ANOMALOUS RADIOCARBON AGES FOUND IN CAMPANIAN IGNIMBRITE DEPOSIT OF THE MEDITERRANEAN DEEP-SEA CORE CT85-5

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ABSTRACT. A detailed radiocarbon chronology has been established for the deep-sea core CT85-5 from the Tyrrhenian Sea. This chronology, which is based on the analysis of foraminifera shells, shows a set of reversed ¹⁴C ages for sediments deposited during the eruption of the Campanian Ignimbrite (~40 ka cal BP). The anomalous young ¹⁴C ages coincide with elevated concentrations of ¹⁰Be measured in the same core. Although reversals in ¹⁴C ages have been previously found in other records at 40 ka cal BP, such extreme changes have not been observed elsewhere. The enhancement in ¹⁴C concentration in CT85-5 sediments associated with the Campanian Ignimbrite is equivalent to an apparent age ~15 ka younger than the age for the sediments deposited shortly before the eruption. Here, we present consistent results of repeated measurements showing no analytical problems that can explain the observed rapid changes in ¹⁴C of this particular record.

STUDY BACKGROUND

The detection of the "¹⁰Be peak" dated to 35 and 40 ka BP found in Antarctic ice cores (Raisbeck et al. 1987) and later in Greenland ice cores (Yiou et al. 1997) prompted studies of variable production rates of cosmogenic isotopes. The low geomagnetic field of the Laschamp event, which was first discovered in a lava flow of the Massif Central, France (Bonhommet and Zähringer 1969), and dated to ~40 ka cal BP (see Valet et al. 2007 for a review) and more recently to 40.65 ± 0.65 ka cal BP (Singer et al. 2009), has been discussed as the most likely cause of these changes. In the 1990s, measurements of ¹⁰Be concentrations in deep-sea sediments became available showing enhanced ¹⁰Be levels around 40 ka cal BP (McHargue et al. 1995, 2000; Robinson et al. 1995). Finally, reconstructions of geomagnetic field intensity in deep-sea sediments (Laj et al. 2000, 2002) and correlation with the ice-core record of cosmogenic isotopes ¹⁰Be and ³⁶Cl has shown that the intervals of low geomagnetic field coincided with periods of increased flux of cosmogenic isotopes (Baumgartner et al. 1998; Wagner et al. 2000).

Castagnoli et al. (1995) published ¹⁰Be measurements on sediment core CT85-5 recovered from the Tyrrhenian Sea. The detection of a clear ¹⁰Be peak, which was coincident with the Campanian Ignimbrite (CI) ash layer, gave the background to our radiocarbon analysis across the sediment core and in particular the section corresponding to the ¹⁰Be enhancement. The surprising results of our ¹⁴C dating were first reported as a figure in Giaccio et al. (2006). In this study, we present the complete set of the results including all the tests carried out to assess the validity of the accelerator mass spectrometry (AMS) ¹⁴C dating performed on the planktonic foraminifera shells of CT85-5 core. These results show that the high ¹⁴C content coincides with the previously observed ¹⁰Be enhancement (Castagnoli et al. 1995). The detailed ¹⁴C-¹⁰Be comparison is allowed by using measurements of these 2 cosmogenic radioisotopes in the same core.

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MATERIALS AND METHODS

The core CT85-5 was recovered from the Tyrrhenian Sea ($40^{\circ}19'02''N$, $11^{\circ}15'42''E$) in 1985 at a water depth of 2833 m. This core had been the subject of various studies including measurements of ¹⁰Be (Castagnoli et al. 1995), carbonate content (Castagnoli et al. 1992), and magnetic properties (Castagnoli et al. 1998).

