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Preservation of sub-microscopic structural relicts in micas from the Gran Paradiso Massif (Western Alps): Implications for 40Ar-39Ar geochronology

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PRESERVATION OF SUB-MICROSCOPIC STRUCTURAL RELICS IN MICAS FROM THE GRAN PARADISO MASSIF (WESTERN ALPS): IMPLICATIONS FOR 40 AR-39 AR GEOCHRONOLOGY

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Marco Beltrando^{1*}, Gianfranco Di Vincenzo², Cristiano Ferraris³

1 Dipartimento di Scienze della Terra, Università di Torino, Via Valperga Caluso 35, 10125 Torino, Italy 2 Consiglio Nazionale delle Ricerche, Istituto di Geoscienze e Georisorse, Via Moruzzi 1, 56124 Pisa, Italy 3 Muséum National d'Histoire Naturelle, Laboratoire de Minéralogie et Cosmochimie, 61, Rue Buffon, 75005 Paris, France

*email: marco.beltrando@unito.it

Abstract 11

> Permian igneous biotite and white mica that were re-heated at 500-550°C in the Middle-Late Eocene were investigated by laser step-heating and in-situ ⁴⁰Ar-³⁹Ar techniques, in conjunction with scanning electron microscopy (SEM), electron microprobe (EMP) and transmission electron microscopy (TEM) to assess the influence of short-lived metamorphism on K-Ar systematics. Large intra- and inter-grain age variations, ranging from the Triassic to the Late Eocene, are primarily related to the extent of metamorphic reequilibration. Brown biotite porphyroclasts from weakly re-equilibrated samples are characterized by coreto-rim zoning in major element composition, mineral structure and ⁴⁰Ar/³⁹Ar ratios. Titanium concentration and Fe/(Mg+Fe) ratios decrease from crystal cores towards the rims, which are compositionally indistinguishable from green biotite aggregates in metamorphic coronas. Compositional changes are coupled with structural modifications towards the crystal edges, where the disordered igneous $1M-2M_1$ stacks are replaced topotactically by highly ordered metamorphic 1M politypes. Apparent ages up to ~66 Ma were determined in crystal cores by in situ laser-probe analyses, while ages as young as ~45 Ma are typical of crystal rims. Step-heating experiments of these zoned biotites yielded a discordant saddle-shaped age spectrum, with a concordant central segment yielding an error-weighted mean age of 44.9 ± 0.3 Ma. Biotite porphyroclasts from a more re-equilibrated specimen gave markedly different results, with uniform major elements compositional profiles, homogeneous argon distribution throughout the crystals and a flat age spectrum at 36.5 ± 0.3 Ma. TEM investigations of these optically homogeneous porphyroclasts revealed that the original igneous structure had been replaced by highly ordered 1M polytypes during Alpine metamorphism. White mica yielded comparable results, with spot ages up to ~218 Ma in domains characterized by high Na/(Na+K) ratios and low Si and Ti content, representing the best preserved igneous relicts. Significantly younger apparent ages, down to 44-39 Ma, are typical of domains affected by partial re-equilibration at high-pressure conditions, as indicated by increasing Si and Ti contents, or by exhumation-related exsolution of amoeboidal quartz. These results indicate that different intra-grain and inter-grain Ar concentration patterns arise primarily from the relative amount and distribution of relict igneous and metamorphic crystal-chemical domains. Core-to-rim age gradients in biotite result from inward-directed recrystallization from the crystal edges, eventually leading to flat age profiles as

metamorphic replacement of pre-existing porphyroclasts runs to completion. This study calls for caution when estimating thermal histories from apparent age gradients or from age vs. grain size correlations, unless the presence of pre-metamorphic mineral relics can be ruled out.

Keywords: ⁴⁰Ar-³⁹Ar geochronology, biotite, white mica, TEM, recrystallization

1. Introduction

Geochronological studies conducted in metamorphic terranes have repeatedly shown that Ar concentration within tri- and di-octahedral micas can be markedly heterogeneous. Published patterns include core-to-rim (e.g. Smith et al., 2005) or rim-to-core younging trends (e.g. Pickles et al., 1997) or, more commonly, complex age variations unrelated to crystal edges (e.g. Scaillet et al., 1992; Hodges et al., 1994; Scaillet, 1996; Hames and Cheney, 1997; Giorgis et al., 2000; Kramar et al., 2001; Sherlock and Kelley, 2002; Di Vincenzo et al., 2006; Schneider et al., 2008). Different processes have been proposed to account for the observed age spread, including: (1) thermally activated partial loss of radiogenic argon, dependent upon the distance from crystal edges (Dodson, 1973; Smith et al., 2005) or crystal defects (Hames & Cheney, 1997; Kramar et al., 2001; Mulch and Cosca, 2004; Hames et al., 2008; Cosca et al., 2011); (2) episodic influx of externally derived argon, possibly followed by subsequent diffusional loss (Arnaud & Kelley, 1995; Pickles et al., 1997, Giorgis et al., 2000; Sherlock & Kelley, 2002; Warren et al., 2011); (3) fluid-mediated recrystallization, leading to the local re-equilibration of mica porphyroclasts (Villa, 1998; Di Vincenzo et al., 2001 and 2006; Tartése et al., 2011).

Deciphering the relative contribution of these mechanisms to argon re-distribution within metamorphic minerals has fundamental bearings on the interpretation of ⁴⁰Ar-³⁹Ar data and on its use in geo- and thermo-chronology. This task may be complicated by the concomitant occurrence of more than one of the above mechanisms within most metamorphic rocks, which commonly undergo heating-cooling cycles associated with multiple stages of deformation and recrystallization. Therefore, a fair assessment of the mechanisms that controlled Ar re-distribution within a specific rock should include, alongside ⁴⁰Ar-³⁹Ar step heating and in-situ analyses, spatially-controlled mineral chemistry and mineral structural data. These different data sets have been combined only in a handful of studies on white mica (e.g. Giorgis et al., 2000; Di Vincenzo et al., 2001), phlogopite (Camacho et al., 2012) and K-feldspar (Villa and Hanchar, 2013).

The study presented here is aimed at investigating the relationship between K-Ar systematics, nano-structural, micro-structural and compositional modifications that affected igneous Permian micas during Tertiary Alpine metamorphism. While this study is mainly focused on biotite, white mica has also been analysed to provide a more complete picture of Ar mobility and availability within the system. The contribution of recrystallization to argon redistribution is assessed with a multi-disciplinary approach

combining laser step-heating and in-situ ⁴⁰Ar-³⁹Ar techniques with scanning electron microscopy (SEM), electron microprobe (EMP) and transmission electron microscopy (TEM) analyses.

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2. Geological Setting

The Gran Paradiso Massif crops out in the axial zone of the Western Alps (Fig. 1). It is usually grouped together with the Monte Rosa and Dora Maira Massifs under the common label of Internal Crystalline Massifs. It is tectonically overlain by the ophiolitic Eclogitic Piemonte Units and rests tectonically above the Money gneisses, attributed to the Briançonnais domain (Fig. 1). The Gran Paradiso Massif consists of orthogneiss originated from a porphyritic granite intruded at 275-265 Ma (Bertrand et al., 2005; Ring et al., 2005) at P=0.25-0.35 GPa (Gabudianu-Radulescu et al., 2011) into a Variscan basement consisting of metapelites and minor meta-mafic rocks (e.g. Callegari et al., 1969; Le Bayon and Ballevre, 2006). Following Mesozoic rifting, leading to the deposition of a locally preserved sedimentary cover (Elter, 1971), the Gran Paradiso Massif underwent Alpine eclogite-facies metamorphism (e.g. Compagnoni and Lombardo, 1974). Metamorphic pressures in excess of 1.8 GPa have been estimated in different lithologies, with local differences possibly reflecting the presence of several sub-units (Le Bayon et al., 2006; Gabudianu Radulescu et al., 2009; Gasco et al., 2010). The orthogneiss generally lacks any evidence of high-pressure mineral assemblages, apart from occasional jadeite (Saliot, 1973). Temperatures ≥550°C were reached during exhumation, at P=0.8-1.5 GPa (Le Bayon et al., 2006; Gasco et al., 2010). Pervasive re-equilibration under epidote-amphibolite facies conditions was accompanied by the formation of a sub-horizontal fabric, which is characteristically observed in the field. Syn-deformation shear zones in orthogneiss contain quartz, K-feldspar, plagioclase (An₀₋₅ and An₁₇₋₂₇), biotite, white mica, titanite \pm epidote \pm garnet \pm ilmenite (Menegon and Pennacchioni, 2010). Despite this multi-stage Alpine metamorphic evolution, large areas of the massif completely escaped fabric re-equilibration, preserving mineralogical and textural record of their pre-Alpine evolution (Callegari et al., 1968; Le Bayon and Ballevre, 2006; Menegon and Pennacchioni, 2010; Gabudianu Radulescu et al., 2011).

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3. Existing Geochronology

Alpine metamorphism in the Gran Paradiso Massif can be bracketed to the 48-33 Ma interval. The lower boundary at ~48-44 Ma is provided by tectonic events recorded in the Eclogitic Piemonte Units and in the Briançonnais domain (Fig. 2): the Piemonte ophiolites, originally located in a more distal part of the rifted margin (e.g. Elter, 1972; Beltrando et al., 2010a), underwent Alpine high-pressure metamorphism at 48-44 Ma (e.g. U-Pb on zircon Rubatto et al., 1998; ⁴⁰Ar-³⁹Ar on white mica, Beltrando et al., 2009; U-Pb on zircon, Beltrando et al., 2010b); the Briançonnais domain, which belonged to a more proximal part of the

European paleo-margin compared to the Internal Crystalline Massifs, instead, was still lying at the surface in the Lutetian (46-40 Ma), at attested by the deposition of flysch sediments (e.g. Stampfli et al., 1998). Furthermore, Zircon Fission Track ages in the 33-30 Ma range for the Gran Paradiso Massif (Hurford and Hunziker, 1989) and ~29 Ma for the southern Dora Maira Massif (Gebauer et al., 1997) provide a reliable upper limit for the metamorphic evolution (Fig. 2). Indeed, a large number of radioisotopic data for the different stages of Alpine metamorphism in the Internal Crystalline Massifs fall within the interval defined above, generally in the 43-33 Ma range (see Fig. 2 and Rosenbaum and Lister, 2005 and Beltrando et al., 2010a for reviews). High pressure metamorphism has been dated at 42.6 ± 0.6 Ma in the Monte Rosa Massif (U-Pb on rutile; Lapen et al., 2007) and at ~38-35 Ma in the southern Dora Maira Massif [U-Pb on zircon (Tilton et al., 1991; Gebauer et al., 1997), Sm–Nd and Lu-Hf on garnet (Tilton et al., 1991; Duchêne et al., 1997), U-Pb on titanite (Rubatto and Hermann, 2001), Rb–Sr on phengite-whole rock (Di Vincenzo et al., 2006)].

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The timing of the different stages of the tectonometamorphic evolution of the Gran Paradiso Massif is somewhat less well defined than in the Monte Rosa and Dora Maira Massifs. Contrasting estimates have been proposed, depending upon the technique and the rock type investigated (Fig. 2). Crystallization of monazite and allanite at high-pressure conditions in talc-chloritoid micaschists has been constrained at 37.4 ± 0.9 Ma and 33.7 ± 1.6 Ma, respectively (U-Pb; Gabudianu-Radulescu et al., 2009). These results are at odds with Rb-Sr estimates of 43.0 ± 0.5 Ma for high-pressure metamorphism in the same rock sample (phengite-apatite; Meffan Main et al., 2005), followed by retrogression at greenschist facies conditions after 36 Ma (phengite-paragonite and paragonite-albite; Meffan Main et al., 2005). Rb-Sr geochronological studies on the Gran Paradiso orthogneiss are broadly consistent with the picture arising from the talcchloritoid micaschists, as age estimates on deformed orthogneiss usually fall in the 42-33 Ma range (Inger and Ramsbotham, 1997; Rosenbaum et al., 2012). Shear zone activity at P≥1.0 GPa along the western boundary between the Gran Paradiso Massif and the Piemonte Units has been dated at 41.2 ± 1.1 Ma (Rb-Sr on white mica-epidote-apatite; Rosenbaum et al., 2012). Greenschist facies shear zones bounding the Gran Paradiso Massif to the north and to the east yielded younger estimates at 33.2 ± 0.4 Ma (Rb-Sr on white mica-feldspar; Freeman et al., 1997) and 39-34 Ma (Rb-Sr on white mica-feldspar-epidote; Inger and Ramsbotham, 1997). Importantly, pre-Alpine apparent ages have commonly been reported. Ages of 207.3 ± 2.7 Ma and >80 Ma were estimated in a meta-granitoid largely preserving the original mineral assemblage and texture (Rb-Sr on muscovite-feldspar; Inger and Ramsbotham, 1997) and in a sheared meta-granitoid (Rb-Sr on quartz-feldspar-white mica; Rosenbaum et al., 2012), respectively. As noted in both studies, different white mica grain fractions are characterized by contrasting Sr-isotopic signatures, with larger mica crystals yielding older apparent ages. This feature has been interpreted as evidence for isotopic inheritance and incomplete recrystallisation of the original igneous white mica during Alpine deformation/ metamorphism (Inger and Ramsbotham, 1997; Rosenbaum et al., 2012).

