Contents lists available at ScienceDirect

Physics Letters B



www.elsevier.com/locate/physletb

Three loop calculations and inclusive V_{ch}

Marzia Bordone, Bernat Capdevila, Paolo Gambino*

Dipartimento di Fisica, Università di Torino & INFN, Sezione di Torino, I-10125 Torino, Italy

ARTICLE INFO

ABSTRACT

Article history: Received 13 July 2021 Received in revised form 19 September 2021 Accepted 22 September 2021 Available online 27 September 2021 Editor: B. Grinstein

We discuss the impact of the recent $\mathcal{O}(\alpha_s^3)$ calculations of the semileptonic width of the *b* quark and of the relation between pole and kinetic heavy quark masses by Fael et al. on the inclusive determination of $|V_{cb}|$. The most notable effect is a reduction of the uncertainty. Our final result is $|V_{cb}| = 42.16(51) 10^{-3}$. © 2021 The Author(s). Published by Elsevier B.V. This is an open access article under the CC BY license (http://creativecommons.org/licenses/by/4.0/). Funded by SCOAP³.

1. Introduction

The purpose of this note is to study the implications of the recent $\mathcal{O}(\alpha_s^3)$ calculations by Fael, Schönwald and Steinhauser [1–3] on the determination of the Cabibbo-Kobayashi-Maskawa matrix element $|V_{ch}|$ from inclusive semileptonic *B* decays, see Refs. [4,5] for the most recent results. As is well-known, the values of $|V_{cb}|$ determined from inclusive semileptonic B decays and from $\bar{B} \rightarrow$ $D^*\ell\bar{\nu}$ have differed for a long time and, despite significant experimental and theoretical efforts, the situation remains quite confusing. Recent accounts of the V_{cb} puzzle can be found in Refs. [6–9]. The latest lattice calculations [10-13], which for the first time explore the $\bar{B} \to D^* \ell \bar{\nu}$ form factors at non-zero recoil, have not clarified the issue and a preliminary comparison of these results shows a few discrepancies [14].

The third order perturbative correction to the $b \rightarrow c \ell \bar{\nu}$ decay width computed in Ref. [1] and partially checked in Ref. [15] represents a fundamental step to improve the precision in the extraction of $|V_{cb}|$ from inclusive B decays. Indeed, perturbative corrections are sizeable - they reduce the semileptonic rate by over 10% - and provide the dominant theoretical uncertainty. In the following we will study the impact of the $\mathcal{O}(\alpha_s^3)$ corrections on the central value and uncertainty of $|V_{cb}|$.

The other relevant three-loop calculation in our analysis is that of Refs. [2,3], which concerns the relation between the pole (or MS) and the kinetic masses of a heavy quark. This calculation allows us to convert recent high-precision determinations of the b quark $\overline{\text{MS}}$ mass [16,17] into the kinetic scheme [18] with an uncertainty of about 15 MeV, or 50% less than the previous two-loop conversion [19,20]. We will investigate the effect of such high-

Corresponding author.

E-mail addresses: marzia.bordone@to.infn.it (M. Bordone), bernat.capdevilasoler@unito.it (B. Capdevila), paolo.gambino@unito.it (P. Gambino).

precision input in the global fit to the semileptonic moments and in the extraction of $|V_{ch}|$. We should add that Ref. [3] also computed the charm mass effects at $\mathcal{O}(\alpha_s^2, \alpha_s^3)$ in the kinetic scheme relations. We will also include $\mathcal{O}(\alpha_s \rho_D^3/m_b^3)$ effects from Ref. [21] in the semileptonic width and slightly update the global fit to the semileptonic moments.

2. The total semileptonic width

Our starting point is the Operator Product Expansion (OPE) for the total semileptonic width (see Ref. [22] for a complete list of references):

$$\Gamma_{sl} = \Gamma_0 f(\rho) \Big[1 + a_1 a_s + a_2 a_s^2 + a_3 a_s^3 - \left(\frac{1}{2} - p_1 a_s\right) \frac{\mu_{\pi}^2}{m_b^2} \\ + (g_0 + g_1 a_s) \frac{\mu_G^2(m_b)}{m_b^2} + d_0 \frac{\rho_D^3}{m_b^3} - g_0 \frac{\rho_{LS}^3}{m_b^3} + \dots \Big]$$
(1)

