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A 10,000 yr record of high-resolution Paleosecular Variation from a flowstone of Rio Martino Cave, Northwestern Alps, Italy

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1	A 10,000 yr record of high-resolution Paleosecular Variation from a flowstone of
2	Rio Martino Cave, Northwestern Alps, Italy
3	
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17 Abstract

Speleothems are potentially excellent archives of the Earth's magnetic field, capable of recording its past variations. Their characteristics, such as the continuity of the record, the possibility to be easily dated, the almost instantaneous remanence acquisition and the high time -resolution make them potentially unique high-quality Paleosecular Variation (PSV) recorders. Nevertheless, speleothems are commonly characterized by low magnetic intensities, which often limits their resolution. Here we present a paleomagnetic study performed on two cores from a flowstone from the Rio Martino Cave (Western Alps, Italy). U/Th dating indicates that the flowstone's 25 deposition covers almost the entire Holocene, spanning the period ca. 0.5-9.0 ka, while an estimation of its mean growth rate is around 1 mm per 15 years. The flowstone is composed of 26 columnar calcite, characterized by a highly magnetic detrital content from meta-ophiolites in the 27 28 cave's catchment. This favourable geological context results in an intense magnetic signal that permits the preparation and measurement of thin(~3 mm depth equivalent) samples, each 29 30 representing around 45 yr. The Characteristic Remanent Magnetization (ChRM), isolated after systematic stepwise Alternating Field demagnetization, is well defined, with Maximum Angular 31 Deviation (MAD) generally lower than 10°. Paleomagnetic directional data allow the 32 33 reconstruction of the PSV path during the Holocene for the area. Comparison of the new data with archeomagnetic data from Italian archeological and volcanic records and using the predictions of 34 the SHA.DIF.14k and pfm9k.1a global geomagnetic field models shows that the Rio Martino 35 36 flowstone represents an excellent recorder of the Earth's magnetic field during the last 9,000 37 years. Our high resolution paleomagnetic record, anchored by a high-quality chronology, provide promising data both for the detection of short term geomagnetic field variations and for 38 39 complementing existing regional PSV curves for the prehistoric period, for which well-dated data 40 are still scarce.

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42 Keywords: Paleosecular variation, Rock magnetism, Speleothem, Italy

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44 **1. Introduction**

To investigate geomagnetic field behavior in the past and to explore its short-term features, highresolution records from globally distributed archives of different origin are necessary (Mandea and Olson, 2009). For Paleosecular Variation (PSV) reconstructions, an ideal paleomagnetic record 48 should satisfy several requirements, such as having a stable remanent magnetization, being well dated, offering a continuous record and presenting high-time resolution. Even though some Earth 49 materials may satisfy a number of these, the characteristics of continuity and high-resolution are 50 rarely coupled. Marine and lacustrine sediment sequences are best at ensuring continuous records 51 and have therefore been intensively studied to obtain geomagnetic data over long time scales (e.g. 52 53 Turner and Thompson, 1981; Rolph et al., 2004; Vigliotti, 2006). However, sometimes data reliability may be questionable: the remanence acquisition mechanisms, the smoothing effects of 54 55 bioturbation, the inclination error and the remanence acquisition delay are just some of the problems that may affect this kind of record. On the other hand, volcanic rocks and fired 56 archeological artifacts may preserve very reliable paleomagnetic data but they are highly 57 58 discontinuous in time. The age uncertainties of the volcanic products, as well as the lack of 59 continuity and the limited time extension of available in situ archeological baked clay structures, restrict their use for high-resolution record studies. 60

Several research groups have studied speleothems for both PSV and paleoenvironmental 61 reconstructions (e.g. Latham et al., 1989; Lean et al., 1995; Openshaw et al., 1997; Osete et al., 62 2012; Xie et al., 2013; Font et al., 2014; Jaqueto et al., 2016; Lascu et al., 2016), revealing their 63 64 high potential for magnetic and secular variation reconstructions (Lascu and Feinberg, 2011). Paleomagnetic time series from speleothems, although still sparse, can provide excellent temporal 65 66 resolution if the speleothem has grown continuously and over a considerable age range, as, for example, in the case of the Mexican stalagmite studied in the pioneering work of Latham et al. 67 (1986). The key features of speleothems are that they can grow continuously for 10^3 - 10^5 yr and 68 69 can be accurately dated by the uranium-series method (e.g. Richards and Dorale, 2003). They normally show little or no secondary alteration, and are generally easy to orient and sample 70 71 (though with obvious consideration of natural heritage values).