Samples were taken at the Turin Laboratory in 1992, 1993, and 1998 for our original study. More recently, in 2009, the second half of the core CT85-5 was sampled. For each ¹⁴C analysis, at least 2000 specimens of *Globigerina bulloides* or mixed planktonic foraminifera were handpicked and processed for AMS analysis. In addition, 1 sample was prepared from pteropods, i.e. shells of pelagic sea snails (Table 1). Prior to sieving, the sediment was soaked in 10% solution of H_2O_2 . Shells were then picked and treated in an ultrasonic bath of H₂O to clean the surface and dissolved in concentrated (85%) phosphoric acid. The weight of all the picked samples was over 20 mg, sufficient for preparation of targets that contained 2 mg of pure carbon. Carbon dioxide was reduced to graphite (Hajdas et al. 2004b) and the ¹⁴C/¹²C ratios were measured using the ETH 6MV tandem accelerator (Bonani et al. 1987) and the 200kV MICADAS, which is a dedicated ¹⁴C AMS instrument (Synal et al. 2007). Finite ¹⁴C ages were obtained down to the layer of tephra at a depth of 290– 333 cm and then down to a depth of 359 cm, where the limit of the 14 C dating method was reached for tandem measurements in the earlier stages of this study. The ages of the 6 deepest samples obtained in that set of data were indistinguishable from the background value (i.e. the measured $^{14}C/$ 12 C ratio of sample was smaller than double 14 C/ 12 C ratio of the blank). For these samples, activity was not corrected for the blank and the calculated ages are referred to as "older than" and are considered to be minimum ages (Figure 1, Table 1).

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	Deptha		¹⁴ C age	
Lab nr	(cm)	$F^{14}C$	(BP)	Comments ^b
ETH-12683	95	0.30170 ± 0.00383	9625 ± 100	
ETH-12261	100	0.27640 ± 0.00272	$10,330 \pm 80$	G. bull
ETH-19538	105	0.25800 ± 0.00344	$10,880 \pm 110$	
ETH-12262	111	0.25200 ± 0.00320	$11,070 \pm 100$	G. bull
ETH-12686	120	0.20918 ± 0.00315	$12,570 \pm 120$	G. bull
ETH-12688	139	0.18121 ± 0.00257	$13,720 \pm 120$	G. bull
ETH-12098	140	0.18214 ± 0.00218	$13,680 \pm 95$	
ETH-12689	147	0.15935 ± 0.00226	$14,750 \pm 110$	
ETH-12690	165	0.12564 ± 0.00191	$16,660 \pm 120$	
ETH-14163	180	0.10188 ± 0.00162	$18,350 \pm 130$	
ETH-14164	190	0.08093 ± 0.00145	$20,200 \pm 140$	
ETH-12691	200	0.05675 ± 0.00138	$23,050 \pm 195$	
ETH-13222	230	0.03738 ± 0.00111	$26,400 \pm 240$	
ETH-13223	250	0.02756 ± 0.00111	$28,850 \pm 330$	
ETH-13224	270	0.01732 ± 0.00096	$32,580 \pm 450$	
ETH-14048	275	0.01843 ± 0.00085	$32,080 \pm 370$	
ETH-37858	275	0.01590 ± 0.00054	$33,265 \pm 275$	1, 2
ETH-14049	279	0.01794 ± 0.00082	32.300 ± 370	

Table 1 Measured concentrations of ${}^{14}C$ (F ${}^{14}C$ as defined by Reimer et al. 2004) and radiocarbon ages of mixed planktonic foraminifera shells (samples made up of only *Globigerina bulloides* are marked as *G bull*) and 1 pteropod shell sample from the deep-sea core CT85-5.