Similar to Rb-Sr data, ⁴⁰Ar-³⁹Ar studies on white mica also yielded a large age spread from ~160 to ~30 Ma (Fig. 2; Chopin and Maluski, 1980; Reddy et al., 2003). A narrow range has been found exclusively in phengite from a highly sheared granitoid, with step ages of 44-37 Ma from step-heating experiments (Chopin and Maluski, 1980) and in white mica + biotite aggregates from an epidote-amphibolite shear zone, yielding apparent ages in the 40-36 Ma range (Rosenbaum et al., 2012). More commonly, step-heating experiments resulted in highly discordant age spectra, with pre-Alpine apparent ages up to ~65 Ma in metagranites (Chopin and Maluski, 1980) and in the 120-40 Ma range in a sheared meta-pelite (Rosenbaum et al., 2012). Importantly, in the latter example coarser-grained mineral separates yielded older apparent ages. Furthermore, apparent ⁴⁰Ar-³⁹Ar ages in the 77-30 Ma range are characteristic of biotite-white mica aggregates from samples where isotopic inheritance was inferred from Rb-Sr data (samples GP03-163 and GP03-161 in Rosenbaum et al., 2012). In the Teleccio Lake area, where the samples that are the subject of this study were collected, several steps of previous step-heating experiments yielded apparent ages between ~80 and ~40 Ma (Rosenbaum et al., 2012). Differently from white mica, biotite has rarely been the focus of geochronological studies: apparent ages in the 31-27 Ma range have been reported for partly chloritized relics of magmatic biotite (Chopin and Maluski, 1980), while K-Ar dating of magmatic biotite yielded 45 ± 1 Ma (Krommenacher and Evernden, 1970).

In summary, syn-kinematic mineral assemblages from different rock types generally yielded apparent ages in the 44-33 Ma range (Fig. 2). However, largely undeformed meta-granites and, occasionally, sheared orthogneisses, yielded older ages, when analysed by both Rb-Sr and ⁴⁰Ar-³⁹Ar techniques. These ages were initially believed to indicate the timing of Alpine high pressure metamorphism (Chopin and Maluski, 1980), but have subsequently been tentatively related to an excess argon component (Reddy et al., 2003) or, alternatively, to the preservation of pre-Alpine mineral relicts partly re-equilibrated during the Alpine orogenic cycle (Inger and Ramsbotham, 1997; Rosenbaum et al., 2012). The study presented here aims at addressing this controversy, testing whether pre-Alpine ⁴⁰Ar-³⁹Ar ages may be related to incomplete resetting of the pre-existing mineral reservoirs in biotite and white mica.

4. Petrographic Description

Samples were collected along the banks of the Teleccio Lake. Details on the pre-Alpine and Alpine tectonometamorphic evolution of this small area, where Compagnoni and Lombardo (1974) reported the first evidence of Alpine eclogites in the Gran Paradiso Massif, can be found in an extensive body of literature (e.g. Compagnoni and Lombardo, 1974; Le Goff and Ballevre, 1990; Menegon and Pennacchioni, 2010; Gabudianu Radulescu et al., 2011; Rosenbaum et al., 2012). Two meta-granitoids, characterized by the best preserved igneous assemblage (JT1007) and by the partial static re-equilibration of the original igneous assemblage, without the formation of a new Alpine fabric (JT1008), were selected.

Sample JT1007 is a meta-granite, with a well preserved igneous assemblage consisting of quartz + Kfeldspar + plagioclase + biotite and accessory apatite, zircon and monazite (Fig. 3a and Supplementary Fig. 1). This sample is coarse grained, with sub-centimeter sized quartz and K-feldspar. The latter commonly displays Karlsbaab twinning and exsolutions along the cleavage planes. Igneous plagioclase, often preserving polysynthetic twinning, is generally zoned, with crystal rims of oligoclase composition and cores largely replaced by fine-grained aggregates consisting of albite + white mica + zoisite. Brown biotite (Bt1) is usually found in aggregates of millimeter to pluri-millimeter sized crystals, often associated with apatite and zircon inclusions, surrounded by the characteristic dark halo (Fig. 3b). The igneous mineral assemblage is partly re-equilibrated as a result of Alpine metamorphism, as indicated by coronitic mineral assemblages. Bt1 is frequently rimmed by thin fine-grained aggregates of white mica + ilmenite (Fig. 3c), locally replaced by a second rim of green biotite (Bt2) + titanite. Ilmenite is common both along the edges of Bt1 crystals and along the cleavage planes. When Bt2 aggregates are directly in contact with Bt1 porphyroclasts (Fig. 3b), igneous brown biotite displays a weak zoning towards the green-brown rims. Further evidence of Alpine re-equilibration include epidote rims around magmatic monazite, albite + white mica + zoisite aggregates formed at the expense of the cores of magmatic plagioclase, albite + white mica aggregates along the edges of magmatic K-feldspar and rare chlorite replacing biotite. Minor evidence of Alpine deformation is restricted to occasional kinking of the igneous biotite and to the local development of subgrains along the edges of large quartz crystals, which display weak wavy extinction (Fig. 3a).

Sample JT1008 is a two-mica tourmaline-bearing meta-aplite, with an original mineral assemblage consisting of medium grained K-feldspar + plagioclase + biotite + white mica + quartz, with accessory tourmaline, apatite and zircon (Supplementary Fig. 1). The main differences with respect to JT1007 concern the extent of Alpine re-equilibration. Pluri-millimeter sized magmatic quartz is rare, being generally replaced by medium- to fine-grained aggregates of quartz crystals with lobated margins and wavy extinction, often preserving the original crystallographic orientation of the parent crystal (Fig. 3d). Brown biotite relics are locally kinked and are commonly surrounded by coronas of ilmenite + medium sized white mica (Wm2; ~100 µm in diameter; Fig. 3e). Rare rutile is observed both associated with Wm2 and along the cleavage planes of igneous biotite. Ilmenite + Wm2 aggregates are then locally replaced by green biotite (Bt2) + titanite. Green biotite aggregates, occasionally hosting white mica (Wm3) are common around Bt1 (Fig. 3f). Small crystals of igneous biotite may be completely replaced by Bt2 aggregates. Chlorite is very rarely observed to replace small Bt1 crystals.

Igneous pluri-mm white mica crystals (Wm1), often kinked (Figs. 3f and 3g), are common in this sample. When in contact with igneous biotite, they are generally overgrown by rims consisting of ilmenite + Wm2 in optical continuity with Wm1 (Fig. 3f). Patchy re-equilibration of Wm1 is indicated by a number of petrographic observations, including crystallographically-unrelated variations in back-scattered electron (BSE) emission (Fig. 3g), due to different Ti contents (see below) and the presence of amoeboidal quartz

exsolutions elongated along cleavage planes (Fig. 3h). Furthermore, green biotite (Bt2) or ilmenite are observed both along kink walls and cleavage planes of large Wm1 crystals, which are also occasionally overgrown by albite (Fig. 3i).

5. Mineral chemistry

Mineral chemistry was determined with a JEOL JXA-8200 electron microprobe at the Dipartimento di Scienze della Terra, Università degli Studi di Milano (Italy). Conditions of analyses were 15 kV and 5 nA, and natural standards were employed. Spot sizes were 5 μ m for white mica and feldspars and 1 μ m for biotite. Representative analyses of biotite and white mica are listed in Table 1 and 2.

Igneous brown biotite and metamorphic green biotite from both samples differ mainly for the Ti content (Figs. 4 a, 4b and 4c). This difference is especially noticeable in JT1007 biotite, where Ti contents of 0.14-0.17 and 0.07-0.10 atoms per formula unit (a.p.f.u.) have been determined for Bt1 and Bt2, respectively (Fig. 4a). Other compositional differences include F content in JT1007, typically in the 0.02-0.07 a.p.f.u. range for Bt1, while it is invariably <0.02 a.p.f.u. in Bt2 (Fig. 4b) and the Mg/(Mg+Fe) ratio in JT1008, where Bt2 is slightly more Mg-rich than Bt1 (Fig. 4c). No significant variation in the Mg/(Mg+Fe) ratio has been detected in JT1007.

Compositional profiles across relict igneous biotite revealed important differences between the two samples (see Figs. 3b, 3e and Supplementary Fig. 2 for the location of the different profiles). Igneous biotite from JT1007 is typically zoned, with mineral composition of the crystal rims converging towards the composition of adjacent green biotite (Figs. 4a, 4b, 4d and 4e). The width of these domains, which are typically green-brown under the optical microscope (Fig. 3b), appears to correlate with the presence of Alpine mineral phases in the specific micro-site. Re-equilibrated areas extend for ~ 80 µm in Bt1 crystals surrounded by abundant Bt2 and/or white mica + ilmenite rims (Figs. 4d and 4e), while they are nearly absent in domains that escaped metamorphic re-equilibration at the microscale (Fig. 4f). Compositional profiles of biotites from JT1008 are instead remarkably homogeneous, even for domains directly adjacent to green biotite aggregates (Figs. 4g, 4h and 4i).

Igneous and metamorphic white mica can be distinguished based on marked compositional differences (Fig. 5). The paragonite component is significantly higher in Wm1 [Na/(Na+K)~0.1] than in Wm2 [Na/(Na+K)~0.02-0.05]. Furthermore, alpine white mica displays a higher celadonitic component, with silica content in the 3.10-3.22 a.p.f.u. range, as opposed to the values of 3.05-3.09 typical of the igneous white mica (Wm1). The compositional plots shown in Fig. 5 also reveal that Wm1 displays a continuous range of composition from Na-rich, Ti-poor end members, typical of large crystals, to Na-poor and Ti-rich compositions. The latter compositional end member is typical of smaller crystals of igneous white mica adjacent to igneous biotite. BSE images of large Wm1 crystals reveal a rather complex Ti zoning, which

appears to be primarily controlled by the presence of kink walls and the proximity of Bt1. Higher Ti contents (0.05-0.1 a.p.f.u.) are also characteristic of alpine white mica formed in optical continuity with igneous white mica at the expense of igneous biotite. These domains are also characterized by higher Si, Fe and Mg and lower Al than the original Wm1, as typically observed for the celadonitic substitution (Mg,Fe)^{VI}+Si^{IV} for Al^{VI}+Al^{IV}. A rather large compositional range is characteristic of alpine white mica (Fig. 5). Alpine white micas from JT1007 generally fall within the same compositional range as those analysed in JT1008, albeit towards the low-Ti end (Fig. 5).

6. Transmission Electron Microscopy

Biotite crystals from both samples were investigated by TEM (see the Supplementary Material for details of the analythical methods). The two samples yielded significantly different results, consistent with different extent of metamorphic re-equilibration of the two specimens during the Alpine orogeny. Biotite from sample JT1008 is generally homogeneous and mainly consists of a 1M polytype, with rare stacking faults (Figs. 6a and 6b). Minor chloritization of the 1M biotite is generally restricted to the crystal edges (Figs. 7a and 7b), probably indicating a late-stage metamorphic overprint. Biotite from sample JT1007 is structurally more complex than JT1008 biotite. Core-to-rim structural and chemical zoning is common. Crystal cores display several features that are indicative of crystallization under plutonic conditions (e.g. Fregola and Scandale, 2011), including the reflections with $k \neq 3n$ in Selected Area Electron Diffraction (SAED) patterns, showing streaking typical of stacking disorder (Fig. 6d). In this central portion the 1M polytype sequence is rare and sometimes appears alternated to the more common $2M_1$ polytype (Fig. 6c). No evidence of chloritization is found within these biotite relics. Crystal edges are statistically Ti- and Fe-poorer compared to nucleus relics, as also apparent from electron microprobe analyses. This compositional difference is accompanied by a major structural change, as peripheral areas of biotite porphyroclasts consist mainly of the 1M polytype. Chloritization is locally observed.