72 Based on magnetic properties, the remanent magnetization of speleothems can be divided into two main genetic forms, detrital (DRM) or chemical (CRM) (Lascu and Feinberg, 2011). Detrital 73 input can originate from flood or drip water sources (Openshaw et al., 1997; Fairchild et al., 2006). 74 Moreover, speleothems present the advantage of acquiring their magnetization rapidly after 75 76 formation, meaning that the registered magnetic remanence variations reliably reflect the PSV 77 path in the past. Nevertheless, these promising features are confounded by a speleothem's 78 generally low concentration of magnetic minerals, and thus their low magnetic signal, which limits 79 their use in magnetic studies. To bypass this problem, large samples have been commonly used in paleomagnetic studies, but this reduces the time-resolution of the sample. Generally, a sample of 80 81 around 2 cm may average ca 100-4000 yr (Strauss et al., 2013) and thus the obtained SV time-82 resolution is very low.

83 This paper reports the results of a paleomagnetic study performed on a flowstone sampled at Rio Martino Cave (North Western Alps, Italy). The favourable geologic context of the cave, which is 84 mainly surrounded by meta-ophiolites, makes this flowstone very rich in detrital ferromagnetic 85 components, and thus an ideal geomagnetic field recorder due to its high magnetic remanence 86 87 properties. Although a high content of detrital material can compromise U/Th dating (Hellstrom, 88 2006), we have been able to produce a continuous, radiometrically-dated, directional SV record for the area during the last ~10 kyr, at a sampling resolution averaging 45 yr. Comparison of the 89 90 new data with archeomagnetic data from Italian artifacts and volcanic rocks and using predictions 91 of global geomagnetic field models, shows that the Rio Martino flowstone represents an excellent 92 recorder of the Earth's magnetic field in the past and demonstrates the potential of speleothems 93 for PSV studies and for the investigation of short-term variations of the geomagnetic field.

94

95 **2. Geological setting and sampling**

96 Rio Martino Cave (44°42′ N, 7°09′ E) is located in the inner sector of the Western Alps (Northern 97 Italy), which consists of a range of continental and oceanic tectono-metamorphic units bounded 98 by major orogen-scale faulting (Balestro et al., 2014), and exhumed and stacked in the axial sector 99 (Fig. 1). The cave is developed within the Mesozoic carbonate cover of the Palaeozoic Dora Maira 100 (Balestro et al., 2013). This unit is overlain by the Monviso meta-ophiolite complex, a major 101 eclogized remnant of the Ligurian-Piedmont oceanic lithosphere, which in turn is tectonically 102 overlain by the Queyras Schistes Lustrés, interpreted as a fossil accretionary wedge.

The surface above the cave is overlain mainly by glacial deposits. The cave is located at 1530 m a.s.l. on the right flank of the upper Po valley. It is a spring cave, ca. 3000 m long, with 200 m of elevation difference, and is crossed by a small river with an average discharge of 50 l/s (maximum 200 l/s) (Badino and Chiri, 2005).

107 The presence of highly magnetized rocks in the cave's catchment (Fig. 1) and the strong magnetic anomalies observed in the Monviso Massif area (Lanza and Meloni, 2006) could induce a magnetic 108 109 deflection effect in the area. To evaluate the possible effect exerted by the meta-ophiolitic masses and to confirm that it does not exert a significant influence on the paleomagnetic sampling, we 110 111 used a triaxial fluxgate magnetometer to measure the geomagnetic field components outside and 112 next to the entrance, as well as inside the cave. The computed magnetic inclination values of 60.7° outside the cave and 60.5° on the flowstone surface are fully comparable to the 2013 IGRF model 113 114 of 60.6° (<u>http://www.ngdc.noaa.gov/geomag-web</u>). Besides, outside the cave we performed some 115 orientation checks by using both the magnetic and the solar compass. The difference between the two declinations was small, ranging from -5° to +2°. Such differences are insignificant and indicate 116 117 that any local magnetic effects on the paleomagnetic sampling can be considered negligible.

118 Two sampling campaigns were carried out to collect two cores from the same flowstone, which 119 has accumulated on the side of a seasonally active stream with a high-detrital content. The cores were taken ~20-30 cm apart and drilled using an adapted electric-powered drill. The first core (RMD1), sampled during a campaign in 2010, was not azimuthally oriented. The second core (RMD8), sampled in 2013, was oriented *in situ* by magnetic compass and inclinometer (figure S1 in the supplemental material). Each core was ca. 60 cm long and was drilled perpendicular to the flowstone growth axis.

A quarter of each core was dedicated to paleomagnetic analysis. The investigated sub-samples consisted of small slices, about 3 mm thick (varying from 2.5 to 4 mm), cut almost perpendicular to the speleothem's growth direction. Slicing was performed using a very thin non-magnetic saw, which ensured that only 1 mm of material was consumed during the cut. Following this systematic sampling, we obtained 146 slices from RMD1 and 143 from RMD8. Each slice was positioned in the centre of a non-magnetic plastic cylinder (2.5 cm diameter, 2.3 cm height) that allowed its handling as per standard paleomagnetic samples (Fig. 2).