where $\Gamma_0 = A_{ew} |V_{cb}|^2 G_F^2 m_h^{kin}(\mu)^5 / 192\pi^3$, $f(\rho) = 1 - 8\rho + 8\rho^3 - 6\rho^2 + 8\rho^2 + 8\rho^2$ $\rho^4 - 12\rho^2 \ln \rho$, $a_s = \alpha_s^{(4)}(\mu_b)/\pi$ is the strong coupling in the $\overline{\text{MS}}$ scheme with 4 active quark flavours, $\rho = (\overline{m}_c(\mu_c)/m_b^{kin}(\mu))^2$ is the squared ratio of the $\overline{\text{MS}}$ charm mass at the scale μ_c , $\overline{m}_c(\mu_c)$, and of the *b* quark kinetic mass with a cutoff $\mu \sim 1 \text{ GeV}$, $m_b^{kin}(\mu)$. $A_{ew} \simeq 1.014$ is the leading electroweak correction. The parameters $\mu_{\pi}^2, \rho_{\rm D}^3$, etc. are nonperturbative expectation values of local operators in the *B* meson defined in the kinetic scheme with cutoff μ . They are generally extracted from a fit to central moments of the lepton energy and of the hadronic invariant mass distributions in semileptonic *B* decays [4,5], for which the same contributions as in Eq. (1) are included, with the exception of the $\mathcal{O}(\alpha_s^3)$ corrections which are available only for the width.

The coefficients in Eq. (1) depend on three unphysical scales: the scale of the $\overline{\text{MS}}$ strong coupling constant μ_b , that of the $\overline{\text{MS}}$

https://doi.org/10.1016/j.physletb.2021.136679

^{0370-2693/© 2021} The Author(s). Published by Elsevier B.V. This is an open access article under the CC BY license (http://creativecommons.org/licenses/by/4.0/). Funded by SCOAP3.



Fig. 1. Scale dependence of Γ_{sl} at fixed values of the inputs and $\mu^{kin} = 1$ GeV. Dashed (solid) lines represent the two (three) loop calculation. In the left plot (μ_b -dependence) the blue (red) curves are at $\mu_c = 3(2)$ GeV; in the right plot (μ_c -dependence) the blue (red) curves $\mu_b = m_b^{kin} (m_b^{kin}/2)$.

charm mass μ_c , and the Wilsonian cutoff μ employed in the kinetic scheme definition of the *b* mass and of the OPE matrix elements. We choose the \overline{MS} scheme for the charm mass because all high-precision determinations of this mass are expressed in this scheme and we prefer to escape uncertainties related to the scheme conversion; in the following we choose $1.6 \,\text{GeV} \lesssim \mu_c \lesssim$ 3GeV, avoiding scales which are either too low or too high to provide a good convergence of the perturbative series. The kinetic scheme [3,18,19] provides a short-distance, renormalon-free definition of m_b and of the OPE parameters by introducing a hard cutoff μ to factor out the infrared contributions from the perturbative calculation. The cutoff μ should ideally satisfy $\Lambda_{OCD} \ll \mu \ll m_h$; in the following we will vary it in the range 0.7–1.3 GeV. Finally, the scale of the strong coupling constant will be varied in the range 2-8 GeV. Table I of Ref. [4] shows the size of the various coefficients in Eq. (1) for a couple of typical scale-settings. The third order coefficient a_3 is new and stems from the calculation of Ref. [1]. We reproduce the numerical results of Ref. [1] for the coefficients a_i .

As a first step in our analysis, we employ the results of the default fit of Ref. [4] with $\mu_b = m_b^{kin}$, $\mu_c = 3 \,\text{GeV}$, and $\mu = 1 \,\text{GeV}$, to extract $|V_{cb}|$ from Eq. (1). To this end, we employ the total semileptonic branching fraction obtained in the same fit and $\tau_B = 1.579(5)$ ps [23]. Notice that the 2014 default fit included a constraint on \overline{m}_c (3 GeV), but not on m_b^{kin} . Small shifts have to be applied to the values of m_{h}^{kin} , μ_{π}^{2} and ρ_{D}^{3} extracted from the fit in order to account for missing two-loop charm mass effects in the kinetic scheme definition adopted in the 2014 fit. These effects have now been computed in Ref. [3], where it was found that they reduce to decoupling effects and that they can be taken into account by expressing the kinetic scheme definitions in terms of $\alpha_s^{(3)}$. For $\mu_b = m_b^{kin}$ the shifts amount to +4 MeV, -0.003 GeV^2 , -0.002 GeV^3 , respectively. We adopt the PDG value for $\alpha_s^{(5)}(M_Z) = 0.1179(10)$, from which we get $\alpha_s^{(4)}(4.557 \,\text{GeV}) =$ 0.2182(36). We employ RunDec [24] to compute the running of all relevant scale-dependent quantities. We eventually find $|V_{cb}| =$ $42.49(44)_{th}(33)_{exp}$ 10⁻³, where the uncertainty refers to the inputs only and is split into an experimental and a theoretical component. If we neglect the new three-loop result we recover the same $|V_{cb}|$ central value as in [4]. The three loop correction therefore shifts $|V_{cb}|$ by +0.6%, well within the theoretical uncertainty of 1.3% estimated in [4]. The perturbative series is