	Depth ^a		¹⁴ C age	
Lab nr	(cm)	$F^{14}C$	(BP)	Comments ^b
ETH-12099	280	0.01642 ± 0.00065	$33,010 \pm 320$	
ETH-13225.1	280	0.09058 ± 0.00170	$19,290 \pm 150$	
ETH-13225.2	280	0.08206 ± 0.00143	$20,090 \pm 140$	
ETH-19537.1	280	0.07071 ± 0.00172	$21,280 \pm 195$	
ETH-19537.2	280	0.08202 ± 0.00128	$20,090 \pm 125$	2
ETH-37859	280	0.09332 ± 0.00086	$19,050 \pm 75$	1, 2
ETH-14050	282	0.03347 ± 0.00103	$27,290 \pm 250$	
ETH-17306.1	282	0.04071 ± 0.00120	$25,710 \pm 240$	Outside
ETH-17306.2	282	0.03885 ± 0.00114	$26,090 \pm 240$	Inside, 35% leached away
ETH-17307.1	283	0.11660 ± 0.00177	$17,260 \pm 120$	Outside
ETH-17307.2	283	0.11787 ± 0.00176	$17,180 \pm 120$	Inside 50%
ETH-24997.1	283	0.04760 ± 0.00107	$24,460 \pm 180$	
ETH-24997.2	283	0.05232 ± 0.00124	$23,700 \pm 190$	
ETH-24997.3	283	0.04398 ± 0.00112	$25,095 \pm 205$	2
ETH-24856.1	284	0.01677 ± 0.00075	$32,840 \pm 360$	Outside
ETH-24856.2	284	0.01773 ± 0.00090	$32,390 \pm 410$	Inside, 75% leached away
ETH-24856.3	284	0.02182 ± 0.00109	$30,730 \pm 400$	Pteropods
ETH-24856.4	284	0.02241 ± 0.00078	$30,510 \pm 280$	-
ETH-37860	284	0.06021 ± 0.00079	$22,570 \pm 105$	1, 2
ETH-12100.1	288	0.07324 ± 0.00143	$21,000 \pm 160$	
ETH-12100.2	288	0.05177 ± 0.00207	$23,790 \pm 320$	
ETH-12263	303	0.05395 ± 0.00106	$23,760 \pm 180$	G. bull
ETH-12264	314	0.04032 ± 0.00122	$25,975 \pm 245$	G. bull
ETH-38332	335	0.01126 ± 0.00111	$36,040 \pm 790$	2
ETH-12693	337	0.01181 ± 0.00073	$35,660 \pm 500$	G. bull
ETH-12694	344	0.01213 ± 0.00076	$35,450 \pm 510$	G. bull
ETH-10629	346	0.00768 ± 0.00074	$39,110 \pm 770$	
ETH-10630	352	0.00504 ± 0.00069	$42,500 \pm 1100$	
ETH-38333	358	0.00543 ± 0.00106	$41,900 \pm 1570$	2
ETH-10631	359	0.00588 ± 0.00069	$41,260 \pm 940$	
ETH-10634	381	0.00631 ± 0.00044	>40,690	
ETH-10635	385	0.00709 ± 0.00048	>39,760	
ETH-11970	388	0.00879 ± 0.00039	>38,030	
ETH-10636	395	0.00769 ± 0.00049	>39,100	
ETH-38334	395	0.00270 ± 0.00026	>47,520	2
ETH-37865	414	0.00453 ± 0.00037	>43,350	2
ETH-37866	459	0.00337 ± 0.00036	>45,720	2

Table 1 Measured concentrations of ${}^{14}C$ (F ${}^{14}C$ as defined by Reimer et al. 2004) and radiocarbon ages of mixed planktonic foraminifera shells (samples made up of only *Globigerina bulloides* are marked as *G bull*) and 1 pteropod shell sample from the deep-sea core CT85-5. *(Continued)*

^aTop of the 1-cm slice of sediment.

^b1= Additional samples 2009; 2 = AMS with MICADAS.



Figure 1 14 C ages of planktonic foraminifera shells are plotted versus depth in core CT85-5: arrows show minimum ages (i.e. close to blank value, see Table 1). Lower panel shows concentration of 10 Be in sediments of CT85-5 (Castagnoli et al. 1995); the Campanian Ignimbrite is marked as a box. Note that this sediment layer, which is 43 cm thick, was an instantaneous deposit (see text).

RESULTS

The ¹⁴C chronology of the last 40 ka in the core CT85-5 shows regular features, but a reversal in ¹⁴C ages appears between 280 and 336 cm (Figure 1), which corresponds to a zone containing the deposit of the Campanian Ignimbrite tephra layer (290 to 333 cm) and just above the ash. Also, the interval partially overlaps with the layer where an enhanced ¹⁰Be concentration in sediment was found (Castagnoli et al. 1995). As shown in Figure 1, a twofold ¹⁰Be peak occurred: the ¹⁰Be concentration increased within the interval 273–384 cm and then the increase was interrupted at 290–339 cm depth by the CI eruption (dilution effect of ¹⁰Be atoms in the volcanic material). At 289 cm, where the glass shards disappear, the ¹⁰Be concentration increased as a spike, reaching its maximum value at 280 cm depth. We note that the highest ¹⁴C content were reached at about the same depth, i.e. at 280–283 cm. Because of the unusual magnitude of the age reversal, various tests have been performed to exclude any possible contamination as well as problems with measurements of very old samples.

Firstly, the possibility of contamination of the material is considered. The most probable source would be the deposition of secondary calcite on the surface of dated shells. We have performed leaching experiments of foraminifera in samples that had high ¹⁴C content and appeared "too young." Our procedure involves on-line acidification of up to 50% of the sample material (surface) and subsequent separation of the produced CO_2 from the remaining residue (inside) (Hajdas et al. 2004a). Both fractions of those samples were dated subsequently, showing no significant difference

in ages between fractions (282 cm, ETH-17306; 283 cm, ETH-17307; and 284 cm, ETH-24856) (Table 1).