132

133 **3. Methods**

134 *3.1. U/Th dating and age modelling*

Nineteen solid prisms of ~40 mg (~2 mm wide along the lamina and 1 mm thick on growth axis) 135 136 from RMD1 were used for age determination (Table S1 in Supplementary Material). The U/Th dating was performed at the University of Melbourne (Victoria, Australia) following the method of 137 Hellstrom (2003). Briefly, samples were dissolved and a mixed ²³⁶U-²³³U-²²⁹Th spike was added 138 prior to removal of the carbonate matrix with ion-exchange resin. The purified U and Th fraction 139 was introduced in a dilute nitric acid to a multi-collector inductively coupled plasma mass 140 spectrometer (MC-ICPMS, Nu-Instruments Plasma). The ²³⁰Th/²³⁸U and ²³⁴U/²³⁸U activity ratios 141 142 were calculated from the measured atomic ratios using an internally standardised parallel ioncounter procedure and calibrated against the HU-1 secular equilibrium standard. Correction for detrital Th content was applied using initial activity ratios of detrital thorium $(^{230}Th/^{232}Th)_i$ of 1.3 ± 0.45. This value, and its relative 2 σ uncertainty, was calculated using a Monte Carlo 'stratigraphic constraint' procedure based on the series of U/Th ages (Hellstrom, 2006). A depth-age model was constructed using a Bayesian Monte Carlo approach following the method described by Drysdale et al. (2005) and Scholz et al. (2012).

149 3.2. SEM-EDS analysis

150 The mineralogy of the detrital inclusions in the studied flowstone was investigated by dissolving 151 different portions of various thin slabs of the RMD1 core in diluted hydrochloric acid and passing the digests through 0.45 micrometre cellulose acetate filters. The residues, bearing almost all of 152 the non-carbonate mineral inclusion types contained in the speleothem, were observed and 153 analysed with a Cambridge Stereoscan 360 Scanning Electron Microscope housed at the Earth 154 155 Science Department of the University of Torino, Italy. Analyses were performed using an Oxford 156 Inca X-Act 200 EDS microanalysis equipped with a Link Pentafet detector (thin window), allowing qualitative/quantitative determination of light elements (down to boron). All data were obtained 157 at 15 kV HT, 25 mm WD, probe current range 800 pA – 1.2 nA and analysis time from 60 to 500 s. 158 Primary standardization was performed on SPI Supplies and Polaron Equipment standards, and the 159 system was regularly calibrated against a high-purity metallic Co standard before each 160 161 experimental session. Data were processed with the Inca 200 Microanalysis Suite Software, 162 version 4.08, and calibrated on natural mineral standards using the ZAF correction method. Analytical data are considered to be only semi quantitative due to the nature of the samples 163 (rough surface of the particles, lack of horizontality, lack of surface polishing). A total of about 164 1500 analyses was performed on seven samples coming from different portions of the core, 165

166 corresponding to about 200 measurements for each filter, randomly scattered on the filter surface167 for better representativeness.

Despite the results of magnetic analysis, very few magnetite particles were found in the filtered material, most likely because the single-domain magnetic particles were not retained by the 0.45 μm filter. Magnetite was indeed observed in sandy materials from Rio Martino, being found in the bed sediments of the relatively high-energy environment of the cave stream, rather than as detritus in carbonate flowstone speleothems.

173

174 *3.3. Rock magnetic measurements*

175

176 All magnetic measurements were performed at the ALP Paleomagnetic Laboratory (Peveragno, Italy). Rock magnetic experiments were performed on representative samples from both cores. 177 Rock magnetism was investigated by low-field susceptibility (k_m) and natural remanent 178 179 magnetization (J_r) measurements using a KLY3 kappabridge and a JR6 spinner magnetometer with a sensitivity of the order of 10^{-8} SI and 10^{-6} A/m, respectively. Susceptibility was measured at least 180 five times per sample in order to calculate a mean value. Standard deviation is low and normally 181 182 less than 5% for specimens associated with a susceptibility spike; uncertainty grows to 20-35% for the remaining specimens, characterized by negative (diamagnetic) susceptibility values. All 183 samples were weighed to get the mass-normalized susceptibility (χ , m³kg⁻¹) and intensity (J, 184 Am²/Kg). Their values are represented as a function of the core depth in figure S2 of the 185 supplemental material. 186

Isothermal Remanent Magnetization (IRM) curves were obtained with an ASC pulse magnetizer,
applying stepwise increasing fields up to 1 T. Thermal demagnetization of a three-axis composite
IRM was also performed on representative samples (Lowrie, 1990). An IRM was imparted with an

ASC pulse magnetizer along the sample's three orthogonal axes, applying first a maximum 1.5 T, then a medium 0.3 T and finally a minimum 0.1 T magnetic field. Crossover plots of IRM curves and alternating field (AF) demagnetization of the saturation IRM (SIRM) were carried out to investigate the magnetic grain size (Symons and Cioppa, 2000).