$$\Gamma_{sl} = \Gamma_0 f(\rho) \Big[0.9255 - 0.1162_{\alpha_s} - 0.0350_{\alpha_s^2} - 0.0097_{\alpha_s^3} \Big]$$

= 0.5401 \Gamma_0, (2)

where the first term differs from 1 because of the power corrections. We can also repeat the same exercise evolving the value of $\overline{m}_c(3 \text{ GeV}) = 0.987(13) \text{ GeV}$ from the fit to $\mu_c = 2 \text{ GeV}$, which gives $\overline{m}_c(2 \text{ GeV}) = 1.091(14) \text{ GeV}$ and extract again $|V_{cb}|$ using $\mu_c = 2 \text{ GeV}$ in Eq. (1). We get $|V_{cb}| = 42.59(44)_{th}(33)_{exp} \ 10^{-3}$ and

$$\Gamma_{sl} = \Gamma_0 f(\rho) \Big[0.9258 - 0.0878_{\alpha_s} - 0.0179_{\alpha_s^2} - 0.0005_{\alpha_s^3} \Big] = 0.5374 \,\Gamma_0. \tag{3}$$

As noted in Ref. [1], the better convergence of the perturbative expansion with $\mu_c = 2 \text{ GeV}$, already observed in Ref. [25], carries on at the three loop level, but the cancellations appear somewhat accidental. Since the physical scale of the decay is actually lower than m_b , we believe a more appropriate choice for the scale of $\alpha_s^{(4)}$ is $\mu_b = m_b^{kin}/2$, which with $\mu_c = 2 \text{ GeV}$ leads to $|V_{cb}| = 42.59(44)_{th}(33)_{exp} \ 10^{-3}$ and

$$\Gamma_{sl} = \Gamma_0 f(\rho) \Big[0.9255 - 0.1140_{\alpha_s} - 0.0011_{\alpha_s^2} + 0.0103_{\alpha_s^3} \Big] = 0.5381 \,\Gamma_0. \tag{4}$$

We see from Eqs. (2)-(4) that the typical size of the three-loop corrections is 1% and that the perturbative series converge well at different values of the scales. A conservative estimate of the residual perturbative uncertainty on Γ_{sl} is therefore 0.5%, but it is worth studying the scale dependence of the width in more detail. In Fig. 1 we show the μ_b and μ_c dependence of Eq. (1) at two and three loops, using the inputs of the 2014 default fit. The scale dependence is reduced by the inclusion of the three loop contribution by over a factor 2, and the red curves appear to be flatter than the blue ones. The region of minimal scale dependence is the one around $\mu_b \sim \mu_c \approx 2 - 3$ GeV. We also studied the dependence of the width on the kinetic scale in the range $0.7 < \mu < 1.3$ GeV for different values of $\mu_{b,c}$, finding similar results. Defining $\Delta_{\mu_{b,c}}$ the maximum percentage deviation of the red solid lines in Fig. 1 and Δ_{μ} accordingly, with $\mu_b = m_b^{kin}/2$ and $\mu_c = 2$ GeV, we get

$$\Delta_{\mu_b} = 0.44\%, \qquad \Delta_{\mu_c} = 0.44\%, \qquad \Delta_{\mu} = 0.67\%.$$
(5)

Based on all this, we conservatively estimate a residual perturbative uncertainty of 0.7% in Γ_{sl} and consequently of 0.35% in $|V_{cb}|$ for our new default scenario, corresponding to $\mu = 1 \text{ GeV}$, $\mu_c = 2 \text{ GeV}$ and $\mu_b = m_b^{kin}/2 \simeq 2.3 \text{ GeV}$.

