}\left(T, \hat{\mu}_{B}\right)\right)\right] \\
& =\hat{\mu}_{B}\left[f_{B S}\left(T_{B S}^{\prime}\right)+\frac{\partial f_{B S}\left(T_{B S}^{\prime}\left(T, \hat{\mu}_{B}\right)\right)}{\partial \hat{\mu}_{B}}\right] \tag{38}
\end{align*}
$$

where we used the shorthand notation $f_{B S}(T) \equiv$ $\mathrm{M}(T, 0)=-\frac{\chi_{11}^{B S}}{\chi_{2}^{S}}\left(T, \hat{\mu}_{B}=0\right)$. Our results for $\chi_{11}^{B S} / \chi_{2}^{S}$ are shown in Fig. 10. In addition to being needed for extrapolation to finite strangeness, this ratio is also of interest for freeze-out phenomenology [78].

## B. Beyond strangeness neutrality

The quantities calculated so far can also be used to extrapolate the equation of state to small values of the strangeness density, slightly off the $\chi_{1}^{S}=0$ line. Let us denote the value of the dimensionless strange quark chemical potential that solves $\chi_{1}^{S}=0$ at fixed $T$ and $\hat{\mu}_{B}$ as $\hat{\mu}_{S}^{\star}$. Still considering fixed $\hat{\mu}_{B}$ and $T$, but changing $\hat{\mu}_{S}$ slightly from the strangeness neutral choice by a small amount

$$
\begin{equation*}
\Delta \hat{\mu}_{S} \equiv \hat{\mu}_{S}-\hat{\mu}_{S}^{\star} \tag{39}
\end{equation*}
$$

the dimensionless strangeness and baryon densities become:

$$
\begin{gather*}
\chi_{1}^{S}\left(\hat{\mu}_{S}\right) \approx \chi_{2}^{S}\left(\hat{\mu}_{S}^{\star}\right) \Delta \hat{\mu}_{S}  \tag{40}\\
\chi_{1}^{B}\left(\hat{\mu}_{S}\right) \approx \chi_{1}^{B}\left(\hat{\mu}_{S}^{\star}\right)+\chi_{11}^{B S}\left(\hat{\mu}_{S}^{\star}\right) \Delta \hat{\mu}_{S}, \tag{41}
\end{gather*}
$$

where we only kept the linear leading order terms in $\Delta \hat{\mu}_{S}$. We will express thermodynamic quantities in terms of the strangeness-to-baryon fraction:

$$
\begin{equation*}
R=\frac{\chi_{1}^{S}}{\chi_{1}^{B}}=\frac{\chi_{2}^{S}\left(\hat{\mu}_{S}^{\star}\right) \Delta \hat{\mu}_{S}}{\chi_{1}^{B}\left(\hat{\mu}_{S}^{\star}\right) \Delta \hat{\mu}_{S}+\chi_{11}^{B S}\left(\hat{\mu}_{S}^{\star}\right)} \tag{42}
\end{equation*}
$$

Inverting this equation we get:


FIG. 8. The dimensionless baryon density (top left panel), pressure (top right panel), entropy (middle left panel), energy density (middle right panel), strangeness chemical potential-to-baryochemical potential ratio (bottom left panel) and strangeness susceptibility (bottom right panel) as functions of the temperature at different values of the real chemical potential. The solid lines show the predictions of the hadron resonance gas model for the corresponding temperature.

$$
\begin{equation*}
\Delta \hat{\mu}_{S}=\frac{R \hat{\chi}_{1}^{B}\left(\hat{\mu}_{S}^{\star}\right)}{\chi_{2}^{S}\left(\hat{\mu}_{S}^{\star}\right)-R \chi_{11}^{B S}\left(\hat{\mu}_{S}^{\star}\right)} . \tag{43}
\end{equation*}
$$

This quantity is shown for $\hat{\mu}_{B}=2$ as a function of the temperature for various values of $R$ in Fig. 11. Substituting Eq. (43) into Eq. (40) we obtain-to leading order in $R$ :

$$
\begin{equation*}
\frac{\chi_{1}^{B}\left(T, \hat{\mu}_{B}, R\right)}{\chi_{1}^{B}\left(T, \hat{\mu}_{B}, R=0\right)} \approx 1+R \frac{\chi_{11}^{B S}\left(T, \hat{\mu}_{B}, R=0\right)}{\chi_{2}^{S}\left(T, \hat{\mu}_{B}, R=0\right)}, \tag{44}
\end{equation*}
$$

where all quantities on the right hand side are along the strangeness neutral line. We show the results of a leading order extrapolation of the dimensionless baryon density as a function of $T$ at $\hat{\mu}_{B}=2$ for several values of $R$ in Fig. 12.