In summary, the structural analyses performed on biotite from both samples indicates that they were affected by Alpine metamorphic recrystallization to different extents. Relics of the original igneous structure are preserved exclusively within cores of the porphyroclasts from sample JT1007. Metamorphic recrystallization resulted in the complete replacement of the igneous structure in JT1008 and in the crystal rims of biotite from sample JT1007.

7. ⁴⁰Ar-³⁹Ar geochronology

7.1 Step-heating analyses

Details of the analythical methods for the 40 Ar- 39 Ar analyses are in the Supplementary Material. Infrared laser step-heating experiments on biotite separates from the two samples gave contrasting results (Figs. 7a and 7b): JT1008 yielded a flat spectrum, for more than 90% of the released 39 Ar_K, with an error-weighted mean age of 36.5 \pm 0.3 Ma; JT1007 yielded a slightly saddle-shaped profile, with initial steps in the 48-52 Ma range followed by a concordant central segment (75% of released 39 Ar_K), with an error-weighted mean age of 44.9 \pm 0.3 Ma. A slight increase of apparent ages to 46-48 Ma is again characteristic of the last 3 steps, which comprise <6% of the total released 39 Ar_K.

White mica from JT1008 yielded a highly discordant age profile, with most step ages falling in the 79 to 104 Ma range (Fig. 7c). Step 3 yielded the lowest apparent age of 63.4 ± 5.0 Ma, followed by a progressive increase to an age maximum of ~129 Ma (step 7). Steps 10 to 18, with the exception of step 12, are characterized by rather constant apparent ages, within the 83-76 Ma range, comprising >26 % of the total 39 Ar_K.

7.2 In-situ analyses on single grains

Millimeter-sized single crystal separates of JT1007 biotite (Figs. 8a and 8b) and JT1008 white mica (Figs. 8c, 8d and 8e) were also analyzed in situ, to directly image potential age gradients within individual crystals. All analyses were performed on the basal cleavage and they revealed significant intra-grain and inter-grain variations in apparent ages. The best example of intra-grain age variations is provided by a biotite crystal with a diameter >1mm, preserving the original igneous crystallographic shape (Fig. 8a). A broad core-to-rim age gradient is immediately apparent, with the oldest age of 66.1 ± 1.0 Ma in the crystal core and apparent ages consistently <50 Ma along the rims (minimum age is 45.2 ± 0.9 Ma). Importantly, core-to-rim age gradients are not constant, as revealed by a closer inspection of Fig. 8a. Another smaller crystal (Fig. 8b), affected by a larger extent of Alpine re-equilibration, as indicated by minor aggregates of white mica, is characterized by a narrower age spread, from 44.5 ± 0.6 Ma to 51.4 ± 0.7 Ma. Significantly, there is no correlation between apparent age and position within the crystal.

Large intra-grain and inter-grain variations in apparent ages are also typical of white micas (Figs. 8c, 8d and 8e). The relationship between apparent ages and position with respect to crystal core and rim is often difficult to assess, since original crystal edges can be determined only for euhedral crystals. However, being white mica probably a late-crystallizing mineral phase in the igneous protolith, the original crystal edges were often difficult to be detected in the mineral separate. Therefore, only crystallographically controlled crystal edges have been outlined in Figs. 8c, 8d and 8e. Despite these limitations, it is immediately apparent that broad core-to-rim age gradients are characteristic of the analyzed white micas.

Apparent ages older than 200 Ma were repeatedly obtained within crystal cores, which yielded ~218 Ma as the oldest age (Fig. 8c). The youngest apparent age of 53.6 ± 2.4 Ma was determined from a crystallographically controlled mineral edge (Fig. 8c). Neighboring sites, only a few hundred micrometers apart, normally yielded widely different apparent ages, with the notable exception of the crystal shown in Fig. 8e, where three analyses performed at relative distances of 0.3-0.5 mm yielded similar ages, in the 90-85 Ma range.

Therefore, laser in-situ dating on single crystal separates indicate that both JT1007 biotite and JT1008 white mica are characterized by a very heterogeneous distribution of ⁴⁰Ar. Although a general decrease of apparent ages toward the rim is observed, ⁴⁰Ar* concentration gradients cannot be considered concentric. Occasionally, single crystals yielded a narrower age spread, with most apparent ages comparable to those characteristic of the flat portion of the step-heating age profiles (cfr. Figs. 8b and 8e with Figs. 7a and 7c).

7.3 In-situ analyses on thick section

In-situ analyses on polished rock chips of sample JT1008 were performed to investigate both the spatial distribution of argon isotopes in biotite and the relationship between composition and apparent ages in white mica, which could not be assessed on the single crystal separates.

Ten Infrared laser spot analyses were performed on different parts of biotite relics (Bt1), including crystal core, rims and kinked domains. All ages are statistically indistinguishable, ranging from 34.5 ± 1.3 Ma to 37.0 ± 1.7 Ma (Figs. 8g and 8h and Supplementary Fig. 3), with an error-weighted mean age of 35.5 ± 0.5 Ma (MSWD=0.73) for the ten analyses. Two analyses, which yielded slightly older apparent ages (in italic in Table 3 and in Fig. 8) were excluded from the mean calculation because the laser pit also sampled white mica, as apparent from BSE imaging (Fig. 8g) and from the chemical correlation diagram (age vs. 38 Ar_{Cl}/ 39 Ar_K in Supplementary Fig. 3). Detection of contamination by white mica is based upon the its lower Cl content compared to biotite (Table 1 and 2), resulting in lower Cl/K ratios estimated from neutron-produced Ar isotopes (Supplementary Fig. 3).

Seven analyses were performed on a large white mica crystal, with a diameter of ~0.9 mm and with the basal cleavage oriented nearly parallel to the section surface (Fig. 8f and Supplementary Fig. 3). Internal inhomogeneity within this igneous relict is indicated by several observation on the BSE image, including (1) the presence of discrete bright linear domains, consisting of ilmenite along kink walls, with variable spacing in the range of 100-160 μ m, (2) the 'porous' appearance of specific domains, due to small quartz exsolutions (Figs. 3g and 3h) and (3) the marked brightness contrast between Ti-rich and Ti-poor domains. Note that the distribution of kink walls, quartz or Ti-enriched mica does not correlate in a simple way with the position within the crystal. No core-to-rim age gradients are found and apparent ages range widely from 128.2 \pm 3.3 to 41.7 \pm 2.0 Ma, with the two extremes located less than 50 μ m apart (Fig. 8f). More

significantly, the youngest age was estimated near the crystal core, but comparable apparent ages <50 Ma were also found near the crystal edges. Importantly, the two oldest ages (128.2 ± 3.3 and 79.3 ± 1.8 Ma) were obtained in the only sites that do not cross kink walls and display the lowest BSE emission (Fig. 8f).

Another Wm1 crystal, with the basal plane oriented roughly perpendicular to the section surface and located less than 1 mm from the larger crystal described above, was also analysed (Fig. 8g and 8h). Quartz exsolutions, alpine biotite and fine-grained alpine white micas are common within this microstructural domain (Figs. 8g and 8h). An apparent age of 43.8 ± 1.6 Ma was obtained from a domain with abundant quartz exsolutions and minor Alpine biotite (Fig. 8h). A trench spanning the transition area between the Wm1 crystal and a fine-grained aggregate of Wm2 + rare biotite yielded a younger apparent age of 40.8 ± 1.2 Ma. An even younger age of 34.8 ± 1.4 Ma was obtained for the fine-grained aggregate of white mica and biotite (Fig. 8h).

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8. Discussion

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8.1 Relationship between crystal-chemical reservoirs and apparent ages

Results from the present study document significant inter-grain and intra-grain variations of apparent ages in relict igneous biotite and white mica that underwent an Alpine metamorphic cycle, which culminated with eclogite facies metamorphism under conditions of P≥2.0 GPa and T=500-550°C (Fig. 8). Similar age variations in other metamorphic terranes have been ascribed to diffusional loss of radiogenic ⁴⁰Ar from pre-existing sheet silicates (e.g. Hodges et al., 1994; Kramar et al., 2001; Hames et al., 2008). However, in the samples studied here, different crystal-chemical domains can be distinguished within both biotite and white mica relics, based on a combination of petrographic observations (Fig. 3), EMP (Figs. 4 and 5) and TEM analyses (Fig. 6), hinting at a potential link between metamorphic re-equilibration and Ar age variations. The apparent age of the different crystal-chemical reservoirs obtained by in-situ laser analyses or step-heating experiment can be assessed either with chemical correlation diagrams (e.g. Villa et al., 2000) or with a tight compositional control on the spatial distribution of the reservoirs (e.g. Giorgis et al., 2000; Di Vincenzo et al., 2006). In the studied samples, the different generations of igneous and metamorphic biotite and white mica in sample JT1007 do not differ significantly and systematically in their Cl and Ca content (Table 1 and 2), thus preventing the use of chemical correlation diagrams. However, compositional profiles, calibrated with nano-structural analyses, provide a consistent picture as to the relative distribution of the two crystal-chemical sites. As a result, the relationships between intra-grain variations in ⁴⁰Ar^{*} concentrations and specific crystal-chemical reservoirs can be unravelled.

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8.1.1 Biotite

Two different crystal-chemical sites can be distinguished in porphyroclastic biotite. They are related to the original igneous structure, with disordered stacking of 2M₁ and 1M polytypes, and to metamorphic re-equilibration, resulting into a highly order structure consisting of 1M polytypes. While the igneous structure has been totally erased in JT1008 biotite, which consists entirely of 1M polytypes (Figs. 6a and 6b), biotite from JT1007 is zoned, with igneous nano-structures widely preserved and 1M polytypes restricted to crystal rims (Figs. 6c and 6d). Areas characterized by different polytypes display distinctive compositions, with higher Ti contents and lower Mg/(Mg+Fe) ratios typically associated with the igneous structure. A similar decrease in Ti content between relict igneous biotite and Alpine biotite has already been documented in other high-pressure orthogneisses (e.g. Biino and Compagnoni, 1992). Our TEM study provides new evidence supporting this earlier suggestion. EMP compositional profiles also display a marked Ti decrease and a slight Mg/(Mg+Fe) increase towards the crystal edges, generally restricted to the outermost 80-100 μm, providing a useful tool to assess the relative distribution of igneous structural relics and newly formed metamorphic biotite (Fig. 4). Importantly, slightly irregular compositional patterns are characteristic of biotite cores (Fig. 4), suggesting that smaller scale structural re-equilibration may also be partly affecting areas that are not directly located along the crystal edges. These deviations may be controlled, at least partly, by the presence of kink walls, which are commonly observed with the optical microscope (Supplementay Fig. 2).

Differently from sample JT1007, biotites from sample JT1008 underwent complete nano-structural re-equilibration during Alpine metamorphism, despite preserving the brown color and the crystal size of the former igneous biotite (Fig. 3e and Supplementary Fig. 2). Compositional profiles are invariably flat, with only minor compositional variations, possibly suggesting that more than one biotite generation may be present, albeit locally (Fig. 4). The different structural and compositional characteristics of the two samples are closely matched by the distribution of in situ laser probe Ar data (Table 3 and Fig. 8g) and by the shape of the step-heating age profiles (Fig. 7b). Biotite from JT1008, which is compositionally and structurally homogeneous as a result of Alpine metamorphism, yielded both a flat age spectrum with an error-weighted mean age of 36.5 ± 0.3 Ma (Fig. 7b) and homogeneous intra-grain Ar concentrations (Fig. 8g). Apparent ages determined by in-situ analyses are statistically indistinguishable, ranging between 34.5 ± 1.3 Ma and 37.0 ± 1.7 Ma (Fig. 9). Biotite from the better preserved JT1007 is instead characterized by a discordant age spectrum, with a significantly older error-weighted mean age of 44.9 ± 0.3 Ma (Fig. 7a) and by significant core-to-rim age variations (Fig. 8a). Apparent ages as old as ~66 Ma were estimated in a crystal core, with a progressive decrease to ~45 Ma at the crystal edge.