Beside the purely perturbative contributions, there are various other sources of uncertainty in the calculation of the semileptonic width [26], but the work done in the last few years has been fruitful. After the $\mathcal{O}(\alpha_s/m_b^2)$ corrections [27,28], the $\mathcal{O}(\alpha_s\rho_D^3/m_b^3)$

Table 1

Results of the updated fit in our default scenario ($\mu_c = 2 \text{ GeV}, \mu_b = m_b^{kin}/2$). All parameters are in GeV at the appropriate power and all, except m_c , in the kinetic scheme at $\mu = 1 \text{ GeV}$. The first and second rows give central values and uncertainties, the correlation matrix follows

m_b^{kin}	$\overline{m}_c(2{\rm GeV})$	μ_{π}^2	$ ho_D^3$	$\mu_G^2(m_b)$	$ ho_{LS}^3$	$BR_{c\ell\nu}$	$10^3 V_{cb} $
4.573 0.012	1.092 0.008	0.477 0.056	0.185 0.031	0.306 0.050	-0.130 0.092	10.66 0.15	42.16 0.51
1	0.307 1	-0.141 0.018 1	0.047 -0.010 0.735 1	0.612 -0.162 -0.054 -0.157	-0.196 0.048 0.067 -0.149	-0.064 0.028 0.172 0.091	-0.420 0.061 0.429 0.299
				1	0.001	0.013 0.033 1	-0.225 -0.005 0.684 1

corrections to Γ_{sl} have been recently computed in Ref. [21] (the $\mathcal{O}(\alpha_s \rho_{LS}^3)$ corrections to Γ_{sl} follow from the $\mathcal{O}(\alpha_s \mu_G^2/m_b^2)$ and are tiny). They are expressed in terms of m_b in the on-shell scheme and of $\overline{m}_c(m_b)$. After converting their result to the kinetic scheme and changing the scale of \overline{m}_c , we find that this new correction, together with all the terms of the same order generated by the change of scheme, enhances the coefficient of ρ_D^3 by 8 to 18%, depending on the various scales. However, after the conversion to the kinetic scheme the $\mathcal{O}(\alpha_s \rho_D^3)$ terms generate new $\mathcal{O}(\mu^3 \alpha_s^2)$ and $\mathcal{O}(\mu^3 \alpha_s^3)$ contributions that tend to compensate their effect. The resulting final shift on $|V_{cb}|$ is +0.05% with $\mu_c = 3 \text{ GeV}, \mu_b = m_b^{kin}$ and +0.1% for $\mu_c = 2 \text{ GeV}, \mu_b = m_b^{kin}/2$, and we choose to neglect it in the following.

After the calculation of the $\mathcal{O}(\alpha_{s}\rho_{D}^{3})$ contribution, the main residual uncertainty in Γ_{sl} is related to higher power corrections. The Wilson coefficients of the $O(1/m_h^4, 1/m^5)$ contributions have been computed at the tree level [29] – here the $O(1/m^5)$ effects include $O(1/m_b^3 m_c^2)$, sometimes referred to as Intrinsic Charm - but little is known about the corresponding 27 matrix elements. The Lowest Lying State Approximation (LLSA) [29] has been employed to estimate them and to guide the extension [5] of Ref. [4] to $O(1/m_h^4, 1/m^5)$. In the LLSA, the $O(1/m_h^4, 1/m^5)$ contributions increase the width by about 1%, but there is an important interplay with the semileptonic fit: as shown in Ref. [5], the $\mathcal{O}(1/m_h^4, 1/m^5)$ corrections to the moments and their uncertainties modify the results of the fit in a subtle way and the final change in Γ_{sl} is about +0.5%, a result stable under changes of the LLSA assumptions [5]. We therefore expect the $\mathcal{O}(1/m_h^4, 1/m^5)$ corrections to decrease $|V_{ch}|$ by 0.25% with respect to the default fit. Although the uncertainty attached to this value is mostly included in the theoretical uncertainty of the 2014 fit results, we may consider an additional 0.3% uncertainty for the width. Further uncertainties stem from unknown $\mathcal{O}(\alpha_s^2/m_b^2)$, and $\mathcal{O}(\alpha_s^2\rho_D^3/m_b^3)$ corrections, but they are all likely to be at or below the 0.1% level. The so-called Intrinsic Charm contributions, related to soft charm, lead to the $\mathcal{O}(1/m_b^3 m_c^2)$ corrections mentioned above, but also to terms of $\mathcal{O}(\alpha_s/m_b^3m_c)$ which may contribute up to 0.5% to the width [30]. Finally, one expects quark-hadron duality to break down at some point. Combining all the discussed sources of uncertainty, we estimate the total remaining uncertainty in Γ_{sl} to be 1.2%.