Finally, hysteresis cycles were obtained by a Vibrating Sample Magnetometer (VSM, LakeShore 7410 - Maximum applied field $B_{max} = \pm 1$ T; $H = \pm 10.000$ Oe / 976.000 A/m) at the Istituto Nazionale di Ricerca Metrologica (INRIM, Torino) and interpreted by RockMag Analyzer 1.0 software (Leonhardt, 2006). All specimens were AF demagnetized stepwise up to 100 mT with a ASC-D 2000 equipment. Representative twin specimens were also stepwise thermally demagnetized with a Schonstedt TSD-1 furnace.

200

201 **4. Results**

202 4.1. Chronology

203 All the U/Th ages obtained from RMD1 were in stratigraphic order within the associated uncertainties, except for two samples that were consequently rejected as outliers (Table S1 in 204 Supplementary Material). Macroscopic and thin-section analyses of core RMD1 shows no growth 205 206 interruption along its length. Age modelling performed on RMD1 indicates that the flowstone grew continuously between 0.56 \pm 0.06 ka and 9.7 \pm 1.6 ka b2k (Fig. 3). The mean growth rate is 207 208 0.058 mm/yr, which implies a mean time-resolution of ca. 60 yr (3 mm specimen + 1 mm cut) for 209 the PSV record. The time averaged in each 3 mm slice sample is ca. 45 yr. The age of RMD8 was inferred by comparing clearly visible growth layers (Fig. 2a) between the two cores, associated 210 with spikes in both the magnetic mass susceptibility and mass magnetization (figure S2 of the 211 supplemental material). 212

213 *4.2. EDS*

214 The mineralogy of the detrital portion in the RMD1 core is in strong accord with the composition of the surrounding lithology. Apart of the calcareous formation in which the cave has developed, 215 the main rocks in the area are prasinites, amphibolites and serpentines. Minerals were grouped by 216 similar chemistry, with some simplifications: as stated above, analyses were only semi-217 quantitative. The main identified groups are: iron oxides (without magnetite, discriminated by 218 morphological features), magnesium silicates (other than serpentine), serpentine group, white 219 mica group, feldspar, tremolite-actinolite amphiboles, other amphiboles (mainly hornblende), 220 epidote group, chlorite group, quartz and accessories. The main minerals (Fig. 4) are represented 221 222 by iron oxides (not distinguishable by chemistry for the reason explained above), magnesium silicates and serpentine group minerals. Iron oxides are mostly irregular in shape as if they had 223 224 undergone reworking from the stream or by feedwater (Perkins, 1996). In few cases, a framboidal 225 shape suggests in situ growth.

226

227 4.3. Magnetic mineralogy

The mass magnetic susceptibility of the specimens strongly varies. It mostly shows a prevailing 228 diamagnetic phase with small negative values (from -7 to 0 x 10^{-9} m³ kg⁻¹ with a mean value of -4 x 229 10^{-9} m³ kg⁻¹), alternating with high positive spikes, up to 970 x 10^{-9} m³ kg⁻¹, suggesting a very low 230 concentration of magnetic minerals in these specimens. Calcite bulk susceptibility is -12.09 µSI; its 231 mass susceptibility is about -4.46 x 10⁻⁹ m³ kg⁻¹ (Almqvist et al., 2010). Since the literature value for 232 233 the susceptibility of calcite refers to single crystal, we can assume that the mass susceptibility for calcite in the speleothem is slightly higher, because of mineral porosity. Assuming a constant 234 diamagnetic contribution mostly due to calcite, the relative variability of magnetic susceptibility is 235

236 indicative of variations of the concentration of magnetic minerals: a mean χ value of -4 x 10⁻⁹ m³ 237 kg⁻¹ can be assumed to be representative of the "standard" content in magnetite, while high 238 values represent for pulses of higher detrital input.

The natural magnetization intensity (J_r) strongly varies from specimen to specimen, being on average around 1-10 x 10⁻⁶ Am² kg⁻¹ with spikes up to 80 x 10⁻⁶ Am²kg⁻¹. The variations of these two bulk parameters are correlated; the computed correlation coefficients are r = 0.87 and r =0.76 for RMD1 and RMD8, respectively. This corroborates the hypothesis that their values are essentially controlled by changes in concentration of the magnetic oxide.