Therefore, the results from sample JT1007 and JT 1008 indicate that there is a strong spatial correlation between structural/compositional re-equilibration assessed by EMP and TEM analyses and apparent ⁴⁰Ar-³⁹Ar ages. Complete resetting of the argon clock was achieved only in the biotites from

JT1008, which underwent thorough nano-structural re-equilibration during Alpine metamorphism, despite preserving the brown color and the crystal size of the pre-existing igneous biotite. The age spread observed in the least re-equilibrated biotites, instead, is best explained as related to the coexistence of two different argon reservoirs tied to the two different structural and compositional sites identified within biotites. A larger relative contribution of the magmatic biotite relics results in geologically meaningless ages >60 Ma, while Eocene ages are related to a larger relative contribution of mineral reservoirs formed during Alpine metamorphism. Different degassing behavior of the two compositionally distinct crystal-chemical reservoirs may be responsible for the slightly discordant age spectrum, as widely documented for mixtures of different generations of white mica (e.g. Wijbrans and McDougall, 1986; Di Vincenzo et al., 2006; Beltrando et al., 2009), amphibole (e.g. Villa et al., 2000; Di Vincenzo and Palmeri, 2001) and possibly biotite (Berger, 1975; Maluski, 1978; Dallmeyer, 1982; York and Martinez, 1986).

8.1.2 White mica

Petrographic observations and electron microprobe analyses reveal that three main chemical reservoirs can be distinguished within relict igneous white mica. Compositional relicts of the original igneous muscovite, characterized by relatively high Na/(Na+K) and low BSE emissions are occasionally overgrown epitactically by phengitic mica at the contact with adjacent brown biotite (Figs. 3f and 3h). BSE imaging reveals that igneous muscovite hosts Si, Mg, Fe and Ti rich domains extending irregularly within the relict crystals. This compositional change, mainly controlled by the celadonitic substitution (Mg,Fe)^{VI}+Si^{IV} for Al^{VI}+Al^{IV}, is largely related to the presence of kinks or brown biotite nearby. Importantly, chemically reequilibrated domains are in optical continuity with the igneous muscovite, suggesting that phengitic substitution occurred topotactically. Subsequent re-equilibration is apparent in discrete domains hosting amoeboidal quartz, generally elongated along the basal plane of micas (Fig. 3h). Similar quartz exsolutions in white mica from meta-granitoids of the Dora Maira Massif originated from breakdown of Si-rich phengitic mica into quartz + Si-poor phengite, with rare talc detected exclusively with TEM observations (Ferraris et al., 2000). We suggest that the white micas from the Gran Paradiso Massif record a similar petrographic evolution. Therefore, quartz-rich domains within the sites of the original igneous mica are interpreted as areas that underwent early re-equilibration under HP conditions and subsequent exhumation-related re-equilibration.

As already noted above, the three different crystal-chemical sites are distributed irregularly within mica porphyroclasts. The apparent ages obtained with the in-situ laser probe technique on thick section also vary widely, but correlate with the crystal-chemical sites. The oldest apparent ages of ~128 Ma were obtained in areas preserving the least re-equilibrated igneous muscovite (Fig. 8f). Younger apparent ages are instead associated with Ti- and Si-enriched sites (Fig. 8f). The irregular boundaries between the different micro-chemical sites and the size of UV laser pits (300x100 µm) prevented sampling of single

crystal-chemical sites. However, the youngest apparent ages in the 47.0 ± 1.6 to 41.7 ± 2.0 Ma range were estimated from domains hosting the largest relative percentage of the amoeboidal quartz-rich mica. Younger ages of 40.8 ± 1.2 Ma and 34.8 ± 1.4 Ma were determined for domains characterized by variable amounts of fine-grained, newly formed white mica and biotite (Fig. 8h). Therefore, we suggest that the intra-grain age spread from ~128 Ma to ~42 Ma within white mica porphyroclasts resulted from variable mixing between a pre-Alpine argon reservoir, associated with the igneous relics, and two distinct Alpine reservoirs, related to HP metamorphism and subsequent decompression. Older apparent ages up to ~218 Ma have been obtained by in-situ analyses on single crystal separates (Fig. 8c). Therefore, this age can be considered as the minimum apparent age of the argon reservoir associated with the igneous mica relic, while the age of 41.7 ± 2.0 Ma provides an estimate for the maximum timing of formation of the youngest crystal-chemical reservoir, during decompression (Fig. 9a). Once again, the highly discordant spectrum resulting from the step-heating experiment on white mica separate (Fig. 7c) reflects the different degassing behavior of the three crystal-chemical reservoirs identified within relict white mica porphyroclasts.

8.2 Effects of partial argon pressure at the crystal edges

As discussed above, the studied biotites and white micas yielded apparent ages pre-dating the timing of Alpine metamorphism, despite having undergone re-heating at T≥550°C. Similar results from other metamorphic terranes have been alternatively ascribed to incomplete diffusional loss or to the episodic intake of extraneous argon from the rock matrix during the metamorphic cycle. However, a significant role of the latter process can be excluded for the Gran Paradiso samples, based on several lines of evidence. Firstly, occurrences of extraneous argon intake within pre-existing sheet silicates commonly result in a rim-to-core decrease of apparent ages (e.g. Pickles et al., 1997; Di Vincenzo et al., 2007; Warren et al., 2011), i.e. opposite to what has been documented for JT1007 biotite and JT1008 white mica. Furthermore, newly formed crystals are more prone to the intake of extraneous argon, as implied by metamorphic white mica and biotite yielding apparent ages older than pre-existing sheet silicates (Arnaud and Kelley, 1995). However, in the examples documented here, texturally younger white mica and biotite invariably yielded younger apparent ages with respect to the igneous relics.

It has been argued that unexpectedly old ages may also arise from non-zero argon concentration at the grain boundaries during metamorphism, slowing down diffusion of argon away from the crystal lattice (e.g. Wheeler, 1996; Kelley and Wartho, 2000). Partial argon pressure may result either from extraneous argon or from the progressive accumulation of radiogenic ⁴⁰Ar along the crystal edges, in the absence of fast fluid migration pathways (e.g. Kelley, 2002). The presence of non-zero Ar concentrations along the grain boundaries during specific stages of the tectonometamorphic evolution is very difficult to assess with independent evidence, as fluids can be consumed/produced through subsequent mineral reactions (e.g., Proyer, 2003) or removed from the system. However, these processes should affect minerals from a single

specimen equally, unless obvious heterogeities (i.e. veins) are present. However, the studied samples lack any evidence of metamorphic veining and white micas and biotites of similar grain size display very different apparent age distributions (Figs. 8a, 8b, 8c, 8d and 8e). This pattern is more compatible with locally variable metamorphic re-equilibration, as apparent from petrographic and compositional analyses. Furthermore, the homogeneous ages of JT1008 biotite suggest that, at least at the timing of biotite recrystallization, ⁴⁰Ar* stored within pre-existing white mica reservoirs was being removed efficiently from the system. Therefore, we conclude that the measured age gradients in white mica are not the result of local hindrance to diffusion of Ar away from the crystal lattice.

8.3 Argon re-distribution in biotite and white mica during short-lived metamorphism

Our study supports previous evidence from the Gran Paradiso Massif on the preservation of pre-Alpine radiogenic Sr and Ar in large igneous white mica relics in meta-granitoids (Inger and Ramsbotham, 1997; Rosenbaum et al., 2012). These previous suggestions, which can now be confidently extended to biotite, are here substantiated by detailed structural information by TEM analyses for biotite and by in-situ laser probe geochronology on both white mica and biotite. Therefore, argon loss from the studied igneous micas heated up to T≥550°C was primarily controlled by metamorphic re-equilibration, rather than volume diffusion. It is worth noting that the preservation of apparent ages up to ~216 Ma in white mica with a diameter of ~1.2 mm (Fig. 8c) from the Gran Paradiso Massif is not surprising, as it is in accordance with the argon diffusivity determined experimentally for white mica (Harrison et al., 2009). In principle, the observations presented in this study do not allow excluding that the observed age pattern in white mica resulted from the interplay of volume diffusion from the igneous reservoirs combined with episodic recrystallization.

On the other hand, the preservation of pre-Alpine ages in igneous biotite is incompatible with experimentally determined Ar diffusivity for biotite (Harrison et al., 1985; Grove and Harrison, 1996), which predicts complete diffusive loss from millimeter-sized crystal flakes at the metamorphic temperature reached by the Gran Paradiso Massif. Anomalously old apparent ages have already been reported in several studies (e.g. Villa and Puxeddu, 1994; Kelley et al., 1997; Kelley and Wartho, 2000; Maurel et al., 2003). These results have been alternatively interpreted as indicating that biotite is significantly more retentive than commonly assumed (Villa and Puxeddu, 1994; Maurel et al., 2003) or ascribed to high Ar concentration along the grain boundaries, slowing down diffusional loss from the mica crystal (Kelley et al., 1997; Kelley and Wartho, 2000), or attributed to <<1Myr heating-cooling cycles (Lister and Baldwin, 1996). In the case presented here, the pre-Alpine ages are specifically tied to igneous crystal-chemical relicts, while they are absent in Alpine crystal-chemical sites, supporting the view that Ar mobility in the studied biotite was lower than predicted by experimental studies. In principle the Cretaceous ages estimated for the crystal cores might alternatively result from mixing between igneous and metamorphic micro-

reservoirs, from partial argon loss from the igneous reservoir only, or from a combination of both processes. The different scales of observations adopted in this study, ranging from $\sim 0.01~\mu m^2$ for the nanostructural analysis by TEM, to 1-25 μm^2 for compositional analyses by EMP to 10000-30000 μm^2 for insitu geochronology prevent clarifying this latter aspect. Despite these limitations, the significant discrepancy between measured and predicted ages suggests caution when applying experimentally determined diffusion parameters to naturally heated biotite.

Most importantly, the study presented here suggests that, in the absence of detailed mineral structure data, great caution should be taken when estimating the thermal history of metamorphic rocks from ⁴⁰Ar* concentration profiles, under the assumption that Ar mobility is primarily controlled by volume diffusion. The study on relict biotite from sample JT1008 shows that flat age spectra and homogeneous Ar distribution, rather than resulting from complete argon loss during Alpine heating, arose from complete recrystallization of the pre-existing brown biotite crystal. Importantly, detection of this process may escape inspection by optical microscopy and, possibly, compositional determinations by EMP. On the other hand, apparent core-to-rim age variations of the kind detected in biotite from JT1007 may be the result of incomplete recrystallization, rather than reflect incomplete diffusional loss. Therefore, in specific circumstances 'diffusion profiles' might be an artifact due to variable mixing of two distinct sub-microscopic argon reservoirs.

The presence of compositionally distinct argon reservoirs, likely characterized by different degassing behavior during the step-heating experiments, is also responsible for the discordant age spectra resulting from the step-heating experiments on mineral separates (Figs. 7a and 7c). Saddle-shaped age spectra of the kind obtained from JT1007 biotite might, in principle, arise for the presence of extraneous argon (e.g. Harrison and McDougall, 1981) or the input of melt and fluid inclusions (Esser et al., 1997; Boven et al., 2001). The study presented here, instead, shows that an older gas reservoir, rather than being related to externally-derived ⁴⁰Ar, is microstructurally tied to gneous crystal-chemical relics.

Another important feature emerging from this study, which may deserve future research, is related to the broad correlation between crystal size, Alpine re-equilibration and apparent ages observed in specific micro-sites. Compositional analyses, combined with petrographic observations, showed that smaller crystals are generally more prone to undergo metamorphic re-equilibration (see the 'large' and 'small' white micas in Fig. 5) and therefore be affected by loss of previously accumulated ⁴⁰Ar* (cfr. Figs. 8f and 8h). This process results into a broad correlation between crystal size and apparent age. Similar relationships have been reported for pre-orogenic micas that underwent subduction/orogeny-related metamorphism (e.g. Goodwin and Renne, 1991; Lo and Onstott, 1995; Reddy et al., 1996). Our study suggests caution when using this type of data to estimate diffusion parameters in mica or cooling rates of metamorphic terranes, unless the presence of multiple reservoirs of the kind documented here can be

ruled out, as correlations between crystal size and apparent age may reflect different extents of metamorphic re-equilibration, rather than different diffusion length scales.