In the end, using the inputs of the 2014 default fit and setting $\mu_c = 2 \text{ GeV}, \ \mu_b = m_b^{kin}/2$ for the central value, we obtain

$$|V_{cb}|_{2014} = 42.49(44)_{th}(33)_{exp}(25)_{\Gamma} \, 10^{-3} = 42.48(60) \, 10^{-3} \quad (6)$$

where the uncertainty due to Γ_{sl} has been reduced by a factor 2 with respect to Ref. [4].

3. Updating the semileptonic fit

Despite ongoing analyses of the q^2 and M_X -moments at Belle and Belle II [31,32], no new experimental result on the semileptonic moments has been published since the 2014 fit [4]. On the other hand, new lattice determinations of m_b and m_c have been presented, improving their precision by roughly a factor 2. We use the FLAG 2019 averages [17] with $N_f = 2 + 1 + 1$ for m_b and m_c ,

$$\overline{m}_c(3 \text{ GeV}) = 0.988(7) \text{ GeV},$$

$$\overline{m}_b(\overline{m}_b) = 4.198(12) \text{ GeV},$$
(7)

which correspond to $\overline{m}_c(2 \text{ GeV}) = 1.093(8)$ and $m_b^{kin}(1 \text{ GeV}) = 4.565(19) \text{ GeV}$, where for the latter we have used option B of [3] for the definition of m_b^{kin} . We now repeat the 2014 default fit with both these constraints, slightly updating the theoretical uncertainty estimates. In view of the small impact of the $\mathcal{O}(1/m_b^4, 1/m^5)$ and $\mathcal{O}(\alpha_s \rho_D^3)$ corrections discussed in the previous section, we reduce the theoretical uncertainties used in the fit to the moments with respect to Ref. [4]. In particular, we consider a 20%, instead of a 30%, shift in ρ_D^3 and ρ_{LS}^3 , and reduce to 4 MeV the safety shift in $m_{c,b}$. For all of the other settings and for the selection of experimental data we follow Ref. [4].

While the central values of the fit are close to those of 2014, the uncertainty on m_b^{kin} ($\overline{m}_c(3 \text{ GeV})$) decreases from 20(12) to 12(7)MeV, and we get $|V_{cb}| = 42.39(32)_{th}(32)_{exp}(25)_{\Gamma} \ 10^{-3}$ with $\chi^2_{min}/dof = 0.46$. The very same fit performed with $\mu_c = 2 \text{ GeV}$ and $\mu_b = m_b^{kin}/2$ gives

$$|V_{cb}| = 42.16(30)_{th}(32)_{exp}(25)_{\Gamma} \ 10^{-3}$$
(8)

with $\chi^2_{min}/dof = 0.47$ and we neglect the very small shift due to the $\mathcal{O}(\alpha_s \rho_D^3)$ correction to Γ_{sl} . This is our new reference value and in Table 1 we display the complete results of this fit.

Let us now comment on the interplay between the fit to the moments and the use of Eq. (1). First, we observe that the fit to the moments is based on an $\mathcal{O}(\alpha_s^2)$ calculation [20,33–36] without $\mathcal{O}(\alpha_s \rho_D^3)$ contributions, and that the lower precision in the calculation of the moments with respect to the width inevitably affects the determination of $|V_{cb}|$. This is clearly visible in Eq. (6), where the theoretical component of the error is larger than the residual theory error associated with the width. However, only a small part of that uncertainty is related to the purely perturbative corrections, which are relatively suppressed in some semileptonic moments but sizeable in Γ_{sl} , as we have seen above. In other words, an $\mathcal{O}(\alpha_s^3)$ calculation of the moments is unlikely to improve the precision of the fit significantly, and the inclusion of $\mathcal{O}(\alpha_s^3)$ corrections only in Γ_{sl} is perfectly justified. On the other hand, an $\mathcal{O}(\alpha_s/m_b^3)$ calculation of the moments can have an important impact on the $|V_{cb}|$