FIG. 9. Comparison of the results for the baryon density in the case of strangeness neutrality (this work, darker shades), and in the case with $\mu_{Q}=\mu_{S}=0$ (Ref. [69], lighter shades), at different values of the real chemical potential. The solid lines show the corresponding predictions of the hadron resonance gas model.


FIG. 10. The ratio $\chi_{11}^{B S} / \chi_{2}^{S}$ as a function of the temperature for several values of $\hat{\mu}_{B}$. The solid lines show the predictions of the hadron resonance gas model.


FIG. 11. Shift of the strangeness chemical potential as a function of the temperature at $\hat{\mu}_{B}=2$, at various values of the strangeness-to-baryon ratio $R=\chi_{1}^{S} / \chi_{1}^{B}$. The solid lines show the exact solution of $0.4 \chi_{1}^{B}=\chi_{1}^{S}$ in the hadron resonance gas model, while the dashed lines show the model results for the approximation of Eq. (43).


FIG. 12. The dimensionless baryon density as a function of the temperature at $\hat{\mu}_{B}=2$, for various values of the strangeness-tobaryon ratio $R=\chi_{1}^{S} / \chi_{1}^{B}$.


FIG. 13. The leading order Taylor expansion coefficient of the pressure in the strangeness-to- baryon ratio $R$ on the strangeness neutral line as a function of the temperature for several fixed values of $\hat{\mu}_{B}$.

At the strangeness neutral line the $\mathcal{O}(R)$ correction of the pressure vanishes. The leading order correction gives:

$$
\begin{equation*}
\hat{p}\left(T, \hat{\mu}_{B}, R\right) \approx \hat{p}\left(T, \hat{\mu}_{B}, R\right)+\frac{1}{2} \frac{\mathrm{~d}^{2} \hat{p}}{\mathrm{~d} R^{2}}\left(T, \hat{\mu}_{B}\right) R^{2}, \tag{45}
\end{equation*}
$$

where

$$
\begin{equation*}
\frac{\mathrm{d}^{2} \hat{p}}{\mathrm{~d} R^{2}}\left(T, \hat{\mu}_{B}\right)=\frac{\left(\chi_{1}^{B}\left(T, \hat{\mu}_{B}\right)\right)^{2}}{\chi_{2}^{S}\left(T, \hat{\mu}_{B}\right)} . \tag{46}
\end{equation*}
$$

This leading order coefficient is show in Fig. 13 for several values of $\hat{\mu}_{B}$.

## VI. SUMMARY AND DISCUSSION

In this work, we generalized the extrapolation procedure we introduced in Ref. [69] for the equation of state to the case of strangeness neutrality. Using zero and imaginary chemical potential simulations we determined the first two coefficients in the scheme $\lambda_{2}^{i j}$ and $\lambda_{4}^{i j}$ for the baryon density, the strangeness chemical potential and the strangeness fluctuations. Using these coefficients, we extrapolated the equation of state up to a baryochemical potential-to-temperature ratio of $\hat{\mu}_{B}=3.5$, considerably improving the range covered by first principle lattice calculations. Just like for the case of $\mu_{S}=\mu_{Q}=0$ in Ref. [69], also for the strangeness neutral case we observed that our extrapolation scheme is free of the unphysical oscillations that plague the fixed order Taylor expansions at higher chemical potentials.

We also introduced a Stefan-Boltzmann correction to improve the rate of convergence of our scheme at high temperatures. Thanks to this, the coefficient $\lambda_{2}^{B B}$ goes to zero at high temperatures.