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8.4 Geological significance of the measured ages

Crystallization ages in the 44-33 Ma range have generally been provided for syn-kinematic mineral assemblages from the Gran Paradiso Massif (see 'Geological Setting'). However, the presence of undetected mineral relics of the kind documented here might skew Rb-Sr and ⁴⁰Ar-³⁹Ar data towards ages pre-dating the metamorphic/deformation event of interest. Biotite from sample JT1008 yielded very consistent results (Fig. 9) and an error-weighted mean age of 36.5 ± 0.3 Ma can be calculated from the step-heating experiment. Due to the compositional and structural homogeneity of the dated biotites, which have been completely transformed into 1M metamorphic polytypes, this age is interpreted to date the recrystallization event. Metamorphic biotite within the original igneous biotite site is compositionally very similar to green biotite found in the rock matrix, in equilibrium with low-Si white mica and titanite. Therefore, the age of 36.5 ± 0.3 Ma is interpreted to constrain post-HP re-equilibration, during exhumation. The lack of direct relationships with albite and zoisite, which are also found as late metamorphic minerals in JT1008, does not allow defining whether green biotite formed under greenschist facies conditions or during the previous epidote-amphibolite facies metamorphic stage. Despite these limitations, our estimates suggest that the Gran Paradiso Massif was at relatively low-P conditions at 36.5 ± 0.3 Ma. This estimate is in accordance with the zircon fission track ages of 33-30 Ma (Hurford and Hunziker, 1989), while it is barely within error (2 σ) with respect to the U-Pb age of 33.7 \pm 1.6 Ma for allanite provided by Gabudianu-Radulescu et al. (2009), which probably underestimates the timing of HP metamorphism. We note that the error-weighted mean ages of 36.5 \pm 0.3 and 35.5 \pm 0.5 Ma estimated from the step-heating experiment and the in-situ laser probe analyses, respectively, do not overlap within 2 σ errors. This slight difference may be explained by a larger extent of chloritization of the biotites that have been dated in situ with respect to those separated and hand-picked for the step-heating experiment. Indeed, minor chlorite interlayering has been imaged by TEM (Fig. 6). Any other specific geological inference from the geochronological data presented in this study should be taken with care, as all apparent ages from biotite in JT1007 and white mica in JT1008 should be considered as mixed ages.

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9. Conclusions

This multi-disciplinary investigation of white mica and biotite shows that pre-metamorphic K-bearing minerals might preserve at least part of their radiogenic argon despite having undergone re-heating at T≥550°C. In this context, different patterns of intra-grain variations in Ar concentration may result primarily from the relative distribution of relict and metamorphic crystal chemical sites. Progressive recrystallization

from the edges towards crystal cores may result in core-to-rim age gradients within single crystals. Complete replacement of pre-existing porphyroclasts during metamorphism, instead, may lead to flat compositional profiles. Irregular recrystallization, partly unrelated to crystallographic edges, which is more likely in deformed and/or larger crystals, results in patchy age patterns. None of these age patterns should be interpreted as resulting from volume diffusion alone. Therefore, in metamorphic rocks, care should be taken when extrapolating thermal histories from argon data, unless the presence of pre-metamorphic mineral relics can be ruled out.

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phengitic mica + ilmenite or green-brown biotite and muscovite (e). Phengitic mica also crystallizes epitactically around igneous white mica (Wm1), at the expense of brown biotite (f). Igneous white mica is variably recrystallized, depending upon the presence of kink walls and the crystal size (g, h; note that the brightness contrast is related to variable Ti-content. Dashed orange line marks areas with abundant quartz exsolutions. See text for details). Late-stage albite is locally forming at the expense of white mica (i)

Fig. 4: mineral chemistry data from the different biotite generations in JT1007 and JT1008. Compositional profiles of brown biotite porphyroclasts are shown in (d), (e), (f) for sample JT1007 and (g), (h), (i) for sample JT1008. Analysed crystals are shown in Fig. 3 and Supplementary Fig. 2.

Fig. 5: mineral chemistry data for the different white mica generations in JT1007 and JT1008.

Fig. 6: (a) Low magnification TEM image of sample JT1008. This sample is mainly a 1*M* polytype as for the SAED in the figure inset. Minor chloritization is indicated by the brighter areas parallel to the stacking, perpendicular to c*. This chlorite layers are more evident in (b) where a TEM high resolution image shows how 14 Å thick chlorite (chl) layers enter the 10 Å biotite structure. (c) [110] HRTEM image of disordered biotite (Bt1) found within JT1007 cores. The stacking vectors (0, +, -) are indicated (Baronnet 1992, 1997; Baronnet et al., 1993). The presence of disordered biotite polytypes is confirmed by streaks in the SAED patterns (d).

Fig. 7: age release spectra of biotite from sample JT1007 and biotite and white mica from sample JT1008.

Fig. 8: in-situ age data. Line drawings show the distribution of laser-spot analyses in single crystal separates of JT1007 biotite (a, b) and JT1008 white mica (c, d, e). Red lines indicate crystallographically controlled mineral edges in white mica. SEM images and laser pits in the thick section of JT1008 (f, g, h). Yellow rectangles indicate UV pits, while brown lines define the edges of multiple IR shots. Dashed brown lines indicate IR pits that have been excluded from the calculation of the error-weighted mean age for biotite due to contamination by white mica, as apparent both from the SEM image and the measured ³⁸Ar (CI) content (see Supplementary Fig. 3). Analysed sites falling outside the field of view can be seen on Supplementary Fig. 3.

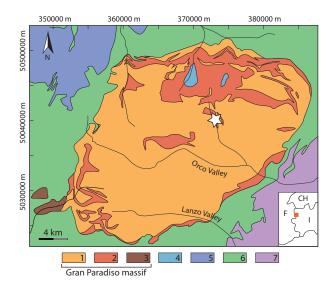
Fig. 9: synthesis of the geochronological estimates from white mica and biotite. Steps of the step-heating experiments are plotted individually. The timing of crystallization of the igneous protolith is taken from Bertrand et al. (2005) and Ring et al. (2005). Estimates of high pressure metamorphism in the Piemonte

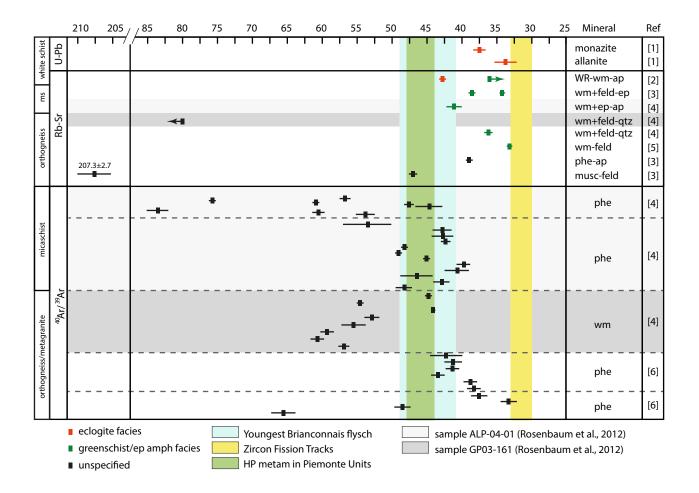
8/5	units is from Rubatto et al. (1998), Beitrando et al. (2009), Beitrando et al. (2010b). Zircon Fission Track
876	data from the Gran Paradiso Massif is from Hurford and Hunziker (1989)
877	
878	Table 1: representative electron microprobe analyses of the studied biotites.
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880	Table 2: representative electron microprobe analyses of the studied white micas.
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882	Table 3: in situ and step-heating ⁴⁰ Ar- ³⁹ Ar data. Argon isotope concentrations are in V. IR= infrared laser;
883	UV= ultraviolet laser. Steps of the step-heating experiments used to calculate the error-weighted mean age
884	are in bold characters.
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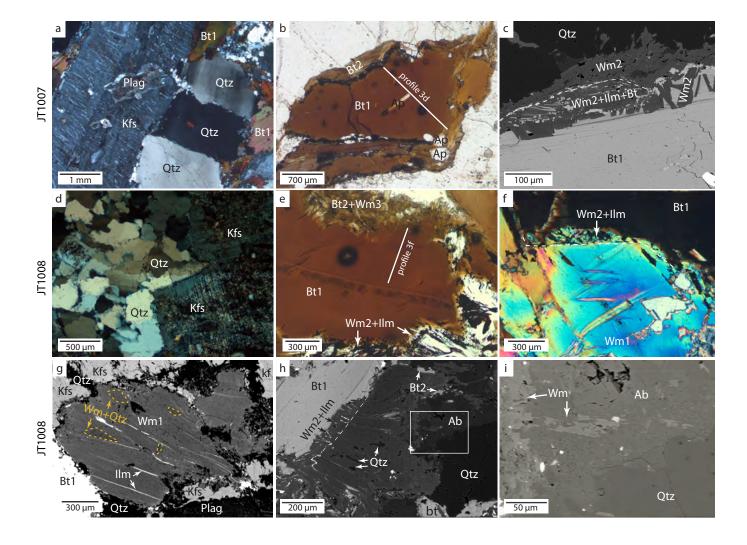
Sample	JT1007	JT1007	JT1007	JT1007	JT1007	JT1007	JT1007	JT1007	JT1007	JT1007	JT1007	JT1007	JT1008							
mineral	Bt 1 core	Bt 1 core	Bt 1 core	Bt 1 core	Bt 1 rim	Bt 1 rim	Bt 1 rim	Bt 1 rim	Bt 2	Bt 2	Bt 2	Bt 2	Bt 1	Bt 1	Bt 1	Bt 1	Bt 2	Bt 2	Bt 2	Bt 2
SiO ₂	34.80	34.89	34.82	34.39	34.63	35.03	35.29	34.96	35.56	35.33	35.52	35.09	35.09	35.00	35.53	36.05	36.55	34.97	35.68	32.53
TiO ₂	2.85	2.92	2.67	2.53	2.15	1.57	1.98	2.09	1.56	1.27	1.34	1.68	2.80	2.70	2.60	3.02	2.44	2.70	1.69	12.44
Al ₂ O ₃	17.79	17.89	18.05	17.34	17.23	17.98	17.94	18.02	18.16	17.84	18.58	18.28	18.45	18.56	18.35	17.90	18.32	18.17	17.57	13.22
FeO	25.74	25.96	25.76	25.75	27.10	25.23	25.81	26.35	26.18	25.80	25.85	26.36	24.54	24.45	24.51	24.00	23.21	24.77	25.15	15.64
MnO	0.22	0.25	0.20	0.24	0.32	0.18	0.23	0.21	0.25	0.23	0.20	0.26	0.35	0.32	0.33	0.31	0.33	0.37	0.34	0.09
MgO	5.30	5.74	5.66	5.35	5.73	5.60	5.51	5.65	5.98	6.15	5.58	5.57	6.34	6.37	6.59	6.56	6.54	6.26	7.57	4.16
CaO	0.05	0.01	0.02	0.06	0.02	0.11	0.04	0.03	0.01	0.04	0.04	0.13	0.05	0.05	0.09	0.05	0.03	0.04	0.01	13.17
Na ₂ O	0.10	0.11	0.12	0.98	0.02	1.06	0.05	0.03	0.04	0.02	0.05	0.05	0.11	0.10	0.07	0.08	0.03	0.03	0.00	0.06
K ₂ O	9.47	9.47	9.54	9.51	9.59	9.51	9.35	9.57	9.61	9.58	9.33	9.11	9.34	9.45	9.34	9.45	9.64	9.55	8.92	5.18
CI	0.04	0.05	0.03	0.46	0.01	0.58	0.05	0.04	0.03	0.03	0.04	0.05	0.04	0.03	0.03	0.04	0.05	0.05	0.07	0.05
F	0.21	0.36	0.44	0.26	0.34	0.16	0.29	0.35	0.11	0.14	0.10	0.10	0.39	0.26	0.35	0.28	0.35	0.38	0.34	1.10
total	96.56	97.81	97.55	96.87	97.17	97.01	96.54	97.31	97.49	96.43	96.63	96.67	97.49	97.29	97.80	97.74	97.49	97.28	97.36	97.72
Si	2.71	2.69	2.70	2.70	2.71	2.73	2.75	2.71	2.74	2.75	2.75	2.73	2.69	2.69	2.71	2.74	2.78	2.70	2.74	2.49
Ti	0.17	0.17	0.16	0.15	0.13	0.09	0.12	0.12	0.09	0.07	0.08	0.10	0.16	0.16	0.15	0.17	0.14	0.16	0.10	0.71
Al	1.63	1.63	1.65	1.60	1.59	1.65	1.64	1.65	1.65	1.64	1.70	1.67	1.67	1.68	1.65	1.60	1.64	1.65	1.59	1.19
Fe	1.68	1.68	1.67	1.69	1.77	1.65	1.68	1.71	1.69	1.68	1.68	1.71	1.58	1.57	1.57	1.53	1.48	1.60	1.62	1.00
Mn	0.01	0.02	0.01	0.02	0.02	0.01	0.01	0.01	0.02	0.02	0.01	0.02	0.02	0.02	0.02	0.02	0.02	0.02	0.02	0.01
Mg	0.62	0.66	0.65	0.63	0.67	0.65	0.64	0.65	0.69	0.71	0.64	0.65	0.73	0.73	0.75	0.74	0.74	0.72	0.87	0.47
Ca	0.00	0.00	0.00	0.00	0.00	0.01	0.00	0.00	0.00	0.00	0.00	0.01	0.00	0.00	0.01	0.00	0.00	0.00	0.00	1.08
Na	0.01	0.02	0.02	0.15	0.00	0.16	0.01	0.01	0.01	0.00	0.01	0.01	0.02	0.01	0.01	0.01	0.00	0.00	0.00	0.01
К	0.94	0.93	0.94	0.95	0.96	0.95	0.93	0.95	0.94	0.95	0.92	0.90	0.91	0.93	0.91	0.92	0.93	0.94	0.87	0.50
Cl	0.00	0.00	0.00	0.04	0.00	0.05	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.01	0.00
F	0.03	0.05	0.06	0.04	0.05	0.02	0.04	0.05	0.01	0.02	0.01	0.01	0.05	0.03	0.05	0.04	0.05	0.05	0.04	0.14
Sum cations	7.78	7.80	7.80	7.90	7.85	7.90	7.78	7.82	7.82	7.83	7.79	7.79	7.78	7.79	7.77	7.75	7.73	7.79	7.81	7.46
Mg/(Mg+Fe)	0.27	0.28	0.28	0.27	0.27	0.28	0.28	0.28	0.29	0.30	0.28	0.27	0.32	0.32	0.32	0.33	0.33	0.31	0.35	0.32
CI/K	0.00	0.00	0.00	0.04	0.00	0.06	0.01	0.00	0.00	0.00	0.00	0.00	0.06	0.04	0.05	0.04	0.05	0.05	0.05	0.29