determination. This is because the semileptonic moments, and the hadronic central moments in particular, are highly sensitive to the OPE parameters. Since the power correction related to ρ_D^3 amounts to about 3% percent in Eq. (1), an $\mathcal{O}(\alpha_s)$ shift on ρ_D^3 induced by perturbative corrections to the moments can have a significant impact in the determination of $|V_{cb}|$. Our estimates of the theoretical uncertainties take this into account. We also note that a fit without theoretical errors is a very poor fit $(\chi^2/dof \sim 2)$ with $|V_{cb}|$ decreased by slightly less than 1σ .

An important problem of the semileptonic fit is the sensitivity to the ansatz employed for the correlation among the theoretical uncertainties associated with the various observables [25]. We have studied the dependence of the result of Eq. (8) on the modelling of the theoretical correlations following Ref. [25] closely. Since the results shown above have been obtained using scenario **D** from Ref. [25] with $\Delta = 0.25 \text{ GeV}$, we have repeated the fit with option **B**, with option **C** using various values of ξ , and with option **D** for Δ in the range 0.15 - 0.30 GeV. The central values for $|V_{cb}|$ vary between 42.05 10^{-3} and 42.28 10^{-3} . These results are very much in line with Fig. 1 and Table 3 of Ref. [25] and therefore we do not add any uncertainty related to the theoretical correlations in Eq. (8).

We have also performed a fit including $\mathcal{O}(1/m_b^4, 1/m^5)$ corrections, in analogy with Ref. [5], to check the consistency with our main result of Eq. (8). We assign an error to the LLSA predictions and assume Gaussian priors for all of the 27 dimension 7 and 8 matrix elements. The error is chosen as the maximum of either 60% of the parameters value in the LLSA or $\Lambda_{LL}^n/2$ (n = 4, 5), with $\Lambda_{LL} = 0.55$ GeV, see Ref. [5] for additional details. As already noticed in Ref. [5], higher power corrections tend to decrease the value of $|V_{cb}|$. A fit performed with the same theory errors of Ref. [5] and $\mu_c = 2$ GeV and $\mu_b = m_b^{kin}/2$ leads to $|V_{cb}| = 42.00(32)_{th}(32)_{exp}(25)_{\Gamma} \ 10^{-3} = 42.00(53) \ 10^{-3}$, which is consistent with Eq. (8). Following the discussion above, one could slightly reduce the theory uncertainties in this fit with the only consequence of a small reduction on the error of $|V_{cb}|$.

Finally, repeating the reference fit of Table 1 without a constraint on m_b we obtain an independent determination of $m_b^{kin}(1 \text{ GeV}) = 4.579(16) \text{ GeV}$, which translates into $\overline{m}_b(\overline{m}_b) = 4.210(22)$ GeV. This determination, which still relies on the lattice determination of m_c reported in (7), is compatible with the FLAG $N_f = 2 + 1 + 1$ average for $\overline{m}_b(\overline{m}_b)$ and competitive with other current determinations of m_b .

4. Discussion

From a theoretical point of view, the reliability of the determination of $|V_{cb}|$ from inclusive semileptonic decays depends on our control of higher order effects. The new three-loop calculation of Ref. [1] shows that higher order perturbative effects are under control, and that they lie within the previously estimated uncertainties. This progress, together with the work done on higher power corrections [5,29] and on perturbative corrections to the Wilson coefficients of power suppressed operators [21,27,28], led us to estimate a residual theoretical error on Γ_{sl} of about 1.2%, and to slightly reduce the theoretical uncertainty in the fit to the moments.