IRM acquisition curves from representative samples saturate at relatively low field (around 0.3 T), 244 indicating the presence of low coercivity minerals (Fig. 5a). The IRM acquisition data were further 245 analyzed applying the MAX UnMix software (Maxbauer et al., 2016) to six flowstone slices and to 246 247 sand collected from the bed of the cave spring. For three specimens, i.e. RM101 and RM134 248 flowstones and the sand, only one magnetic component was identified with Bh and DP values fully consistent with detrital magnetite (Bh = 1.46, DP = 0.40). In the remaining cases, two components 249 were detected, one as above, and the other pointing to higher-coercivity magnetite. The 250 computed S-ratio ($S_{-0.3 \text{ mT}}$) ranges from 0.95 to 1.00. 251

During the thermal demagnetization of the orthogonal IRM components (Lowrie, 1990), two 252 253 typical behaviors were observed, which are independent of the magnetization intensity of the specimens. The first (e.g. sample RM68a), representing about the 80% of the measured 254 specimens, suggests that the primary remanence is dominated by a soft magnetic carrier, 255 demagnetized at ca 350-450 °C, which is interpreted as a titanomagnetite (Fig. 5b). The second 256 (e.g. sample RM44a), in the remaining 20%, is characterized by a first drop in the magnetization 257 intensity between 200 and 300 °C, which may be speculatively related to the existence of 258 maghemite (Pan et al., 2000), even though this evidence is not sufficient to unambiguously 259

260 identify this magnetic phase (Zhu et al., 2012). The presence of (titano) magnetite of detrital origin is easily justified considering the geologic context of the cave and it is probably originated from 261 the highly magnetic rocks of the surrounding area, mostly meta-ophiolites (Balestro et al., 2013). 262 The occurrence of small serpentinite lithics was also detected. In those cases, the Median 263 Destructive Field (MDF), which is normally stable and around 50-60 mT, drops to 5-25 mT. 264 265 Deflections from MDF = 50 mT occur only where both the magnetic susceptibility and the magnetization intensity values are high. To check for these variations, we performed the 266 267 experiment of Symons and Cioppa (2000) on some selected specimens, which were characterized by MDF ranging from 15 to 60 mT. It consists of a crossover plot, where the %SIRM is plotted as a 268 function of the applied field, and using a logarithmic scale (Fig. 5c). The results suggest that, 269 270 except specimen SP212, which is characterized by MD (titano)magnetite, samples mainly contain 271 SD to PSD (titano)magnetite grains.

The modified Lowrie-Fuller method (Johnson et al., 1975), which represents a valid first-order indicator of grain-size composition (Font et al., 2014), was applied to some specimens with MDF in the range 45-50 mT. Results always show L-type behaviors with the Anhysteretic Remanent Magnetization (ARM) dominating the Isothermal Remanent Magnetization (IRM) during AF demagnetization treatment, corroborating the occurrence of SD grains.

277 These results are confirmed by the hysteresis cycles (Figure S3 in the supplemental material)

278 performed on 2 specimens, one characterized by high χ , one by low χ . After correction for the

279 diamagnetic effect, the curves show the occurrence of a low-coercivity phase with saturation at

280 0.3-0.5 T; in both cases, the ratios M_{rs}/M_s and B_{cr}/B_c fall next to the PSD field and SD + MD mixing

curve in the Day plot (Dunlop, 2002).

282 **5. Paleomagnetic directions**

283 5.1. Natural Remanent Magnetization and the Anisotropy of Remanent Magnetization

284 Speleothems can potentially offer very useful records of PSV and the remanence acquisition mechanisms in speleothems have been previously studied in detail (e.g., Lascu and Feinberg, 2011; 285 Strauss et al., 2013, and reference therein). In order to provide a reliable PSV record, the 286 magnetization should be acquired and locked soon after the calcium carbonate film is deposited 287 on a speleothem (almost instantaneously). Following Strauss et al. (2013), lock-time for a 288 289 speleothem is sub-annual and the magnetization is a DRM. Synchronicity between crystallization and magnetization has been tested experimentally by synthetic stalagmite growth (Morinaga et 290 291 al., 1989), confirming the short time-lapse in acquiring magnetization parallel to the ambient field direction. 292

To test if this requirement is encountered in Rio Martino flowstone and thus to check for its 293 reliability as a PSV recorder, we measured both the Anisotropy of Magnetic Susceptibility (AMS) 294 295 and the Anisotropy of Isothermal Remanent Magnetization (AIRM) on two selected sets of samples from RMD1 (azimuthally non-oriented core), each comprising a time interval of ca 1000 296 yr. The first set comprised 16 samples (SP200 to SP260) from 4.26 ± 0.23 to 3.30 ± 0.03 ka, and the 297 second set 14 samples (SP346 to SP397) from 7.76 ± 0.12 to 6.91 ± 0.11 ka. A difference of 20° in 298 the mean magnetic ChRM inclination distinguished these two sets of specimens. The AMS was 299 300 measured by a KLY-3 kappabridge. The results obtained on both sets of samples show a welldefined (confidence angles $< 15^{\circ}$) mean minimum susceptibility axis, k_3 , which is statistically 301 302 vertical and perpendicular to the flowstone growth laminae (Fig. 6a). For the AIRM measurements, each specimen was first AF demagnetized using a tumbling 2G demagnetizer at 60 mT peak field 303 304 and then given an isothermal remanent magnetization (IRM) with a steady field of 20 mT using an 305 AGICO PUM-1 pulse magnet. After measurement with the spinner magnetometer, the sequence 306 was repeated for a total of 12 different orientations of the IRM in order to calculate the anisotropy 307 tensor. The experiment (Fig. 6b-c) shows that for both sets, the maximum IRM anisotropy axis I₁ is