A second potential reason for the younger ages is contamination with modern carbon and an underestimated blank correction. The blank values for IAEA marble C-1 for earlier measurements with the EN tandem were close to 0.005 $F^{14}C$ and samples stratigraphically older than those, showing enhanced ¹⁴C, i.e. below 333 cm, were successfully dated (Figure 1). Moreover, measurements with the MICADAS system (<0.003 $F^{14}C$ on C-1, IAEA marble) confirmed our earlier results (Table 1).

Our leaching experiments and duplicate analyses provide evidence that the enhanced ¹⁴C levels in foraminifera shells deposited during or shortly after the volcanic eruption of Campanian Ignimbrite cannot be explained by limitations of the ¹⁴C measurements or by contamination by secondary carbonates.

There is no plausible explanation for the intrusion of younger sediments into this section of the core. The sediments showing the inversed ages are ~40 cm above the bottom of the core section; thus, contamination at the core break-off can be ruled out (Castagnoli et al. 1995). Additional measurements were obtained on samples from a nearby core CT85-6 (Table 2). The record of CT85-6 is shorter because the core was broken when penetrating the Campanian Ignimbrite. The ¹⁴C ages in this core also show an inversion, however, not to such an extent as the one observed in CT85-5 (Figure 2).

Table 2 Concentrations of ¹⁴ C (F ¹⁴ C) and radiocarbon ages measured on material from the core
CT85-6 (mixed planktonic foraminifera). Correlation between cores was based on the Campanian
Ignimbrite (CI) layer and CaCO ₃ (Castagnoli et al. 1992). The top of the CI layer is at 290 cm depth.

	Depth		¹⁴ C age
Lab nr	(cm)	F ¹⁴ C	(BP)
ETH-14165	270	0.02656 ± 0.00078	$29,150 \pm 240$
ETH-14166	275	0.02484 ± 0.00079	$29,680 \pm 260$
ETH-14167	280	0.02417 ± 0.00074	$29,900 \pm 250$
ETH-14051	286	0.01474 ± 0.00078	$33,880 \pm 430$
ETH-14597	287	0.01438 ± 0.00077	$31,680 \pm 360$
ETH-14598	289	0.01938 ± 0.00087	$29,710 \pm 350$
ETH-15197	290	0.01709 ± 0.00096	$34,070 \pm 480$
ETH-15198	292	0.01382 ± 0.00083	$35,180 \pm 515$
ETH-15051	293	0.02475 ± 0.00108	$32,690 \pm 450$
ETH-14052	295	0.01254 ± 0.00080	$34,070 \pm 430$
ETH-14599	296	0.01438 ± 0.00077	$33,600 \pm 400$

The presence of the tephra layer found in the CT85-5 core at 290–333 cm depth is important for our results because we can use it as a time marker. This layer, which has previously been correlated with the Campanian Ignimbrite (marine tephra C-13 or Y-5) (Castagnoli et al. 1995), has a lower carbonate content than the surrounding sediment (reduced to ~25%). Nevertheless, we were able to obtain 2 samples from this interval. Correlation of nearby Mediterranean deep-sea cores to North Atlantic cores showed that this eruption preceded Heinrich event 4 (Paterne et al. 1999).

As described in Castagnoli et al. (1995), the thick layer of trachytic glass admixed with carbonate mud is found in the CT85-5 core at depths between 333 and 290 cm. The layer contains some grains of sanidine, plagioclase, biotite, and Na pyroxene. In the samples from this interval and sediment above it (280–290 cm), which had been washed for foraminifera, biotite and volcanic glass shards could be observed, confirming that inversed ¹⁴C ages are associated with the ash layer. The chemical



Figure 2 Reversed ¹⁴C ages of planktonic foraminifera shells from the section in core CT85-5 core shown together with ¹⁴C ages from a nearby core CT85-6. In both cores, the top of the CI tephra is at 290 cm.