Our final result is shown in Eq. (8). It is very close to previous determinations of $|V_{cb}|$ [4,5], but the total uncertainty is now 1.2%, one third smaller than in [4]. This reduction of the uncertainty reflects a better control of higher order effects, but it is also due to improved determinations of the heavy quark masses. The dominant single component of the uncertainty in Eq. (8) is now related to the experimental determination of the moments and of the semileptonic branching fraction, which are expected to be improved at Belle II. Future experimental analyses should also consider new observables beyond the traditional moments of the lepton energy and hadronic invariant mass distributions. For instance, the forward-backward asymmetry [37] and the moments of the leptonic invariant mass (q^2) distribution would enhance the sensitivity to the OPE matrix elements and reduce the uncertainty on $|V_{ch}|$. Because of reparametrisation invariance, the q^2 -moments and Γ_{sl} depend on a reduced number of OPE matrix elements [38], so that a fit to the q^2 -moments at $\mathcal{O}(1/m_b^4)$ involves only 8 parameters. This nice property allows for an independent check of the treatment of higher power corrections adopted in [5], but it is unlikely to lead to a competitive determination of $|V_{cb}|$. The q^2 moments will also constrain the soft charm effects considered in [30]. As far as the current experimental analyses are concerned, there are various aspects that require closer scrutiny. We refer in particular to the subtraction of QED corrections made with PHO-TOS [39], to the impact of Coulomb interactions, to the contribution of D^{**} states and to the correlations which play a crucial role in the fit.

Finally, turning to ways in which theory can improve the inclusive determination of $|V_{cb}|$, we have already argued that the most important missing higher order effects are probably the $\mathcal{O}(\alpha_s/m_b^3)$ corrections to the moments. Lattice QCD calculations provide precise constraints on the heavy guark masses, see Eq. (7), which are going to improve in the future, but we now have methods to compute differential distributions and their moments directly on the lattice [40]. While it is still unclear whether a determination of Γ_{sl} competitive with Eq. (1) can be achieved at the physical *b* mass, these lattice calculations might be able to enhance the predictive power of the OPE by accessing quantities which are inaccurate or beyond the reach of current experiments and are highly sensitive to the non-perturbative parameters. The computation of meson masses at different quark mass values [16,41] can also provide useful information when the data are analysed in the heavy guark expansion. At the moment, however, the reduction of the uncertainty in Eq. (8) exacerbates the V_{cb} puzzle, and calls for renewed efforts to solve an unwelcome anomaly, impervious to New Physics explanations [42,43].

Declaration of competing interest

The authors declare that they have no known competing financial interests or personal relationships that could have appeared to influence the work reported in this paper.

Acknowledgements

We are grateful to Matteo Fael for useful communications concerning the results of Refs. [1–3]. This work is supported by the Italian Ministry of Research (MIUR) under grant PRIN 20172LNEEZ.

References

- M. Fael, K. Schönwald, M. Steinhauser, Phys. Rev. D 104 (2021) 016003, arXiv: 2011.13654 [hep-ph].
- [2] M. Fael, K. Schönwald, M. Steinhauser, Phys. Rev. Lett. 125 (2020) 052003, arXiv:2005.06487 [hep-ph].
- [3] M. Fael, K. Schönwald, M. Steinhauser, Phys. Rev. D 103 (2021) 014005, arXiv: 2011.11655 [hep-ph].
- [4] A. Alberti, P. Gambino, K.J. Healey, S. Nandi, Phys. Rev. Lett. 114 (2015) 061802, arXiv:1411.6560 [hep-ph].
- [5] P. Gambino, K.J. Healey, S. Turczyk, Phys. Lett. B 763 (2016) 60, arXiv:1606. 06174 [hep-ph].
- [6] P. Gambino, M. Jung, S. Schacht, Phys. Lett. B 795 (2019) 386, arXiv:1905.08209 [hep-ph].
- [7] M. Bordone, M. Jung, D. van Dyk, Eur. Phys. J. C 80 (2020) 74, arXiv:1908.09398 [hep-ph].
- [8] P. Gambino, et al., Eur. Phys. J. C 80 (2020) 966, arXiv:2006.07287 [hep-ph].
- [9] S. Jaiswal, S. Nandi, S.K. Patra, J. High Energy Phys. 06 (2020) 165, arXiv:2002. 05726 [hep-ph].