308 concordant or statistically indistinguishable from the mean ChRM direction, showing no relation with the speleothem growth laminae. This shows that ChRM direction is due to the statistical 309 alignment of the magnetic particle and fully agrees with the conclusions of Zhu et al. (2012), who 310 performed both AMS and AIRM on stalagmites. They found that the AMS was dominated by the 311 calcite fabric, being the minimum susceptibility axis k₃ aligned perpendicular to the stalagmite 312 313 growth laminae, while the AIRM fabric showed the maximum remanence axis I₁ almost parallel to the NRM direction. These data all point to a detrital origin of the magnetization, with the 314 315 geomagnetic field control in the orientation of the ferromagnetic minerals.

316

317 5.2. Characteristic Remanent Magnetization determination

Demagnetization results are represented by intensity-decay curves and plotted in Zijderveld 318 diagrams (Fig. 7). Most of the specimens are characterized by a small viscous remanent 319 magnetization (VRM), which is easily removed at an AF field of 15-20 mT. The remaining 320 demagnetization path is linear and points to the origin, indicating a stable remanent 321 322 magnetization; this component has been interpreted as the Characteristic Remanent 323 Magnetization (ChRM). The ChRM direction is mostly well defined and characterized by low MAD values (lower than 8° for the 91% of the studied samples). AF and thermal demagnetization results 324 325 obtained from twin specimens are very similar (Fig. 7), confirming the reliability of the ChRM direction (Fig. 8; Table S2 in Supplementary Material). In the thermal demagnetization (Fig. 7b and 326 7d), an inflection in the intensity decay curve it is observable at ca. 150 °C. As suggested by Strauss 327 328 et al. (2013) the decay could be correlated with the occurrence of goethite, not detected however by any other experiment. The AF demagnetization treatment has been preferred rather than the 329 thermal demagnetization, as it permits the further use of the same samples for 330 paleoenvironmental and relative paleointensity investigations through Anhysteretic Remanent 331

Magnetization (ARM) measurements. Therefore, all samples were systematically AF demagnetized and ChRM directions were obtained from the AF demagnetization results. Demagnetization behavior in samples with low- and high-remanence (spike) does not change significantly except for specimens where serpentinite clasts were recognized.

336

6. Directional Paleosecular Variation during Holocene

338 Paleomagnetic directions obtained from the two cores (reported in Table S2 of Supplementary material) are plotted versus depth from the top of the core in Figure 9. Some spikes in declination 339 340 show a strong correspondence with atypical MDF values, lower/higher than 20/60 mT. These 341 deflected directions have been ascribed to the presence of small serpentinite lithic fragments and were thus rejected. The declination of RMD1, which was not azimuthally oriented, has been 342 recovered after adjustment of its mean value to the Geocentrical Axial Dipole (GAD) calculated at 343 Rio Martino according to the following procedure: first, the mean direction of RMD1 has been 344 calculated for the last 10 kyr and then its deviation from the GAD value has been computed. The 345 difference in declination between the core and the GAD has been extracted from each declination 346 value of the RMD1 core. 347

Generally, directions obtained from both RMD1 and RMD8 are in good agreement with each other and data reproducibility is high (Fig. 9). This is particularly evident for the inclination data at depths between 200.0 and 600.0 mm, where the two records match each other. Instead, in some cases, mostly at depths from 150.0 to 200.0 mm, differences in inclination of around 20°-25° are observed. The cause of such differences is not clear, even though uncertainties during sampling (slices not perfectly perpendicular to the flowstone growth) and deflections related to a possible anisotropy effect connected to the calcite crystals growth cannot be completely excluded. To guarantee the high quality of the new data, only ChRM directions characterized by MAD values lower than 6° have been used in the plots. Directional data of each core have been kept clearly distinguishable, because only RMD1 was directly dated and even though the paleomagnetic records from the two cores are consistent and despite the fact that the two cores are only 30 cm apart, their simultaneous growth cannot be fully guaranteed.