Sample	JT1007	JT1007	JT1007	JT1007	JT1008	JT1008	JT1008	JT1008	JT1008	JT1008	JT1008						
Mineral	Alpine	Alpine	Alpine	Alpine	large	large	large	large	small	small	epitactic	epitactic	Wm2 on	Wm2 on	Wm2 on		
	Wm	Wm	Wm	Wm	Wm1	Wm1	Wm1	Wm1	Wm1	Wm1	Wm2	Wm2	Bt1	Bt1	Bt1	Wm3	Wm3
SiO ₂	47.58	47.60	47.53	46.91	46.64	46.43	46.57	46.59	45.97	46.15	46.93	47.48	47.91	48.05	46.76	46.40	47.09
TiO ₂	0.32	0.38	0.52	0.25	0.57	0.58	0.48	0.60	0.79	2.30	1.12	0.71	2.17	1.77	1.97	0.97	1.80
Al2O ₃	32.47	32.40	33.14	34.23	36.26	36.29	36.19	36.05	35.39	33.61	33.61	32.98	29.07	31.28	31.81	33.98	31.96
FeO	2.94	3.03	2.61	2.36	1.07	1.09	1.01	1.28	1.26	1.75	1.72	2.10	3.66	2.22	2.71	2.69	3.13
MnO	0.00	0.01	0.01	0.04	0.00	0.08	0.00	0.00	0.04	0.01	0.09	0.05	0.07	0.01	0.00	0.01	0.07
MgO	1.22	1.22	1.22	0.86	0.50	0.55	0.52	0.61	0.56	0.75	1.04	1.17	1.87	1.44	1.25	0.89	1.26
CaO	0.04	0.03	0.04	0.04	0.05	0.01	0.03	0.07	0.03	0.02	0.04	0.08	0.03	0.01	0.01	0.02	0.01
Na ₂ O	0.20	0.24	0.27	0.26	0.80	0.84	0.45	0.35	0.40	0.28	0.29	0.27	0.11	0.22	0.23	0.21	0.25
K ₂ O	10.98	11.09	10.98	10.92	10.20	10.10	10.70	10.78	10.90	10.92	10.79	10.73	10.85	10.99	10.96	11.00	10.89
Cl	0.00	0.01	0.00	0.01	0.00	0.01	0.00	0.00	0.00	0.01	0.00	0.00	0.00	0.00	0.00	0.01	0.01
F	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00
Total	95.74	96.01	96.33	95.89	96.09	95.97	95.95	96.34	95.33	95.80	95.62	95.58	95.75	95.99	95.70	96.18	96.47
Si	3.18	3.18	3.15	3.12	3.07	3.06	3.07	3.07	3.07	3.08	3.12	3.16	3.22	3.20	3.14	3.09	3.14
Ti	0.02	0.02	0.03	0.01	0.03	0.03	0.02	0.03	0.04	0.12	0.06	0.04	0.11	0.09	0.10	0.05	0.09
Al	2.56	2.55	2.59	2.68	2.81	2.82	2.81	2.80	2.78	2.64	2.64	2.59	2.31	2.45	2.51	2.67	2.51
Fe	0.16	0.17	0.14	0.13	0.06	0.06	0.06	0.07	0.07	0.10	0.10	0.12	0.21	0.12	0.15	0.15	0.17
Mn	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.01	0.00	0.00	0.00	0.00	0.00	0.00
Mg	0.12	0.12	0.12	0.09	0.05	0.05	0.05	0.06	0.06	0.07	0.10	0.12	0.19	0.14	0.12	0.09	0.13
Ca	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.01	0.00	0.00	0.00	0.01	0.00	0.00	0.00	0.00	0.00
Na	0.03	0.03	0.03	0.03	0.10	0.11	0.06	0.04	0.05	0.04	0.04	0.03	0.01	0.03	0.03	0.03	0.03
K	0.94	0.94	0.93	0.93	0.86	0.85	0.90	0.91	0.93	0.93	0.92	0.91	0.93	0.93	0.94	0.93	0.93
Cl	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00
F	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00
Sum cations	7.01	7.02	7.01	7.00	6.98	6.98	6.98	6.98	6.99	6.97	6.98	6.98	6.99	6.97	6.99	7.01	7.00
Mg/(Mg+Fe)	0.42	0.42	0.45	0.39	0.45	0.47	0.48	0.46	0.44	0.43	0.52	0.50	0.48	0.54	0.45	0.37	0.42
Na/(Na+K)	0.03	0.03	0.04	0.04	0.11	0.11	0.06	0.05	0.05	0.04	0.04	0.04	0.02	0.03	0.03	0.03	0.03