We moved beyond strangeness neutrality by calculating the baryon-strangeness correlator to strangeness susceptibility ratio $\chi_{11}^{B S} / \chi_{2}^{S}$ at finite real $\hat{\mu}_{B}$ on the strangeness neutral line. This allows one to perform a leading order extrapolation of the baryon number in the ratio $R=\chi_{1}^{S} / \chi_{1}^{B}$.

Apart from increasing precision through higher statistics, this work can be extended in the future to take the small isospin imbalance $\chi_{1}^{Q}=0.4 \chi_{1}^{B}$ of lead and gold nuclei into account beyond leading order. With significantly more data taking, a full scan of the $\mu_{B}-\mu_{S}$ plane is possible, thus determining the equation of state for higher values of the strangeness-to-baryon ratio. It will be valuable to see whether the calculation of even higher order coefficients (e.g., $\kappa_{6}^{i j}$ ) will alter the presented picture or introduce significant improvements in the range explored by the RHIC Beam Energy Scan program.

We also mention that, while for the ordinary Taylor expansion of the pressure estimates of the radius of convergence exist both on coarse lattices [26,27,71,79] and from phenomenological arguments [80,81], the ultimate convergence properties of our scheme are completely unknown at the moment. As our ansatz is not a Taylor series for the pressure, its domain of convergence is not necessarily a circle in the complex plane. Furthermore, the scheme is generalizable also for non-polynomial approximations of $T^{\prime}$, such as rational functions. The domain of convergence of our scheme is an interesting theoretical question. It, however, has no practical consequence at the moment, as the $\lambda_{4}^{i j}$ are already consistent with zero, leading to fast apparent convergence at low orders.

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## APPENDIX: $\lambda_{2}^{i j}$ COEFFICIENTS FROM TAYLOR COEFFICIENTS

For reference, we list here the relationships between our $\lambda_{2}$ coefficients and ordinary Taylor coefficients. For the expansion coefficient of the baryon density, we get:

$$
\lambda_{2}^{B B}=\frac{1}{6 T f^{\prime}(T)}\left(c_{4}^{B}(0, T)-\frac{\overline{c_{4}^{B}}(0)}{\overline{c_{2}^{B}}(0)} f(T)\right),
$$

where $f(T)=\frac{\mathrm{d}^{2} \log Z}{\mathrm{~d} \mu_{B}^{2}}\left(\mu_{B}=0, T\right)$. For the expansion coefficient of the strangeness chemical potential we get

$$
\lambda_{2}^{B S}=\frac{1}{T f^{\prime}(T)} s_{3}(T)=\frac{1}{6 T f^{\prime}(T)} \frac{\mathrm{d}^{3} \hat{\mu}_{S}}{\mathrm{~d} \hat{\mu}_{B}^{3}}(T),
$$

where $\frac{\hat{\mu}_{S}}{\hat{\mu}_{B}}\left(\hat{\mu}_{B}, T\right)=s_{1}(T)+s_{3}(T) \hat{\mu}_{B}^{2}+s_{5}(T) \hat{\mu}_{B}^{4}+\cdots$ and $f(T)=\lim _{\hat{\mu}_{B} \rightarrow 0} \frac{\hat{\mu}_{S}}{\hat{\mu}_{B}}\left(\mu_{B}, T\right)=-\frac{\chi_{1}^{B S}}{\chi_{2}^{S}}(0, T)$. For the expansion coefficient of the strangeness susceptibility we get

$$
\lambda_{2}^{S S}=\frac{1}{2 T f^{\prime}(T)} S_{2, \mathrm{sym}}^{\mathrm{NLO}}(0, T)
$$

where $f(T)=\chi_{2}^{S}\left(\mu_{B}=0, T\right)$.
In principle, the $\lambda_{4}$ coefficients can also be expressed using the Taylor coefficients at $\mu \equiv 0$. For these one needs the Taylor coefficients up to sixth order and the second temperature derivative of the second order coefficients. For the quantities discussed in this paper we have:

$$
\begin{gather*}
\lambda_{4}^{\mathrm{BB}}(T)=\frac{1}{360 T} \frac{1}{\bar{c}_{2}^{B}(0)^{2} f^{\prime}(T)^{3}} \cdot\left[3 \bar{c}_{2}^{B}(0)^{2} c_{6}^{B}(0, T) f^{\prime}(T)^{2}-10 \bar{c}_{4}^{B}(0) f^{\prime}(T)^{2}\left(\bar{c}_{2}^{B}(0) c_{4}^{B}(0, T)-\bar{c}_{4}^{B}(0) f(T)\right)\right. \\
\left.-5 f^{\prime \prime}(T)\left(\bar{c}_{2}^{B}(0) c_{4}^{B}(0, T)-\bar{c}_{4}^{B}(0) f(T)\right)^{2}\right]  \tag{A1}\\
\lambda_{4}^{B S}(T)=\frac{s_{5}(T)}{T f^{\prime}(T)}-\frac{s_{3}(T)^{2} f^{\prime \prime}(T)}{2 T f^{\prime}(T)^{3}} \\
=\frac{1}{120 T f^{\prime}(T)} \frac{\mathrm{d}^{5} \hat{\mu}_{S}}{\mathrm{~d} \hat{\mu}_{B}^{5}}(T)-\frac{f^{\prime \prime}(T)}{72 T f^{\prime}(T)^{3}}\left(\frac{\mathrm{~d}^{3} \hat{\mu}_{S}}{\mathrm{~d} \hat{\mu}_{B}^{3}}(T)\right)^{2}  \tag{A2}\\
\lambda_{4}^{S S}(T)=\frac{1}{24 T f^{\prime}(T)^{3}}\left(S_{2, \mathrm{sym}}^{\mathrm{NNLO}}(0, T) f^{\prime}(T)^{2}-3 f^{\prime \prime}(T) S_{2, \mathrm{sym}}^{\mathrm{NLO}}(0, T)^{2}\right), \tag{A3}
\end{gather*}
$$

with

$$
\begin{align*}
& S_{2, \mathrm{sym}}^{\mathrm{NLO}}(0, T)=\chi_{22}^{B S}(0, T)+2 s_{1}(T) \chi_{13}^{B S}(0, T)+s_{1}(T)^{2} \chi_{4}^{S}(0, T)  \tag{A4}\\
& S_{2, \mathrm{sym}}^{\mathrm{NNLO}}(0, T)= \chi_{42}^{B S}(0, T)+4 s_{1}(T) \chi_{33}^{B S}(0, T)+6 s_{1}(T)^{2} \chi_{24}^{B S}(0, T)+4 s_{1}(T)^{3} \chi_{15}^{B S}(0, T) \\
&+s_{1}(T)^{4} \chi_{6}^{S}(0, T)+24 s_{3}(T) \chi_{13}^{B S}(0, T)+24 \chi_{4}^{S}(0, T) s_{1}(T) s_{3}(T) \tag{A5}
\end{align*}
$$

where we used the expansion coefficients of $\hat{\mu}_{S}\left(\hat{\mu}_{B}\right)$ :

$$
\begin{gather*}
s_{1}=-\frac{\chi_{11}^{B S}}{\chi_{2}^{S}}  \tag{A6}\\
s_{3}=-\frac{1}{6 \chi_{2}^{S}}\left[\chi_{4}^{S} s_{1}^{3}+3 \chi_{13}^{B S} s_{1}^{2}+3 \chi_{22}^{B S} s_{1}+\chi_{31}^{B S}\right]  \tag{A7}\\
s_{5}=-\frac{1}{120 \chi_{2}^{S}}\left[+\chi_{6}^{S} s_{1}^{5}+5 \chi_{15}^{B S} s_{1}^{4}+10 \chi_{24}^{B S} s_{1}^{3}+60 \chi_{4}^{S} s_{1}^{2} s_{3}+120 \chi_{13}^{B S} s_{1} s_{3}+60 \chi_{22}^{B S} s_{3}+10 \chi_{33}^{B S} s_{1}^{2}+5 \chi_{42}^{B S} s_{1}+\chi_{51}^{B S}\right] \tag{A8}
\end{gather*}
$$

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