360 The paleosecular variations registered by the Rio Martino speleothem are compared with spot archeomagnetic directions obtained from dated archeological structures and volcanic rocks from 361 362 Italy. The data from the *in situ* archeological material are taken from the Italian archeomagnetic dataset (Tema et al., 2006; Tema, 2011) updated by some recently published results (Malfatti et 363 al., 2011; Kapper et al., 2014; Tema et al., 2013; 2015; 2016). The data from the Italian volcanic 364 eruptions have been downloaded from the GEOMAGIA50.v3 database (Brown et al., 2015). All 365 366 data have been relocated at the geographic coordinates of Rio Martino via the virtual geomagnetic pole method (Noel and Batt, 1990). The comparison shows that the archeomagnetic and volcanic 367 data generally fit very well to the speleothem directions (Fig. 10). Some discrepancies in 368 declination can be observed around 1000 AD. For this period the speleothem declination values 369 are very low and quite dispersed. Nevertheless, it is particularly interesting to note that for the BC 370 371 period, the available archeomagnetic data, even if very limited and often accompanied by large 372 error bars, are in excellent agreement with the new data, in spite of the speleothem inclination 373 being systematically higher. Recently, Ponte et al. (2016) noticed that the inclination values in speleothems seem to vary as a function of the orientation of the calcite growth. To check for this 374 effect, we calculated the mean paleomagnetic direction for both RMD1 and RMD8 applying Fisher 375 376 Statistics, selecting ChRMs with MAD < 6°. Their mean directions (RMD1: D = 358.8°, I = 65.1°, α_{95} = 1.9°; RMD8: D = 2.4°, I = 65.8°, α_{95} = 1.7°) are close to the GAD at Rio Martino's geographic 377 coordinate (D = 0°, I = 63°), even if statistically distinguishable. However, AIRM shows a reliable 378

consistency between the mean ChRM and the mean I_1 axis, which confirms that the orientation of the ferromagnetic minerals was mainly controlled by the geomagnetic field present during speleothem accumulation. This substantiates the high potential of Rio Martino speleothems to continuously and reliably register the Earth's magnetic field, offering a unique source of high quality data for the BC period where *in situ* archeological artifacts are very scarce.

384 The new data are also compared with the predictions of global geomagnetic field models. Here, we have used for comparison the pfm9k.1a (Nilsson et al., 2014) and the SHA.DIF.14k (Pavón-385 386 Carrasco et al., 2014) models that are the most recently published global geomagnetic models for the Holocene. There is good agreement between the speleothem records and the global models 387 predictions, confirming some interesting features of the Earth's magnetic field in the past. The 388 eastward declinations around 1000 BC mainly observed in the SHA.DIF.14k model are observed in 389 390 the speleothem data for the same time period, and also show high declination values. For the 4000-2000 BC period, only small declination variations are shown by the speleothem data, in 391 agreement with the pfm9k model's predictions, while the declination peaks seen in the 392 SHA.DIF.14k model (e.g. around 3600 BC) are not confirmed by the speleothem data. For periods 393 394 older than 5000 BC, speleothem records show generally higher declination values compared to the 395 model predictions and other archeomagnetic data. Interestingly, similar eastward declination values were found for cave sediments in Switzerland (Kapper et al., 2014). Regarding the 396 397 inclination data, good agreement can be observed for the periods 6000-3500 BC and 500 BC-500 AD. However, around 1000 BC, speleothems show an interesting high-inclination peak that is not 398 399 observed in the models or sustained by the available archeomagnetic data. This peak is actually 400 only observed on the data from the RMD8 core and definitely more independent records are 401 necessary to investigate if it corresponds to a real abrupt directional change (as it corresponds also 402 to high declination values) of the geomagnetic field at this time period. For the 7500 BC to 6000 BC

403 period, the speleothem records show continuously increasing inclination with a peak around 6000
404 BC that seems to be in agreement with the model's predictions.

405

406 **7. Conclusions**

Some outstanding characteristics of the Rio Martino flowstone, such as its continuous growth, the well-constrained chronology and the intense magnetic signal, make its paleomagnetic directional record for the Holocene in the northwestern Italy particularly appropriate for PSV investigation. The high magnetic signal permits a high-resolution record of around 60 yr per data point; the regular scatter of paleomagnetic data through time shows an almost constant distribution of directional data though the Holocene.

The obtained directional results are well defined and offer a unique, almost continuous, secular variation record for the last ~10,000 years. Although some discrepancies can be observed, comparison with archeomagnetic data and global geomagnetic field models confirms the high potential of these speleothems to the reconstruction of the Earth's magnetic field variations in the past.

Our results show that the Rio Martino flowstones are not affected by recrystallization effects or secondary alterations. The speleothems do not show any inclination shallowing when compared with model predictions, and in some cases show high inclination peaks that are not observed by the models (e.g. around 3800 BC, 1000 BC, 800 AD).