No.	mineral	laser	IR#	³⁶ Ar _(atm)	±2σ	37Ar _(Ca)	±2σ	38Ar _(CI)	±2σ	³⁹ Ar _(K)	±2σ	⁴⁰ Ar _(Tot)	±2σ	Age	±2σ	⁴⁰ Ar* %	type
			shots UV (µm²)	,,		(,		(,				(/		(Ma)			
121A	5482±0.0000344 Bt	IR	2	3.5E-05	4.0E-05	5.9E-04	8.4E-04	3.3E-04	7.6E-05	3.9E-02	2.4E-04	1.5E-01	5.8E-04	35.9	3.0	92.9	thick section A
121C	Bt+Wm	IR	2	2.6E-05	4.3E-05	-		4.1E-04	9.9E-05	6.3E-02	3.6E-04	2.8E-01	1.2E-03	43.3	2.0	97.1	thick section A
121D 121E	Bt+Wm	IR IR	3 2	1.7E-04 2.3E-05	4.0E-05 4.0E-05	4.9E-04 2.2E-04	8.8E-04 9.2E-04	4.6E-04 5.8E-04	9.8E-05 1.3E-04	4.8E-02 6.5E-02	2.3E-04 3.6E-04	2.5E-01 2.4E-01	6.6E-04 5.7E-04	40.3 35.6	2.4 1.8	79.5 97.0	thick section A
121G	Bt Bt	IR IR	4	1.7E-05	3.8E-05	2.4E-04	7.6E-04	1.0E-03	1.3E-04 1.9E-04	1.0E-01	4.9E-04	3.8E-01	8.0E-04	35.9	1.5	98.4	thick section A thick section A
121H	Bt	IR	4	5.4E-05	3.5E-05	2.4E-03	1.2E-03	7.9E-04	1.5E-04	8.5E-02	4.6E-04	3.2E-01	7.7E-04	35.7	1.2	94.8	thick section A
121I 144B	Bt Bt	IR IR	3 2	3.5E-05 1.9E-04	3.2E-05 4.5E-05	1.2E-04 5.0E-04	7.1E-04 2.7E-03	6.8E-04 4.5E-04	1.3E-04 9.3E-05	7.3E-02 5.2E-02	4.0E-04 4.3E-04	2.6E-01 2.4E-01	7.1E-04 3.1E-03	34.5 35.5	1.3 2.6	95.8 76.2	thick section A thick section B
144C	Bt	IR	4	1.8E-04	2.9E-05	8.5E-04	2.6E-03	4.3E-04	8.8E-05	5.3E-02	3.2E-04	2.4E-01	3.2E-03	35.1	1.7	77.5	thick section B
144D	Bt	IR IR	4	1.0E-04	2.8E-05	1.9E-03	3.3E-03	6.4E-04	1.4E-04	6.0E-02	4.2E-04	2.4E-01	3.2E-03	35.1	1.5	87.4	thick section B
144F 144G	Bt Bt	IR IR	3	3.0E-05 2.1E-05	3.1E-05 2.9E-05	1.1E-03 1.6E-03	3.1E-03 2.8E-03	3.4E-04 4.8E-04	8.3E-05 1.1E-04	3.9E-02 5.3E-02	3.1E-04 2.9E-04	1.5E-01 2.0E-01	3.1E-03 3.1E-03	35.1 37.0	2.4 1.7	93.9 96.8	thick section B thick section B
121N	Bt+Wm	UV	100x300	1.8E-04	3.0E-05	2.2E-04	8.2E-04	3.3E-04	8.7E-05	6.0E-02	3.8E-04	2.7E-01	1.3E-03	36.7	1.5	80.5	thick section A
1210 1210	Wm Wm2+Bt2	UV	100x300 100x300	2.0E-05 6.4E-05	3.3E-05 2.7E-05	7.1E-05 4.6E-04	1.0E-03 1.2E-03	1.3E-04 3.0E-05	7.1E-05 6.9E-05	6.0E-02 6.4F-02	3.0E-04 3.7E-04	2.7E-01 2.8E-01	1.1E-03 1.1E-03	43.8 40.8	1.6	97.6 93.1	thick section A thick section A
121R	Wm2+Bt2	UV	100x300	4.3E-05	5.6E-05	5.0E-04	9.6E-04	6.4E-05	6.5E-05	6.7E-02	3.8E-04	2.5E-01	1.0E-03	34.8	2.4	94.6	thick section A
121U	Wm	UV	100x300	7.8E-05	4.4E-05	2.3E-04	7.6E-04	3.0E-05	5.9E-05	3.8E-02	3.0E-04	5.3E-01	1.2E-03	128.2	3.3	95.6	thick section A
121V 121W	Wm Wm	UV UV	150x200 100x300	5.0E-05 1.1E-04	6.3E-05 3.3E-05	1.9E-04 6.4E-05	7.8E-04 6.6E-04	0.0E+00 4.7E-05	0.0E+00 5.9E-05	4.6E-02 4.7E-02	2.8E-04 2.5E-04	2.3E-01 2.3E-01	5.8E-04 7.1E-04	46.1 41.7	4.0 2.0	93.4 85.8	thick section A thick section A
121X	Wm	UV	100x300	8.6E-05	3.4E-05	1.2E-04	7.2E-04	4.5E-05	6.7E-05	5.6E-02	3.7E-04	4.3E-01	7.8E-04	69.6	1.8	93.9	thick section A
121Y 121Z	Wm Wm	UV	100x300 100x270	6.0E-05 8.2E-04	3.0E-05 3.2E-05	1.7E-04 3.4E-04	1.0E-03 8.9E-04	2.5E-05 0.0E+00	7.6E-05 0.0E+00	5.7E-02 5.4E-02	4.2E-04 3.5E-04	2.9E-01 6.8E-01	1.0E-03 1.4E-03	47.0 79.3	1.6 1.8	93.7 64.2	thick section A thick section A
122A	Wm	UV	100x270	2.3E-05	3.9E-05	3.8E-04	8.6E-04	1.3E-05	6.6E-05	4.3E-02	3.8E-04	2.3E-01	9.8E-04	50.1	2.6	96.8	thick section A
JT1008. J=0.002	7175±0.0000174																
12M0019B	Wm	CO2	1	9.6E-04	2.9E+00	2.5E-03	1.9E+03	7.5E-06	5.5E+02	8.9E-02	2.5E-01	1.6E+00	5.1E-01	85.5	1.0	84.9	single crystal separate
12M0019C	Wm	CO2 CO2	1	6.1E-04	2.5E+00 6.4E+00	2.5E-03 -2.8E-03	1.9E+03	5.1E-05 0.0E+00	1.1E+02 0.0E+00	1.1E-01 8.1E-02	2.9E-01	2.0E+00	2.3E-01	86.7	0.6	91.7	single crystal separate
12M0019D 12M0019G	Wm Wm	CO2	1	2.4E-04 2.5E-04	7.0E+00	9.3E-03	1.7E+03 3.3E+02	9.5E-05	6.1E+01	1.0E-01	2.1E-01 3.4E-01	1.7E+00 1.9E+00	2.7E-01 2.8E-01	102.7	0.7	96.1	single crystal separate
12M0019G	Wm	CO2	1	6.4E-04	3.1E+00	1.4E-02	2.3E+02	0.0E+00	0.0E+00	4.8E-01	1.7E-01	8.2E+00	7.9E-02	89.6 82.2	0.8	96.2 97.7	single crystal separate single crystal separate
12M0027A	Wm	CO2	1	3.1E-04	6.7E+00	5.1E-03	7.7E+02	5.3E-06	4.4E+02	4.5E-02	2.4E-01	1.9E+00	3.2E-01	199.5	1.5	95.5	single crystal separate
12M0027B	Wm	CO2	1	2.7E-04	6.8E+00	4.5E-03	1.0E+03	0.0E+00	0.0E+00	2.8E-02	4.0E-01	1.3E+00	4.2E-01	217.8	2.4	94.4	single crystal separate
12M0027C	Wm	CO2	1	4.4E-05	2.7E+01	5.3E-03	5.6E+02	1.0E-05	1.9E+02	1.5E-02	5.4E-01	1.6E-01	2.2E+00	53.6	2.3	92.6	single crystal separate
12M0027E	Wm Wm	CO2 CO2	1	5.3E-04 1.3E-04	4.0E+00 1.1E+01	6.9E-03 5.3E-03	4.3E+02 6.5E+02	3.2E-05 2.8E-05	8.5E+01 1.0E+02	5.1E-02 6.2E-02	3.5E-01 2.3E-01	2.4E+00 1.2E+00	2.8E-01	215.3	1.8	93.7	single crystal separate
12M0027F 12M0027G	Wm Wm	CO2 CO2	1	1.3E-04 2.6E-04	1.1E+01 5.8E+00	5.3E-03 -2.2E-03	6.5E+02 1.6E+03	2.8E-05 4.9E-06	1.0E+02 6.7E+02	6.2E-02 5.6E-02	2.3E-01 3.2E-01	1.2E+00 2.1E+00	3.6E-01 2.6E-01	93.6 171.6	0.8	96.8	single crystal separate
12M0027G	Wm	CO2	1	9.6E-04	2.4E+00	-1.5E-02	2.4E+02	1.6E-05	3.3E+02	1.2E-01	2.4E-01	2.7E+00	2.7E-01	171.6	1.3 0.7	96.4 90.3	single crystal separate single crystal separate
12M0027K	Wm	CO2	1	1.1E-03	2.0E+00	2.3E-03	1.3E+03	5.7E-05	3.7E+02	5.5E-01	2.6E-01	1.6E+01	2.4E-01	133.9	0.7	97.9	single crystal separate
12M0028A	Wm	CO2	1	1.1E-04	1.0E+01	6.2E-03	4.2E+02	3.2E-05	7.8E+01	5.8E-02	2.2E-01	2.6E+00	1.4E-01	204.5	1.0	98.7	single crystal separate
12M0028B	Wm	CO2	1	2.8E-04	8.6E+00	1.4E-02	2.5E+02	5.2E-05	1.0E+02	1.3E-01	2.7E-01	3.4E+00	2.5E-01	128.9	0.9	97.6	single crystal separate
12M0028C	Wm	CO2	1	5.0E-05	2.4E+01	4.6E-03	6.2E+02	0.0E+00	0.0E+00	2.3E-02	5.6E-01	8.4E-01	4.4E-01	171.8	2.3	98.2	single crystal separate
12M0028D 12M0028F	Wm Wm	CO2 CO2	1	4.1E-05 1.5E-04	3.6E+01 8.4E+00	-4.1E-03 -1.3E-02	7.4E+02 1.8E+02	0.0E+00 0.0E+00	0.0E+00 0.0E+00	6.1E-02 6.0E-02	5.1E-01 2.7E-01	1.5E+00 1.4E+00	2.9E-01 2.8E-01	120.3 109.0	1.4 0.8	99.2 96.9	single crystal separate
12M0028F	Wm	CO2	1	6.9E-04	3.9E+00	1.0E-02	4.3E+02	2.9E-04	7.8E+01	5.7E-01	1.6E-01	1.7E+01	6.6E-02	141.6	0.5	98.8	single crystal separate single crystal separate
	7302±0.0000164																3 ,
11M0124B	Bt	IR (12W)	3	5.2E-05	2.5E+01	7.4E-03	5.6E+01	1.3E-03	2.1E+00	6.7E-02	3.4E-01	6.2E-01	1.6E-01	44.5	0.6	97.5	single crystal separate
11M0124C	Bt	IR (12W)	5	3.5E-05	3.5E+01	1.2E-03	4.4E+02	1.0E-03	2.3E+00	4.8E-02	3.2E-01	4.7E-01	1.1E-01	46.5	0.8	97.7	single crystal separate
11M0124E 11M0124F	Bt Bt	IR (12W) IR (12W)	3 5	4.0E-05 4.3E-05	2.8E+01 2.8E+01	6.3E-03 4.9E-03	8.6E+01 8.8E+01	1.1E-03 1.3E-03	2.0E+00 1.5E+00	6.1E-02 6.0E-02	3.6E-01 3.4E-01	6.3E-01 6.5E-01	8.1E-02 6.6E-02	49.2 51.4	0.6	98.1 98.0	single crystal separate single crystal separate
11M0125A	Bt	IR (12W)	4	1.5E-05	7.1E+01	-3.1E-03	1.3E+02	5.8E-04	5.2E+00	3.2E-02	3.1E-01	4.5E-01	1.0E-01	66.1	1.0	98.9	single crystal separate
11M0125C 11M0125D	Bt Bt	IR (12W) IR (12W)	5 3	3.8E-05 1.3E-05	3.1E+01 9.2E+01	1.9E-03 -9.9E-04	2.4E+02 4.2E+02	1.1E-03 8.2E-04	3.4E+00 3.4E+00	5.3E-02 4.6E-02	3.7E-01 3.1E-01	5.2E-01 5.7E-01	1.8E-01 8.0E-02	46.7 59.3	0.7	97.8 99.2	single crystal separate
11M0125E	Bt	IR (12W)	5	5.0E-05	3.5E+01	3.5E-03	1.0E+02	1.1E-03	1.9E+00	5.7E-02	3.8E-01	5.8E-01	1.6E-01	48.2	0.9	97.4	single crystal separate single crystal separate
11M0125H	Bt	IR (12W)	3	6.1E-05	2.5E+01	-5.4E-03	9.6E+01	8.9E-04	2.8E+00	4.5E-02	4.2E-01	5.5E-01	1.1E-01	56.6	1.1	96.6	single crystal separate
11M0125I 11M0125M	Bt Bt	IR (12W) IR (12W)	5 3	7.3E-05 3.6E-05	2.2E+01 3.1E+01	2.3E-03 -9.4E-04	2.2E+02 4.8E+02	7.8E-04 7.2E-04	3.9E+00 1.7E+00	4.1E-02 3.9E-02	5.6E-01 3.9E-01	4.4E-01 3.7E-01	2.1E-01 1.8E-01	49.8 45.2	1.3	95.1 97.0	single crystal separate single crystal separate
11M0125N 11M0125O	Bt Bt	IR (12W)	2 2	4.0E-04 1.4E-05	3.8E+00 9.4E+01	1.2E-02 8.5E-03	5.4E+01 8.0E+01	4.0E-03 6.0E-04	7.7E-01 2.3E+00	2.1E-01 3.4E-02	3.0E-01 4.5E-01	2.8E+00 3.6E-01	1.1E-01 2.3E-01	61.4 51.2	0.4	95.8 98.9	single crystal separate
		IR (12W)	-	1.42-03	3.42.01	0.5E-05	0.01.01	0.02-04	2.32.100	J.4L-02	4.51-01	3.01-01	2.51-01	31.2	1.2	36.3	single crystal separate
JT1008, J=0.002 11M0047A	7175±0000172 Wm	IR (0.20 W)	raster	6.7E-06	4.3E+02	-2.1E-04	1.8E+02	-1.1E-05	1.6E+02	3.4E-04	3.4E+01	1.3E-02	1.9E+01	158.1	256.6	85.2	step heating
11M0047B	Wm	IR (0.40 W)	raster	8.1E-04	3.6E+00	2.0E-04	1.8E+02	2.5E-04	6.4E+00	6.0E-03	2.1E+00	3.4E-01	7.6E-01	83.5	14.6	30.4	step heating
11M0047C 11M0047D	Wm Wm	IR (0.55 W) IR (0.70 W)	raster raster	3.5E-04 1.2E-03	8.2E+00 2.5E+00	5.2E-04 8.5E-04	7.2E+01 4.1E+01	3.1E-04 2.3E-03	7.0E+00 1.8E+00	1.7E-02 1.6E-01	8.1E-01 2.6E-01	3.3E-01 3.9E+00	7.7E-01 7.1E-02	63.4 104.5	5.0 0.8	68.5 90.7	step heating step heating
11M0047E	Wm	IR (0.77 W)	raster	8.0E-04	3.9E+00	4.6E-04	7.8E+01	3.4E-03	7.8E-01	2.4E-01	2.6E-01	5.1E+00	8.4E-02	95.9	0.6	95.3	step heating
11M0047F 11M0047H	Wm	IR (0.82 W)	raster	9.3E-04	3.6E+00	5.6E-04	6.2E+01	2.4E-03	2.1E+00	1.6E-01	2.9E-01	3.9E+00	1.1E-01	104.4	0.9	92.9	step heating
11M0047H 11M0047I	Wm Wm	IR (0.87 W) IR (0.92 W)	raster raster	6.8E-04 5.8E-04	5.7E+00 5.3E+00	1.7E-04 3.5E-04	2.1E+02 1.1E+02	3.2E-03 3.4E-03	1.5E+00 1.2E+00	2.3E-01 2.5E-01	2.8E-01 3.0E-01	6.6E+00 5.5E+00	5.1E-02 1.1E-01	129.1 102.2	0.8	96.9 96.8	step heating step heating
11M0047J	Wm	IR (0.97 W)	raster	6.2E-04	5.1E+00	9.7E-04	3.7E+01	2.7E-03	1.2E+00	1.9E-01	3.3E-01	4.1E+00	8.1E-02	96.5	0.8	95.5	step heating
11M0047L 11M0047M	Wm Wm	IR (1.02 W) IR (1.09 W)	raster raster	2.2E-04 4.2E-04	8.3E+00 4.9E+00	3.0E-04 2.7E-04	1.1E+02 1.2E+02	2.2E-03 1.4E-03	2.9E+00 3.8E+00	1.7E-01 1.0E-01	2.5E-01 2.7E-01	3.0E+00 1.9E+00	3.9E-02 1.2E-01	81.8 79.2	0.5	97.8 93.3	step heating step heating
11M0047N	Wm	IR (1.18 W)	raster	2.6E-04	7.9E+00	3.4E-04	9.8E+01	1.2E-03	3.6E+00	9.3E-02	3.0E-01	1.8E+00	8.5E-02	90.4	0.8	95.8	step heating
11M00470	Wm	IR (1.28 W)	raster	9.2E-05	1.8E+01	1.9E-04	1.7E+02	1.0E-03	4.2E+00	8.7E-02	2.6E-01	1.4E+00	7.9E-02	78.1	0.7	98.1	step heating
11M0047P 11M0047Q	Wm Wm	IR (1.43 W) IR (1.65 W)	raster raster	9.8E-05 9.7E-05	1.7E+01 2.2E+01	3.9E-04 6.8E-05	8.1E+01 4.8E+02	8.5E-04 1.3E-03	5.7E+00 4.0E+00	6.8E-02 1.0E-01	3.1E-01 2.8E-01	1.1E+00 1.7E+00	1.0E-01 7.1E-02	76.9 81.8	0.9 0.7	97.4 98.3	step heating step heating
11M0047R	Wm	IR (2.05 W)	raster	2.3E-04	8.7E+00	8.9E-04	3.7E+01	2.1E-03	2.7E+00	1.7E-01	2.8E-01	2.9E+00	3.4E-02	82.8	0.6	97.6	step heating
11M0047S 11M0047T	Wm Wm	IR (2.40 W) IR (2.80 W)	raster raster	1.5E-04 4.6E-05	1.0E+01 4.4E+01	6.2E-04 4.0E-04	5.1E+01 8.1E+01	2.0E-03 9.9E-04	2.7E+00 4.5E+00	1.6E-01 7.6E-02	2.5E-01 3.5E-01	2.7E+00 1.3E+00	4.6E-02 7.9E-02	79.5 79.1	0.5	98.3 98.9	step heating step heating
11M0047U	Wm	IR (4.00 W)	raster	5.6E-05	3.4E+01	3.6E-03	9.6E+00	3.1E-03	1.9E+00	2.4E-01	2.6E-01	5.2E+00	2.9E-02	104.2	0.6	99.6	step heating
11M0047V 11M0047X	Wm Wm	IR (5.50 W) IR (11.80 W)	raster raster	1.7E-04 7.2E-05	1.5E+01 2.8E+01	6.3E-03 1.6E-03	5.6E+00 1.1E+01	7.6E-03 3.8E-03	1.0E+00 1.1E+00	6.0E-01 3.0E-01	2.5E-01 2.4E-01	1.2E+01 6.0E+00	2.4E-02 3.7E-02	93.4 95.1	0.5	99.5 99.6	step heating step heating
		(11.00 W)	1 03161	7.20-03	2.32 701	1.UL-U3	1.11.701	J.0L*03	2.1L FUU	J.0L*01	2.46-01	5.5E 700	J./L*UZ	55.1	0.5	33.0	step nearing
JT1007, J=0.002 11M0046A	27302±0.0000161 Bt	IR (0.15 W)	raster	2.5E-04	4.5E+00	3.5E-04	1.2E+02	1.8E-04	1.6E+01	8.2E-03	5.9E-01	1.6E-01	3.2E-01	52.6	4.0	54.5	step heating
11M0046B	Bt	IR (0.25 W)	raster	3.7E-04	7.1E+00	6.2E-04	7.3E+01	1.6E-03	2.0E+00	8.3E-02	2.7E-01	9.3E-01	7.0E-02	48.3	0.9	88.1	step heating
11M0046C 11M0046D	Bt Bt	IR (0.35 W) IR (0.42 W)	raster raster	3.0E-04 9.7E-05	7.5E+00 1.7E+01	1.3E-03 1.2E-03	3.3E+01 3.5E+01	5.9E-03 5.7E-03	1.2E+00 9.3E-01	3.0E-01 3.0E-01	2.5E-01 2.5E-01	2.9E+00 2.8E+00	4.4E-02 5.2E-02	45.7 44.7	0.3	96.8 98.9	step heating
11M0046E	Bt	IR (0.42 W)	raster	7.5E-05	2.3E+01	1.2E-03 6.5E-04	6.9E+01	5.7E-03 4.0E-03	9.3E-01 1.1E+00	2.2E-01	2.5E-01 2.7E-01	2.8E+00 2.0E+00	5.2E-02 5.2E-02	44.7	0.3	98.9	step heating step heating
11M0046F	Bt	IR (0.60 W)	raster	1.2E-04	1.7E+01	2.3E-03	1.8E+01	4.7E-03	1.3E+00	2.4E-01	2.5E-01	2.3E+00	5.9E-02	45.1	0.3	98.4	step heating
11M0046G 11M0046H	Bt Bt	IR (0.70 W) IR (0.80 W)	raster raster	1.3E-04 8.0E-05	1.4E+01 2.3E+01	3.9E-03 5.1E-03	1.1E+01 8.6E+00	5.3E-03 4.8E-03	1.6E+00 1.6E+00	2.7E-01 2.5E-01	2.5E-01 2.6E-01	2.5E+00 2.3E+00	5.2E-02 4.6E-02	44.8 44.9	0.3	98.4 98.9	step heating step heating
11M0046I	Bt	IR (0.95 W)	raster	7.9E-05	1.7E+01	1.0E-02	4.2E+00	3.9E-03	1.4E+00	2.0E-01	2.6E-01	1.9E+00	4.0E-02	44.8	0.3	98.7	step heating
11M0046J 11M0046K	Bt Bt	IR (1.15 W) IR (2.00 W)	raster	3.5E-05	5.6E+01	1.7E-02	3.0E+00	2.5E-03	1.4E+00	1.3E-01	2.8E-01	1.2E+00	5.8E-02	45.0	0.5	99.1	step heating
11M0046k	Bt	IR (4.00 W)	raster raster	1.3E-05 6.6E-06	1.3E+02 2.3E+02	3.0E-02 1.3E-02	2.4E+00 1.7E+01	1.5E-03 9.9E-04	1.3E+00 1.5E+01	8.2E-02 5.4E-02	2.7E-01 1.2E+00	8.0E-01 5.1E-01	4.1E-02 2.4E-01	47.5 46.4	1.4	99.7 99.7	step heating step heating
11M0046N	Bt	IR (15.00 W)	raster	5.2E-06	3.7E+02	1.5E-02	3.6E+00	1.6E-04	1.3E+01	7.0E-03	7.2E-01	7.0E-02	4.5E-01	48.2	7.8	99.2	step heating
	7175±0.0000174																
11M0063A	Bt	IR (0.15 W)	raster	4.5E-04	5.8E+00	6.3E-04	2.1E+02	3.2E-04	6.6E+00	1.1E-02	1.2E+00	2.0E-01	1.1E+00	28.1	6.9	33.0	step heating
11M0063B 11M0063E	Bt Bt	IR (0.25 W) IR (0.35 W)	raster raster	8.7E-04 2.8E-04	3.9E+00 9.2E+00	9.9E-04 2.3E-03	1.3E+02 5.7E+01	3.8E-03 1.1E-02	1.3E+00 1.0E+00	1.7E-01 5.3E-01	2.5E-01 2.2E-01	1.6E+00 4.1E+00	1.5E-01 6.2E-02	37.1 36.6	0.6	83.6 97.8	step heating step heating
11M0063F	Bt	IR (0.45 W)	raster	1.9E-04	1.7E+01	2.3E-03	5.6E+01	1.2E-02	1.1E+00	5.8E-01	2.3E-01	4.4E+00	6.2E-02	36.5	0.2	98.6	step heating
11M0063H 11M0063I	Bt Bt	IR (0.52 W) IR (0.60 W)	raster raster	6.8E-05	4.6E+01 2.5E+01	1.1E-03 1.7E-03	2.1E+02 7.4E+01	6.2E-03 4.9E-03	1.3E+00 9.0E-01	3.0E-01 2.4E-01	2.8E-01 2.3E-01	2.3E+00 1.9E+00	3.8E-01 9.6E-02	36.8 36.5	0.5	99.0 98.2	step heating
11M0063J	Bt Bt	IR (0.60 W)	raster	1.0E-04 1.2E-04	2.5E+01 2.3E+01	1.7E-03 2.0E-03	7.4E+01 6.7E+01	4.9E-03 7.2E-03	9.0E-01 7.4E-01	3.4E-01	2.3E-01 2.2E-01	1.9E+00 2.6E+00	9.6E-02 8.3E-02	36.5 36.6	0.4	98.2 98.5	step heating step heating
11M0063K	Bt	IR (0.82 W)	raster	2.2E-04	1.2E+01	2.3E-03	5.5E+01	1.0E-02	6.4E-01	4.8E-01	2.2E-01	3.6E+00	5.9E-02	36.4	0.2	98.1	step heating
11M0063L 11M0063M	Bt Bt	IR (0.95 W) IR (1.10 W)	raster raster	1.5E-04 6.2E-05	1.9E+01 4.7E+01	2.0E-03 1.4E-03	6.6E+01 8.9E+01	7.0E-03 3.7E-03	1.1E+00 1.3E+00	3.3E-01 1.7E-01	2.2E-01 2.4E-01	2.5E+00 1.3E+00	6.1E-02 1.0E-01	36.1 36.4	0.3	98.1 98.5	step heating step heating
11M0063N	Bt	IR (1.50 W)	raster	4.7E-05	6.6E+01	9.2E-04	1.4E+02	4.4E-03	1.4E+00	2.1E-01	2.4E-01	1.5E+00	8.1E-02	36.1	0.5	99.0	step heating
11M0063P 11M0063Q	Bt Bt	IR (2.20 W) IR (4.00 W)	raster	1.5E-05 2.3E-05	1.2E+02 7.7E+01	7.7E-04 1.6E-03	1.9E+02 9.6E+01	1.0E-03 1.4E-03	6.7E+00 4.9E+00	5.0E-02 7.1E-02	6.4E-01 4.4E-01	3.8E-01 5.4E-01	4.6E-01 3.2E-01	36.6 36.3	1.2 0.8	98.7 98.6	step heating
11M0063Q 11M0063S	Bt Bt	IR (4.00 W) IR (15.00 W)	raster raster	2.3E-05 2.1E-05	7.7E+01 7.0E+01	5.6E-04	9.6E+01 2.6E+02	9.1E-04	4.9E+00 7.9E+00	7.1E-02 4.5E-02	4.4E-01 6.8E-01	3.5E-01	5.0E-01	36.3	1.1	98.6 98.1	step heating step heating