composition of the glassy material (Castagnoli et al. 1995) points to the Campanian Ignimbrite eruption in the Phleagrean Fields, which is a well-known stratigraphic marker in the Mediterranean region. This powerful eruption left distinct deposits of ash both on land and in marine sediments in the region (Barberi et al. 1978) and most probably had a large impact on the environments in the region. The first ¹⁴C ages obtained on carbonized wood found in lava of the Campanian Ignimbrite had a large spread between 28 and 38 ka BP (Alessio et al. 1973, 1974). Other independent chronometers such as the ⁴⁰K/⁴⁰Ar or ⁴⁰Ar/³⁹Ar dating method provided ages for the lava that ranged from 34 to 38 ka cal BP (Wulf et al. 2004 and references cited therein). The most recent ⁴⁰Ar/³⁹Ar dating of the Campanian Ignimbrite ash revealed an age of 39.28 \pm 0.11 ka cal BP (De Vivo et al. 2001) and 41.1 \pm 2.1 ka cal BP (Ton-That et al. 2001).

DISCUSSION

The large change in ¹⁴C concentration observed in the CT85-5 core appears to be a rather short event. In spite of its thickness of 43 cm, the sediment layer in which the reversed ages were measured represents a very short event of volcanic eruption. Reversed ages are also found in the overlying 10 cm of sediments that have been described as turbidite deposits, probably associated with the CI event. Therefore, the short ¹⁴C event observed in this core might be unique and difficult to replicate in other records. Most of the published records of the last 50,000 yr show evidence of an increase in atmospheric ¹⁴C level that occurred around 40 ka cal BP due to the Laschamp excursion, but no reversal (Hughen et al. 2004; Reimer et al. 2009 and references therein). However, one of the first published indications of an age reversal at 40.3–41.7 ka cal BP is the record from core PS2644

in the Icelandic Sea (Voelker et al. 2000). This event precedes Heinrich event 4 and has been correlated with low values of paleomagnetic intensity in PS2644 and, via δ^{18} O, with enhanced values of cosmogenic ³⁶Cl and ¹⁰Be measured in ice cores (Wagner et al. 2000; Laj et al. 2002).

Other anomalously young ¹⁴C ages have been observed and reported in marine sediments corresponding to the interval close to the Laschamp event (~40 ka cal BP). For example, aberrant ¹⁴C ages have been reported for paleomagnetic reconstructions in the Arctic Ocean sediment cores (Nowaczyk et al. 2003). This raises the question of whether the observed 14 C anomaly can be explained by the Laschamp event. However, model calculations show that switching off the geomagnetic dipole field increases the ¹⁴C production rate by only about a factor 2.5 compared to the present-day production rate. To quantify these effects, we used a box diffusion carbon cycle model (Oeschger et al. 1975), which showed that a doubling in the production rate for 3 ka leads to an increase of the atmospheric ${}^{14}C/{}^{12}C$ ratio by about 40%, corresponding to an age reduction by 3200 vr, which is much less than the observed 15,000 yr. Reducing the thermohaline circulation is more effective. A complete shutdown, which is a rather unrealistic assumption, leads to an annual increase of the atmospheric ${}^{14}C/{}^{12}C$ ratio by about 0.1%, assuming the present production rate. A more realistic reduction of the eddy diffusion constant from the present value of 4000 to 1000 m² yr⁻¹ causes an increase of about 0.01% per year. In conclusion, neither a change in the production rate nor a drastic reduction in the global thermohaline circulation seems feasible to explain the observed dramatic change in ¹⁴C measured in the CT85-5 record.

CONCLUSIONS

The ¹⁴C chronology of the Mediterranean deep-sea core CT85-5 shows a period of reversed ¹⁴C ages coinciding with high levels of ¹⁰Be that were found in the sediments deposited before and after the eruption of the Campanian Ignimbrite (CI). Based on the radiometric ages of the CI eruption (40 K/ 40 Ar or 40 Ar/ 39 Ar) and on stable isotopic correlations obtained for Mediterranean deep-sea cores, this extreme ¹⁴C event can be dated to around 40 ka cal BP. We note that, despite several anomalously young ages in other records, the amplitude of the ¹⁴C changes measured in the foraminifera of the layer in the CT85-5 core is unique and has not yet been confirmed by any other data. We are convinced that these too-young ¹⁴C ages are not the result of any analytical problems. Although a plausible mechanism that could explain such extreme reversals in ¹⁴C age as observed in this core remains to be found, our results are presented here to stimulate discussion and comparison with other possibly unpublished studies.

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