The record characteristics overcome some typical features affecting both clastic sedimentary and the archeomagnetic PSV records, including the smoothness of the magnetic data in the case of the former and the presence of temporal gaps and uneven data distribution in the case of the latter. The high resolution obtained points to the possibility of detecting short and abrupt geomagnetic field changes by studying a wide variety of Earth Magnetic Field variations at a timescale from tens of years to the millennia and highlights the importance of regional differences when modelling the
Earth's field. The use of speleothem records for PSV reconstructions can be particularly important
for the prehistoric period where other sources if data coming from archeological artifacts or welldated volcanic eruptions are scarce.

431

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577 Figure Caption

578

Figure 1. a) Structural sketch map of the Western Alps; b) 3D reconstruction of the Inner Western Alps in the Rio Martino zone (modified after Balestro et al., 2014). The square indicates the location of the Rio Martino Cave. Note: the region highlighted in the small inset map is not the same as the one shown in the enlargement.

Figure 2. a) Part of core RMD8; b) part of the flowstone systematically cut and sampled in 3 mmthick slices; c) the amagnetic plastic cylindrical holder created in order to fix the small samples in the centre of the cylinder and treat them as standard paleomagnetic samples.

Figure 3. Age-depth model for RMD1 core. The age is expressed both in ka AD and in b2k (before 2ka).

Figure 4. The distribution of the mineral species in the insoluble (detrital) fraction of Rio Martino speleothem. The picture is the sum of ca 1500 EDS determinations from seven different portions from the same core. The "accessory minerals" are species <2.5 % of the analyzed particles, for each sample and includes rutile, zircon, monazite, apatite (mainly apatite-F), sphene, xenotyme, galena, pyrite, ilmenite, barite.

593 Figure 5. a) Isothermal remanent magnetization (IRM) acquisition curves; b) thermal 594 demagnetization of a composite three-axes IRM (Lowrie, 1990); c) crossover plots (Symons and 595 Cioppa, 2000).

Figure 6. Equal area stereographic projections of the eigenvectors for a) the anisotropy of magnetic susceptibility, and for b) and c) isothermal remanent magnetization, where the maximum, intermediate, and minimum eigenvectors are denoted by squares, triangles, and circles, respectively. The 95% confidence ellipses for the eigenvectors are shown by unfilled ellipses. The mean ChRM directions and their alpha95 errors for specimens SP200 to SP260 (b) and
SP346 to SP397 (c) are denoted by stars with grey ellipses.

Figure 7. Thermal and AF demagnetization results from twin specimens from samples a-b) RM7 and c-d) RM20 plotted in intensity decay plots (left) and Zijderveld diagrams (right). Symbols: full dots = declination; open dots = apparent inclination.

Figure 8. Equal-area projections of the ChRM directions for five samples obtained from a) AF and
b) thermal demagnetization on twin specimens. The star represents the mean value calculated for
each group of samples following a Fisherian distribution.

Figure 9. a) Declination and b) inclination data from cores RMD1 (red) and RMD8 (blue) plotted
versus depth in mm from the top of the core.

Figure 10. a) Declination and b) inclination plots of the RMD1 (red) and RMD8 (blue) compared with the Italian archaeomagnetic data from archeological artefacts (green diamonds) and volcanic rocks (black squares) and the pfm9k (magenta line) and SHA.DIF.14k (black line) global geomagnetic field models. All directions are calculated at the geographic coordinates of Rio Martino (44.7° N, 7.15° E). Age is given both as Calendar Age (year AD) and b2k (before 2 ka).

- 615 Figure caption of the Supplemental material
- Figure S1. a) Photograph of the flowstone's sampling and b) of the device to orient the core.
- Figure S2. a) Mass susceptibility and b) mass remanence intensity from RMD1 (red) and RMD8(blue) plotted versus depth in mm from the top of the core.
- Figure S3. Hysteresis curves (mass magnetization versus applied field) for specimens SP421 and
 SP460. In a) uncorrected curves; in b) slope-corrected for diamagnetic effect curves.

621 Table caption

Table S1. Corrected U/Th ages for RMD1 core. Isotope ratios are expressed as activity ratios standardized to the HU-1 secular equilibrium standard. Ages have been calculated using decay constants of 9.195×10^{-6} (²³⁰Th) and 2.835×10^{-6} (²³⁴U). Depths are from top.

Table S2. Characteristic remanent magnetization directions (ChRMs) of the samples from the RMD1 (left) and RMD8 (right) cores. Legend: z = depth in mm from the core top; D, I = magnetic declination and inclination; MAD = Mean Angular Deviation; $D_{corr} = declination$ corrected by subtracting the angular difference between the $D_{GAD} = 0^{\circ}$ and the mean ChRM declination (D = 146.1°).



630







635 Figure 3



637 Figure 4



638











644 a) Af demagnetization

b) Thermal demagnetization

645 Figure 8







649 Figure 10