- [10] A. Bazavov, et al., Fermilab Lattice, MILC, arXiv:2105.14019 [hep-lat], 2021.
- [11] J. Harrison, C.T.H. Davies, LATTICE-HPQCD, arXiv:2105.11433 [hep-lat], 2021.
- [12] T. Kaneko, et al., JLQCD, PoS LATTICE2019 (2019) 139, arXiv:1912.11770 [heplat].
- [13] G. Martinelli, S. Simula, L. Vittorio, arXiv:2105.08674 [hep-ph], 2021.
- [14] T. Kaneko, JLQCD, talk at FPCP 2021.
- [15] M. Czakon, A. Czarnecki, M. Dowling, arXiv:2104.05804 [hep-ph], 2021.
- [16] A. Bazavov, et al., Fermilab Lattice, MILC, TUMQCD, Phys. Rev. D 98 (2018) 054517, arXiv:1802.04248 [hep-lat].
- [17] S. Aoki, et al., Flavour Lattice Averaging Group, Eur. Phys. J. C 80 (2020) 113, arXiv:1902.08191 [hep-lat].
- [18] I.I.Y. Bigi, M.A. Shifman, N. Uraltsev, A.I. Vainshtein, Phys. Rev. D 56 (1997) 4017, arXiv:hep-ph/9704245.
- [19] A. Czarnecki, K. Melnikov, N. Uraltsev, Phys. Rev. Lett. 80 (1998) 3189, arXiv: hep-ph/9708372.
- [20] P. Gambino, J. High Energy Phys. 09 (2011) 055, arXiv:1107.3100 [hep-ph].
- [21] T. Mannel, A.A. Pivovarov, Phys. Rev. D 100 (2019) 093001, arXiv:1907.09187 [hep-ph].
- [22] P. Gambino, Int. J. Mod. Phys. A 30 (2015) 1543002, arXiv:1501.00314 [hep-ph].
- [23] Y.S. Amhis, et al., HFLAV, Eur. Phys. J. C 81 (2021) 226, arXiv:1909.12524 [hepex].
- [24] K.G. Chetyrkin, J.H. Kuhn, M. Steinhauser, Comput. Phys. Commun. 133 (2000) 43, arXiv:hep-ph/0004189.
- [25] P. Gambino, C. Schwanda, Phys. Rev. D 89 (2014) 014022, arXiv:1307.4551 [hep-ph].
- [26] D. Benson, I.I. Bigi, T. Mannel, N. Uraltsev, Nucl. Phys. B 665 (2003) 367, arXiv: hep-ph/0302262.
- [27] A. Alberti, P. Gambino, S. Nandi, J. High Energy Phys. 01 (2014) 147, arXiv: 1311.7381 [hep-ph].

- [28] T. Mannel, A.A. Pivovarov, D. Rosenthal, Phys. Rev. D 92 (2015) 054025, arXiv: 1506.08167 [hep-ph].
- [29] T. Mannel, S. Turczyk, N. Uraltsev, J. High Energy Phys. 11 (2010) 109, arXiv: 1009.4622 [hep-ph].
- [30] I.I. Bigi, N. Uraltsev, R. Zwicky, Eur. Phys. J. C 50 (2007) 539, arXiv:hep-ph/ 0511158.
- [31] F. Abudinén, et al., Belle-II, arXiv:2009.04493 [hep-ex], 2020.
- [32] R. van Tonder, Belle, in: 55th Rencontres de Moriond on Electroweak Interactions and Unified Theories, 2021, arXiv:2105.08001 [hep-ex].
- [33] K. Melnikov, Phys. Lett. B 666 (2008) 336, arXiv:0803.0951 [hep-ph].
- [34] A. Pak, A. Czarnecki, Phys. Rev. Lett. 100 (2008) 241807, arXiv:0803.0960 [hepph].
- [35] A. Pak, A. Czarnecki, Phys. Rev. D 78 (2008) 114015, arXiv:0808.3509 [hep-ph].
- [36] S. Biswas, K. Melnikov, J. High Energy Phys. 02 (2010) 089, arXiv:0911.4142 [hep-ph].
- [37] S. Turczyk, J. High Energy Phys. 04 (2016) 131, arXiv:1602.02678 [hep-ph].
- [38] M. Fael, T. Mannel, K. Keri Vos, J. High Energy Phys. 02 (2019) 177, arXiv:1812. 07472 [hep-ph].
- [39] P. Golonka, Z. Was, Eur. Phys. J. C 45 (2006) 97, arXiv:hep-ph/0506026.
- [40] P. Gambino, S. Hashimoto, Phys. Rev. Lett. 125 (2020) 032001, arXiv:2005.13730 [hep-lat].
- [41] P. Gambino, A. Melis, S. Simula, Phys. Rev. D 96 (2017) 014511, arXiv:1704. 06105 [hep-lat].
- [42] M. Jung, D.M. Straub, J. High Energy Phys. 01 (2019) 009, arXiv:1801.01112 [hep-ph].
- [43] A. Crivellin, S. Pokorski, Phys. Rev. Lett. 114 (2015) 011802, arXiv:1407.1320 [hep-ph].