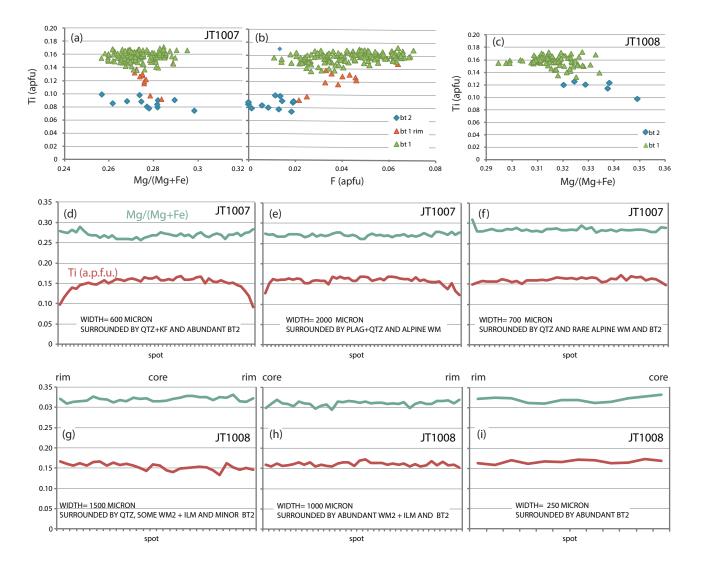
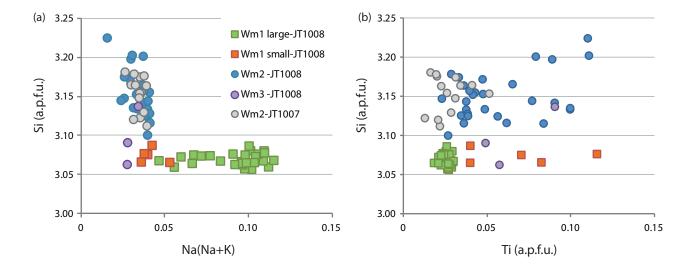


Figure 5



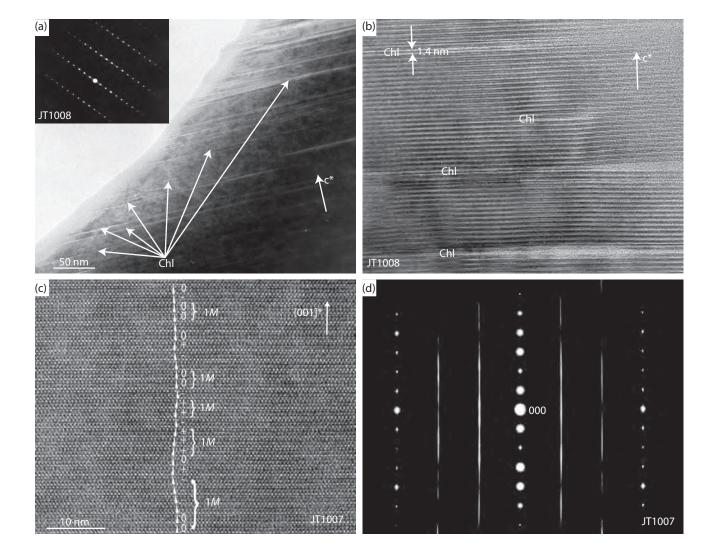


Figure 7

